

Royal Holloway University of London

Doctoral Thesis

**Algorithms and applications of geodynamic
modelling of rapid extension processes in SE Asia**

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Declaration

I hereby declare that except where specific reference is made to the work of others, the contents of this dissertation are original and have not been submitted in whole or in part for consideration for any other degree or qualification in this, or any other university. This dissertation is my own work and contains nothing which is the outcome of work done in collaboration with others, except as specified in the text and Acknowledgements.

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September 2018

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Abstract

The incompressible Boussinesq approximation have become a widely accepted and standardised strategy by the geodynamics community to solve the Stokes equations that describe the thermo-mechanical behaviour of Earth's interior. It is usually reasonable to assume that rocks comprising the lithosphere and mantle are (nearly) incompressible if one is focused on studying processes occurring at the uppermost part of the Earth. However, this hypothesis does not hold if attempt to compute deep mantle calculations or volumetric phase changes at lithospheric depths such as serpentinisation or melt extraction. In this thesis, I focus on the mathematical description, numerical implementation and benchmark of a fully compressible formulation of the Stokes equations so that volumetric strain can be accounted for, when necessary. Furthermore, volumetric increase rising from brittle failure is nearly always neglected. Subsequently, I have developed a visco-elasto-plastic constitutive law using an associated Drucker-Prager flow law for geodynamical processes.

Finally, I combined this tool with geochronological and geothermobarometric data to investigate the exhumation of metamorphic core complexes in Indonesia. South East Asia covers roughly the 15% of the Earth's surface and represents one of the most tectonically active regions in our planet, yet its tectonic evolution remains relatively poorly studied and understood in comparison with other heavily studied regions of the Earth. Recent episodes of extension in SE Asia have been associated with subduction initiation, sedimentary basin growth and phases of crustal melting, uplift and extremely rapid exhumation of young (< 5 Ma) metamorphic core complexes. In this Ph.D. I applied numerical tools to better comprehend some of these recent events that occurred (and many of them still ongoing) in SE Asia. Therefore, numerical models are used to better constrain the thermal conditions of the lithosphere and extension rates at which core complexes might have developed and rapidly exhumed in SE Asia. In particular, I compare available geothermobarometric data obtained from samples of the Palu Metamorphic Complex (PMC) with synthetic p-T paths computed from the numerical models.

*"I am not feeling the green burning flame,
as I gaze back along footprints you have made.
I am not dreaming of more than you have shown.
You're not a foundation, you are not a stone.
But I'm afraid of the way that I'm feeling;
afraid of this new understanding now.
Afraid for the beauty within me,
and that which I hold within my hand.
And this is the ultimate secret,
that many before me have ever known.
So capture me while I am weakest;
I want to know, I want to know.
Here I am wide open, surrendering to your side.
I have laid down my armour, I have no sword at my side.
I leave behind me the ruins of the fortress I swore to defend;
I leave behind me foundations;
I'll leave you a man I'll need you to mend.
And through all the battles around me,
I never believed I would fight.
Yet here I stand, a broken soldier, shivering and naked,
in your winter light."*

Footprints - Warning

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1 | Introduction

1.1 Plate tectonics

The dynamics of the Earth are governed by plate tectonics; in other words, the Earth is divided in a set of rigid plates whose relative motion with respect to each other is accommodated at the plate boundaries. The basics of this theory were drawn by Morgan [1968], following previous pioneering studies by Taylor [1910]; Jeffreys [1924]; Holmes [1931]; Du Toit [1937]; Wegener [1946]; Hess [1962], amongst others.

The plates are typically comprised by approximately 100 km of cool rocks, and they are classified as either oceanic or continental lithosphere. The largest fraction of the Earth's outer shell is made up of oceanic lithosphere, which is created at oceanic ridges (*divergent* plate boundaries), where two plates move away from each other. As the plates are pulled apart, hot and buoyant mantle rocks are up-welled towards the surface to fill the gap left by the divergent motion. These mantle rocks are cooled by conductive heating as they are exhumed at the spreading centre, and new oceanic lithosphere is created. As the new lithosphere moves away from the oceanic ridge, it further cools and thickens. Estimation of the composition of newly accreted oceanic lithosphere is an extremely challenging task from economical and technological points of view because it is covered by several kilometres of sea water. However, it is possible to study the composition of slices of oceanic lithosphere that have been uplifted and brought to the surface. These rock formations are known as *ophiolites* and are found in, for example, Oman, Cyprus, New Guinea, New Zealand and Newfoundland. Their study has revealed that oceanic lithosphere is composed by a thin layer (4-8 km), known as *oceanic crust*, of basaltic and gabbroic rocks, usually covered by sediments (from a few hundred metres to a few kilometres), and underlain by mantle rocks, typically referred as *peridotites*. The oceanic crust and the cold mantle rocks are separated by the so-called Mohorovičić or *Moho* discontinuity. Since the volume of lithospheric material has to be more or less constant over time, oceanic lithosphere is cycled back to the deep mantle at subduction zones (*convergent* plate boundaries); thus the age of ocean floor available at the Earth's surface is relatively young in geological time-scales, and very rarely exceeds 200 Ma (Fig. 1.1).

On the other hand, the composition of continental plates is mainly silicic, and the main hypothesis is that a vast portion of them was created during the early stages of the Earth's history by large amounts of melt extraction. New material is accreted into the continental lithosphere by mechanisms such as intraplate volcanism or accretion at subduction zones. Since the continental material is less dense than the oceanic lithosphere, it is gravitationally stable and cannot be cycled back to the mantle by subduction processes. Mechanisms that can consume continental lithosphere and bring it back to the deep mantle are delamination of its base, and subduction of continental sediments at subduction zones.

While the composition of the uppermost section of the continental crust is accessible and relatively easy to estimate, it remains challenging to estimate the composition of the lower crust. Nonetheless, direct evidence of its composition can be obtained from mineralogical and geochemical analyses of xenolith

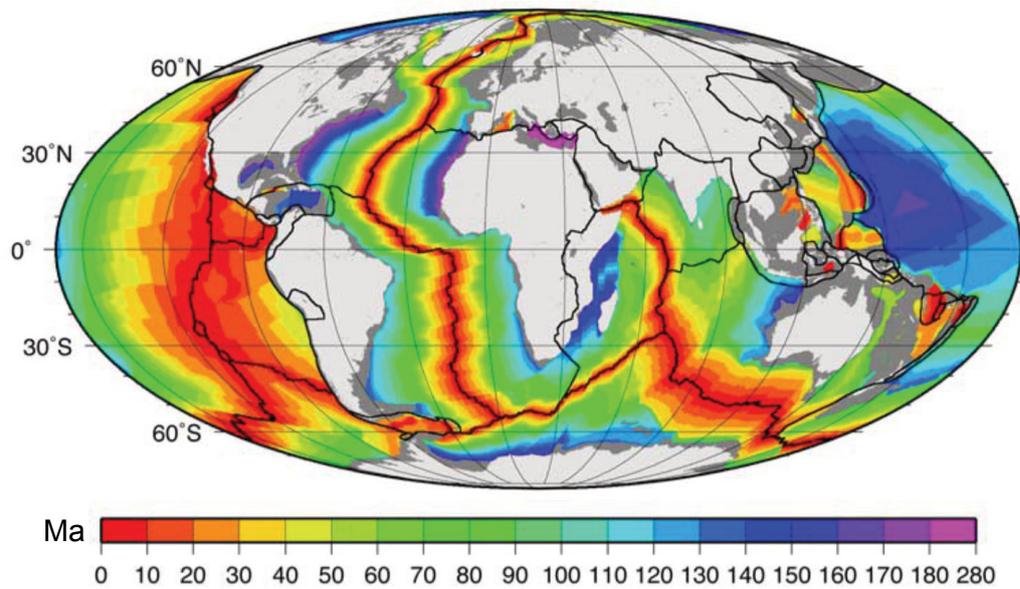


Fig. 1.1 Oceanic floor age. From Müller et al. [2008].

samples brought to the surface at continental volcanic centres, and from exhumed high-grade metamorphic rocks. The latter can be found in, for example, the so-called Metamorphic Core Complexes (MCCs). These are dome-shaped structures comprised of metamorphosed and -frequently- partially molten mid-to-lower crust that are exhumed along a low-angle normal fault (Whitney et al. [2013] and references therein). Despite being described first in continental plates [Coney, 1974; Lister and Davis, 1989], MCCs have also been identified in oceanic plates along oceanic ridges [e.g. Cann et al., 1997; Ranero and Reston, 1999] (Fig. 1.2). Another important fraction of information of the lower crust composition is inferred from indirect measurements, as for example, the speed of elastic waves travelling through the crust [e.g. Miller and Christensen, 1994; Sobolev and Babeyko, 1994; Kern et al., 1996; Musacchio et al., 1997].

1.2 Computer modelling of the Earth's dynamics

Direct observation of many geological processes, including the ones mentioned above, is often not possible or, in the best case, it is extremely difficult and expensive. Fortunately, highly valuable information of these processes can be inferred from indirect observations derived from geophysical and numerical studies. The rapid technological development produced in the last decades has made possible the proliferation of numerical models aiming at studying the dynamics of the Earth's, from the microscale to the macroscale. In spite of being simplified (quite often oversimplified) representations of thermal, mechanical and/or chemical processes, numerical models allow us to test a wide range of hypothesis at an almost insignificant economical cost, while still providing a large volume of invaluable information.

The motion and heat transfer of Earth's interior is described by the Stokes equation and the equation of conservation of energy, which comprise a set of Partial Derivative Equations (PDEs). Analytical solutions of PDEs usually exist only under specific geometry and boundary conditions, which are not representative of the bulk of geologic processes. Hence, these equations must be approximated using different numerical techniques such as the Finite Element Method (FEM) or Finite Differences (FE), which are two of the most widespread methods to tackle geodynamic problems.

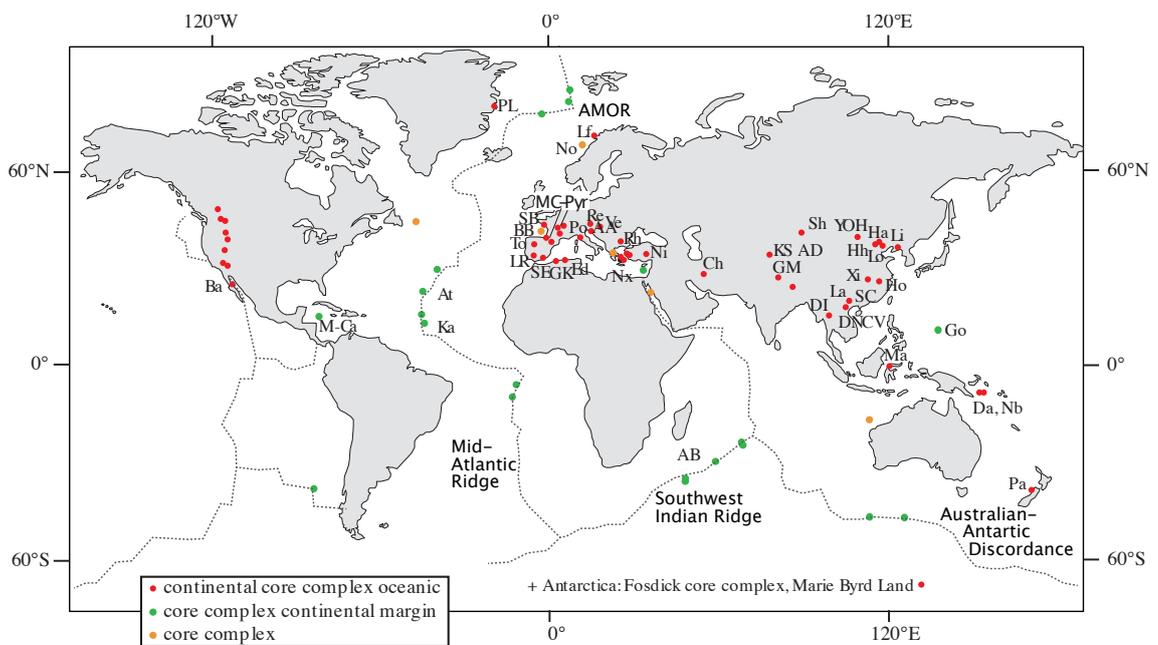


Fig. 1.2 Map of the world showing the locations of some Phanerozoic core complexes in the continents and oceans. Key to abbreviations: AA-Alpi Apuane (Italy); AB-Atlantis Bank (SW Indian Ridge); AD-Ama Drime (Nepal); AMOR-Arctic segment of Mid-Atlantic Ridge; At-Atlantis Massif (Mid-Atlantic Ridge); Ba-Baja (Mexico); BB-Bay of Biscay; Ch-Chapedony (Iran); Da-Dayman (Papua New Guinea); DI-Doi Inthanon (Thailand); DNCV-Day Nui Con Voi (Vietnam); Ed-Edough (Algeria); GK-Grand Kabilye (Algeria); GM-Gurla Mandhata (Pamirs); Go-Godzilla; Ha-Harkin (China/Mongolia); Hh-Hohhot (China); Ho-Hongzhen (China); Ka-Kane (Mid-Atlantic Ridge); KS-Kongur Shan (Pamirs); La-Laojunshan (China); Lf-Lofoten (Norway); Li-Liaodong Peninsula (China); Ma-Malino (Indonesia); Lo-Louzdian (China); LR-Lora del Rio (Spain); M-Ca-Mid-Cayman spreading center; MC-Pyr-Massif Central (France-Pyrenees, France, Spain; includes Montagne-Noire); Nb-Normanby Island (Papua New Guinea); Ni-Niğde (Turkey); No-Norway rifted continental margin; Nx-Naxos (Greece); Pa-Paparoa (New Zealand); PL-Payer Land (Greenland); Po-Pohorje Mountains (Slovenia); Re-Rechnitz (Austria); Rh-Rhodope (Greece, Bulgaria); SB-southern Brittany (France); SC-Song Chay (China); Sh-Shaerdelan (China); SE-Sierra de las Estancias (Spain); To-Tormes (Spain); Ve-Veporic (Slovenia); Xi-Xiaoqinling (China); YOH-Yagan-Onch-Hayrhan (China/Mongolia). From Whitney et al. [2013].

Even though the propagation of seismic waves has proven the solid state of the lithosphere and the mantle, these rheological layers behave as a viscous flow at extremely low velocities, as inferred from Global Isostasy Adjustment (GIA) studies [e.g. Peltier, 1996; Forte and Mitrovica, 1996; Mitrovica and Forte, 2004; van der Wal et al., 2015]. However, unlike hot and ductile mantle rocks, cold lithospheric material is able to release and dissipate stored elastic energy upon brittle failure, which manifests as heat release, faults, and earthquakes. This bimodal behaviour is frequently described by the mathematical model of a Maxwell body with a visco-elastic rheology, where the total strain is assumed to be the sum of its elastic and inelastic components. The mechanical response of the rocks is then defined by the so-called *constitutive equation*, which states the relationship between stress and strain for a given material. In a visco-elastic material, the dynamic viscosity is the physical parameters that controls the viscous response to applied forces, whereas the shear modulus controls the response of the elastic strain. Furthermore, the mantle and lithosphere are often treated as a *non-Newtonian* fluid, meaning that the viscosity strongly depends on temperature, compositional changes and applied stress; however, the stress dependency is sometimes omitted for computational reasons (frequently in 3D computations), and then the fluid is known as *Newtonian*. On the other hand, the shear modulus is usually considered a constant parameter that depends on the material. Nonetheless, the 'bulk' shear modulus of a rock is not strictly constant, and it will degrade if microscopic and macroscopic fractures propagate.

Several hypothesis are often considered in order to simplify the Stokes equations to make them more computationally stable and effective. First, the inertial forces are not considered. While this hypothesis holds for most tectonic processes due to extremely slow motions, it breaks if one wishes to study the propagation of elastic waves (i.e. earthquakes). Another common hypothesis is to assume that the mantle and lithosphere can be treated as (near-)incompressible bodies. This simplifies the problem, as the density becomes independent of the pressure and reduces the non-linear degree of the equations; incompressibility is incorporated by adopting the Boussinesq approximation. This approximation is suitable for many of the tectonic processes involving mantle-lithosphere deformation [e.g. Gerya and Yuen, 2007; Huismans and Beaumont, 2007; Buitter et al., 2006; Rey et al., 2009b; Huet et al., 2011; Schenker et al., 2012; Brune et al., 2012; Taramón et al., 2015; Tetreault and Buitter, 2017; Ros et al., 2017]; however, the Boussinesq approximation becomes inaccurate if density variations are larger than 10%. Deep mantle calculations are an example where the incompressibility approximation is no longer valid as the density changes exceed the accuracy threshold (density is roughly 60% higher at the core-mantle boundary) due to the enormous pressures at which rocks are subject. Moreover, a high compressibility will affect the viscous dissipation and adiabatic heating, thus swaying the thermal structure.

Compressibility was first introduced by the so-called *anelastic approximation* [Jarvis and McKenzie, 1980] to study deep mantle convection. This approach is based on the hypothesis that the dynamic pressure is very small compared to the lithostatic pressure and this allows the density to change as a function of depth, but not with time. Compressibility has been further studied [e.g. Glatzmaier, 1988; Schmeling, 1989; Bercovici et al., 1992; Tackley, 1996; King et al., 2010] in mantle convection calculations either using the anelastic approximation or its variants, such as the *truncated anelastic approximation*. However, if volumetric strain linked to phase changes is considered, the density rate might become non-negligible and the latter hypothesis breaks. Anyhow, the incompressible Boussinesq approximation prevails as the preferred hypothesis employed to reduce the degree of complexity of the Stokes equations.

1.3 Aims

The aims of this Ph.D thesis are to develop a numerical tool able to handle a general formulation in 2D of the compressible Stokes equations for a viscous flow, in order to study the effects of geological processes where volume changes and compressibility should not be overlooked. For completeness and consistency, the constitutive law is expanded so that volume changes raising from plastic deformation are also accounted for.

In parallel to the development of the new numerical tools mentioned above, numerical models are used to help to unravel the tectonic history behind the rapid exhumation and high cooling rates observed in the Palu Metamorphic Complex (PMC), one of the youngest MCCs on Earth (< 5 Ma), located in Sulawesi, Indonesia (Fig. 1.3). The PMC has been chosen as the case of study due to the possibility of comparing the numerical results with newly available geothermobarometric from rock samples coming from the PMC [Hennig et al., 2017].

1.4 Thesis outline

This thesis is divided in 7 chapters. Chapter 2 describes the methodology used in this thesis, which includes a brief description of the Finite Element Method, followed by the description and numerical implementation of the equations governing the thermo-mechanical evolution of tectonic processes. The core of this work is

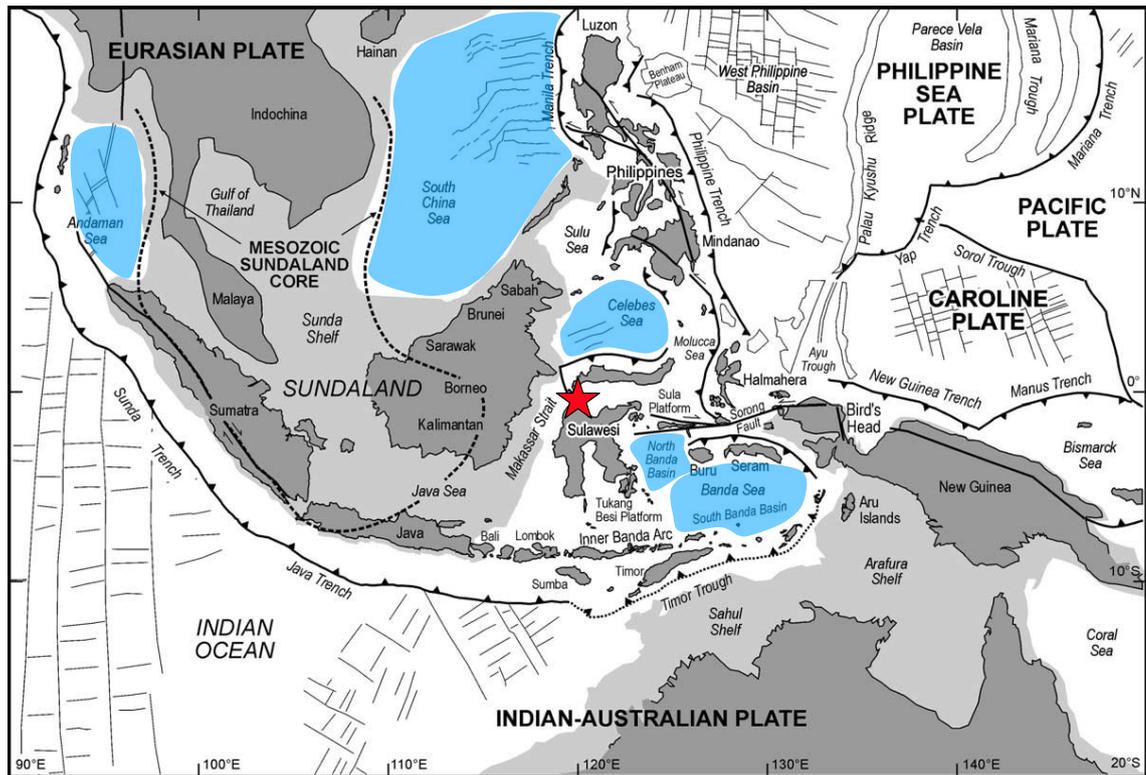


Fig. 1.3 Geographical map showing the main geological features in SE Asia. Regions where sea-floor spreading occurred are delimited by the shaded blue areas and the red star marks the location of the Palu Metamorphic Complex. Modified after Hall [2002].

composed by three journal publications or manuscript drafts to be submitted into scientific journals in the near future.

Large amount of the Ph.D. focuses on the development and implementation of a general compressible formulation of the Stokes equations that accounts for volumetric strain due pressures changes and also to phase changes, such as serpentinisation and melt extraction. The description of compressible formulation and the numerical strategies to solve the resulting set of PDEs are presented in Chapter 3. A set of numerical experiments designed to validate its correct implementation is also included. This chapter consists of a manuscript draft for submission in *Geochemistry*, *Geophysics*, *Geosystems*.

While Chapter 3 focuses on solving the compressible equations governing the motion of a visco-elastic Maxwell body, the numerical model is extended in Chapter 4 by introducing a visco-elasto-plastic constitutive equation able to handle the fully non-associated (non-dilatant) and associated (dilatant) plastic limits. First, the formulation of the new constitutive law is described. Second, the results of shear band initiation obtained with this formulation are compared with other published numerical and analytical studies. And at last, the implications of plastic dilation are further extended to an example of large scale tectonic process; in this case I consider rifting of continental crust with different crustal strengths. This chapter consists of a manuscript draft for later publication in *Tectonophysics*.

In Chapter 5, numerical modelling is applied to a geological case. The Palu Metamorphic Complex (Sulawesi, Indonesia) developed and exhumed in very recent geological times ($< c. 5$ Ma) at rates much higher than previously known for other continental metamorphic complexes. In this chapter I investigate the conditions at which MCCs are exhumed so rapidly. I do so by comparing synthetic cooling paths computed by numerical

models against published geochronological and thermobarometric data from Hennig et al. [2017]. This chapter consists of a manuscript currently under review in *Earth and Planetary Science Letters*.

In Chapter 6 I discuss the methods and results illustrated in this thesis, and the final conclusions are drawn in Chapter 7.

2 | Numerical methods

Due to the large time and length scales of geological process, deformation of Earth's interior is often approximated by the Stokes equations that describe the thermo-mechanical behaviour of a viscous flow. The solvability of this set of partial derivative equations (PDEs) is not trivial, and requires of complex numerical techniques. All the PDEs found in this thesis are solved using the *Finite Element Method* (FEM), and implemented in the MATLAB-based code LaCoDe for bidimensional problems. FEM is a powerful and highly versatile numerical tool to approximate the continuum solution of PDEs with complex geometries and boundary conditions, and it is widely employed in engineering and other fields. The FEM is based on subdividing the domain where any given PDE needs to be solved into a set of smaller discrete regions, the so-called *elements*. The equations are first defined locally in the elements, and then they are assembled into a global system of equations where the contributions of each element are accounted for. Other numerical methods are commonly used to approximate numerically the PDEs that describe different physical processes, such as *Finite Differences Method* (FDM), *Finite Volumes Method* (FVM), *Extended Finite Element Method* (X-FEM) or *Discrete Elements Method* (DEM), but are not be considered nor discussed in this project.

In this chapter I present a description and applicability of the Lagrangian FEM to solve the Stokes equations for a visco-elasto-plastic body and other physical processes implemented in LaCoDe. A brief description of the FEM is presented in Appendix A.1, and the reader is referred to Zienkiewicz [1985]; Hughes [1987]; Bathe [2006] for a more in-depth description of the FEM.

2.1 Stokes equations: incompressible Boussinesq approximation

The general Stokes equations are defined by the coupling of the equations of conservation of momentum, conservation of energy and conservation of mass. In the following sections I will describe these equations and its numerical implementation using the FEM. Large part of this project focuses on developing a code to solve the motion of compressible visco-elastic flows. The mathematical description, numerical implementation and benchmarking of such formulation is given in Chapter 3. However, some of the models presented in this thesis (i.e. Chapter 5) employ the incompressible Boussinesq approximation. This approximation states that the flow is (nearly) incompressible and density variations only intervene in the buoyancy forces. Under this assumption, small density changes due to, for example, thermal expansion and contraction are permitted. However, the Boussinesq approximation becomes inaccurate if these density changes are $\Delta\rho/\rho > 0.1$.

In this section I present the standard mathematical formulation of an viscous flow employing the Boussinesq incompressible approximation, and the numerical implementation using the FEM. For further details in the mathematical description and numerical solutions, the reader is referred to e.g. Hughes [1987]; Donea and Huerta [2003]; Zienkiewicz and Taylor [2005]. First, the equation of conservation of energy that determines

the diffusion and advection of the temperature field is described. Then I proceed to describe the equation of conservation of momentum and conservation of mass.

2.1.1 Thermal diffusion

The time dependent equation of heat advection and diffusion with and external source of heat (or heat consumption) for an incompressible material is described by the strong form the equation of conservation of energy:

$$\rho C_p \frac{\partial T}{\partial t} + \mathbf{u} \nabla T = \kappa \nabla^2 T + Q \quad (2.1)$$

or casting out for the x and z axis:

$$\rho C_p \frac{\partial T}{\partial t} + u_x T_{,x} + u_z T_{,z} = \kappa (T_{,xx} + T_{,zz}) + Q \quad (2.2)$$

where ρ is the density, C_p is the heat capacity, T is temperature, κ is thermal conductivity and $\nabla = e_i \partial / \partial x_i$ is the nabla operator, where e_i is the standard basis. The source term Q can be positive (i.e. heat generated by, for example, radiogenic decay or inelastic work) or negative (e.g. latent heat cooling during melting). In the incompressible approximation of a viscous flow we consider two sources of heat that are included in Q : 1) heat produced by radiogenic decay and 2) shear heating, which is produced by inelastic (i.e. viscous and plastic) work and it is defined as $H_{sh} = \tau_{ij} \dot{\epsilon}_{ij}^{inel} = \tau_{ij} (\dot{\epsilon}_{ij}^{viscous} + \dot{\epsilon}_{ij}^{plastic})$. It must be noted that eq. (2.2) describes the advection and diffusion of heat under an Eulerian frame of reference. Under a Lagrangian frame of reference, the advection part is done by updating the nodal positions accordingly to the velocity field, and only the diffusion terms are solved numerically, thus the advection term $\mathbf{u} \nabla T$ vanishes and the partial time derivative is equivalent to the material time derivative (i.e. $\partial(\cdot) / \partial t = D(\cdot) / Dt$), see Section 2.2 for more details. The numerical Lagrangian formulation and implementation of eq. (2.2) is described in the following subsection.

Numerical implementation

The Lagrangian time-dependent diffusion equation in a domain Ω is defined by the following boundary problem:

$$\rho C_p \frac{DT}{Dt} = \kappa \nabla^2 T + Q \quad (2.3)$$

with the boundary conditions

$$T = g \quad \text{on } \Gamma_D \quad (2.4)$$

$$-\mathbf{n} \kappa T = q \quad \text{on } \Gamma_N \quad (2.5)$$

where n is the unit outward normal vector to the boundary Γ and q is the heat flux. The derivation of the weak formulations of the equation of heat diffusion can be found in Appendix A.2.1 and reads as follows:

$$\int_{\Omega} \mathbf{N} \rho C_p \frac{D(\mathbf{N} \tilde{T})}{Dt} d\Omega + \int_{\Omega} (\nabla \mathbf{N}^T) \kappa (\nabla \mathbf{N} \tilde{T}) d\Omega = \int_{\Omega} \mathbf{N} Q d\Omega - \int_{\Gamma_N} \mathbf{N} q d\Gamma_N \quad (2.6)$$

or in a compact matrix form (and dropping the wide tilde over T for more clarity):

$$\mathbf{M} \frac{DT}{Dt} + \mathbf{K} T = \mathbf{f} \quad (2.7)$$

where \mathbf{M} is the mass matrix and \mathbf{K} is the conductivity matrix:

$$\mathbf{M} = \int_{\Omega} \mathbf{N}^T \mathbf{N} d\Omega \quad (2.8)$$

$$\mathbf{K} = \int_{\Omega} \nabla^T \mathbf{N}^T \kappa \nabla \mathbf{N} d\Omega \quad (2.9)$$

$$\mathbf{f} = \int_{\Omega} \mathbf{N} Q d\Omega - \int_{\Gamma_N} \mathbf{N} q d\Gamma_N \quad (2.10)$$

The Galerkin method approximates the spatial dependency of the problem; however, eq. (2.7) is also time dependent equation as the mass matrix is multiplied by the time derivative of the temperature. In theory, it is possible to further use the FEM to perform the time discretisation; however, using a finite differences approach to compute the time derivatives is a common strategy due to its efficiency and simple implementation. The time derivatives are then approximated as follows:

$$\mathbf{M} \frac{DT^{n+\alpha}}{Dt} + \mathbf{K} T^{n+1} = \mathbf{f}^{n+1} \quad (2.11)$$

$$T^{n+1} = T^n + \Delta t \frac{DT^{n+\alpha}}{Dt} \quad (2.12)$$

$$\frac{DT^{n+\alpha}}{Dt} = (1 - \alpha) \frac{DT^n}{Dt} + \frac{DT^{n+1}}{Dt} \quad (2.13)$$

Subscript n indicates time at t_n and $n + 1$ indicates time at $t_n + \Delta t$, and Δt is the time step. Using this scheme, we obtain the following system of linear equations:

$$(\mathbf{M} + \alpha \Delta t \mathbf{K}) T^{n+1} = (\mathbf{M} - (1 - \alpha) \Delta t \mathbf{K}) T^n + \Delta t (\alpha \mathbf{f}^{n+1} + (1 - \alpha) \mathbf{f}^n) \quad (2.14)$$

or as a more condensed expression:

$$\mathbf{K}^* T = \mathbf{f}^* \quad (2.15)$$

This scheme is part of the generalized trapezoidal family of methods and a more detailed description can be found in Chapter 8 of Hughes [1987]. Depending on the value of α eq. (2.14) yields to different methods of the so-called trapezoidal family: if $\alpha = 0$ it describes the forward Euler method; $\alpha = 0.5$ describes the Crank-Nicolson method; and $\alpha = 1$ describes the backward Euler method. The backward Euler method is used in all the models discussed in this thesis, unless other method is specified, because the solution depends on T and DT/Dt at $t = t_{n+1}$ and is unconditionally stable. On the other hand, the forward Euler method usually requires very small time steps to yield accurate solutions.

2.1.2 Conservation of momentum and conservation of mass

The deformation of a incompressible viscous flow is described by the coupling of the equation conservation of momentum and conservation of mass:

$$\sigma_{i,j,j} = -\rho g_i \quad (2.16)$$

$$u_{i,i} = 0 \quad (2.17)$$

where σ_{ij} is the Cauchy stress tensor, and g_i is the gravitational acceleration. It should be noted that the acceleration in mantle-lithosphere processes is negligible, thus eq. (2.16) describes the conservation of momentum of an inertia-free system. It is convenient to split the Cauchy stress tensor into its deviatoric (shear component that deforms the fluid) and hydrostatic (pressure in equilibrium that does not disturb the

fluid) components:

$$\sigma_{ij} = \tau_{ij} - p\delta_{ij} \quad (2.18)$$

where τ_{ij} is the deviatoric stress tensor, δ_{ij} is the Kronecker delta, and the pressure p is the mean of the principal stresses:

$$p = -\frac{1}{3}\sigma_{kk} \quad (2.19)$$

Substituting eq. (2.16) into eq. (2.18), the conservation of momentum yields:

$$\tau_{ij,j} - p_{,j} = \rho g_i \quad (2.20)$$

The general relationship between the stress and the deformation is given by:

$$\sigma_{ij} = \mathcal{C}_{ijkl}\dot{\epsilon}_{ij} \quad (2.21)$$

where \mathcal{C}_{ijkl} is a 4th rank tensor that includes the material properties, and $\dot{\epsilon}_{ij}$ is the strain tensor, defined as:

$$\dot{\epsilon}_{ij} = \frac{1}{2}(u_{i,j} + u_{j,i}) \quad (2.22)$$

Due to the symmetry of its components, the constitutive law of a viscous flow in (2.21) is reduced to:

$$\sigma_{ij} = 2\eta_s \left(\dot{\epsilon}_{ij} - \frac{1}{3}\dot{\epsilon}_{kk}\delta_{ij} \right) + \eta_b \dot{\epsilon}_{kk}\delta_{ij} - p\delta_{ij} \quad (2.23)$$

where η_s is the shear viscosity, that accounts for the resistance of the material to shear strain rates and the bulk viscosity η_b determines the resistance of the fluid to compressive deformation. Therefore, the first term in eq. (2.23) describes the stresses caused by shear deformation, and the second term accounts for volumetric deformation due to normal stresses. We neglect the effects volumetric resulting from the bulk viscosity, thus the second term of the right-hand-side in eq. (2.23) vanishes:

$$\sigma_{ij} = \overbrace{2\eta\dot{\epsilon}'_{ij}}^{\tau_{ij}} - p\delta_{ij} \quad (2.24)$$

where $\dot{\epsilon}'_{ij} = \dot{\epsilon}_{ij} - \frac{1}{3}\dot{\epsilon}_{kk}\delta_{ij}$ is the deviatoric strain rate tensor. To simplify the notation, the subindex s of the shear viscosity will be dropped from now onwards, hence $\eta_s = \eta$. The incompressibility of the flow is mathematically described by a divergence-free velocity field:

$$\dot{\epsilon}_{ii} = 0 \quad (2.25)$$

This means that the inflow into an infinitesimal volume is equal to the outflow, and therefore, the volume is preserved. Eq. (2.25) will be referred as the *incompressible constraint*. Given the expressions described above, a Stokes flow boundary problem in n dimensions is defined by the following set of coupled equations in the domain Ω :

$$\tau_{ij,j} - p_{,j} = \rho g_i \quad (2.26)$$

$$u_{i,i} = 0 \quad (2.27)$$

and the boundary conditions

$$u_i = h_i \quad \text{on } \Gamma_D \quad (2.28)$$

$$\sigma_{ij}n_j = t_i \quad \text{on } \Gamma_N \quad (2.29)$$

Now we can substitute eq. (2.24) in eq. (2.26) and obtain the following expressions for the x axis:

$$2\eta_{eff} \left(\dot{\epsilon}_{xx} - \frac{1}{3} \dot{\epsilon}_{kk} \delta_{ij} \right)_{,x} + (2\eta_{eff} \dot{\epsilon}_{xz})_{,z} - p_{,x} = 0 \quad (2.30)$$

and for the z axis:

$$2\eta_{eff} \left(\dot{\epsilon}_{zz} - \frac{1}{3} \dot{\epsilon}_{kk} \delta_{ij} \right)_{,z} + (2\eta_{eff} \dot{\epsilon}_{zx})_{,x} - p_{,z} = -\rho g_z \quad (2.31)$$

Using the definition of the strain tensor we obtain the strong forms of the conservation of momentum:

$$\eta_{eff} \left(\frac{4}{3} u_{x,x} - \frac{2}{3} u_{z,z} \right)_{,x} + \eta_{eff} (u_{x,z} + u_{z,x})_{,z} - p_{,x} = 0 \quad (2.32)$$

for the x axis. And:

$$\eta_{eff} \left(\frac{4}{3} u_{z,z} - \frac{2}{3} u_{x,x} \right)_{,z} + \eta_{eff} (u_{x,z} + u_{z,x})_{,z} - p_{,z} = -\rho g_z \quad (2.33)$$

for the z axis. Eqs.(2.32) and (2.33) are then transformed into weak form and solved using the FEM. The reader is referred to Appendix (A.2.2) or Chapter 3 for the derivation of the Stokes equations for a incompressible and compressible visco-elastic flow, respectively.

2.1.3 Numerical implementation

The resulting set of governing equations of the problem is solved numerically using the FEM to generate a system of linear equations. The governing equations (2.16) and (2.17) are transformed into their weak forms with help of the trial solutions and weighting functions employing the Galerkin approximatim (see Appendix A.2.2), reading:

$$\int_{\Omega} \mathbf{B}^T \mathbf{D} \mathbf{B} \tilde{\mathbf{u}} d\Omega - \int_{\Omega} \mathbf{B}^T \mathbf{m} \mathbf{N}_p \tilde{p} d\Omega = \int_{\Omega} \mathbf{N}_u^T (\rho \mathbf{g} + \mathbf{B} \chi \hat{\boldsymbol{\tau}}) d\Omega \quad (2.34)$$

$$- \int_{\Omega} \mathbf{N}_p^T \mathbf{m}^T \mathbf{B} \tilde{\mathbf{u}} d\Omega = 0 \quad (2.35)$$

where the elemental matrix \mathbf{B}^e represents the strain-displacement matrix, and \mathbf{D}^e is the rheology matrix that relates strain rates to deviatoric stresses:

$$\mathbf{B}^e \tilde{\mathbf{u}}^e = \begin{bmatrix} \frac{\partial N_u}{\partial x} & 0 \\ 0 & \frac{\partial N_u}{\partial z} \\ \frac{\partial N_u}{\partial z} & \frac{\partial N_u}{\partial x} \end{bmatrix} \begin{bmatrix} u_x \\ u_z \end{bmatrix} = \begin{bmatrix} \dot{\epsilon}_{xx} \\ \dot{\epsilon}_{zz} \\ \dot{\epsilon}_{xz} \end{bmatrix} \quad (2.36)$$

$$\mathbf{D}^e = \eta_{eff} \begin{bmatrix} C_1 & C_2 & 0 \\ C_2 & C_1 & 0 \\ 0 & 0 & 1 \end{bmatrix} \quad (2.37)$$

and

$$\mathbf{m}^T = [1 \quad 1 \quad 0] \quad (2.38)$$

and the coefficients in D take the values $C_1 = 4/3$ and $C_2 = -2/3$. The system of equations (2.34) and (2.35) can be conveniently written in block matrix form as:

$$\begin{pmatrix} \mathbf{A} & \mathbf{Q} \\ \mathbf{Q}^T & \mathbf{0} \end{pmatrix} \cdot \begin{pmatrix} \tilde{\mathbf{u}} \\ \tilde{\mathbf{p}} \end{pmatrix} = \begin{pmatrix} \mathbf{f} \\ 0 \end{pmatrix} \quad (2.39)$$

Solution scheme

Numerical complications arise due to the presence of the diagonal zeros in the full matrix eq. (2.39). A widespread method to tackle this issue consists in using the following modified continuity equation:

$$\nabla \cdot \mathbf{u} = \frac{p}{\lambda} \quad (2.40)$$

where λ is a mesh- and problem-independent parameter, commonly referred as penalty parameter. It becomes evident that $p/\lambda \rightarrow 0$ for large values of λ , and incompressibility is recovered. This so-called *penalty method* has been widely used in incompressible flow problems and a detailed description can be found in, for example, Hughes [1987]; Donea and Huerta [2003] and references therein. Introducing the penalty term in the systems of equations (A.22) yields:

$$\begin{pmatrix} \mathbf{A} & \mathbf{Q} \\ \mathbf{Q}^T & -\frac{1}{\lambda}\mathbf{M} \end{pmatrix} \cdot \begin{pmatrix} \tilde{\mathbf{u}} \\ \tilde{\mathbf{p}} \end{pmatrix} = \begin{pmatrix} \mathbf{f} \\ 0 \end{pmatrix} \quad (2.41)$$

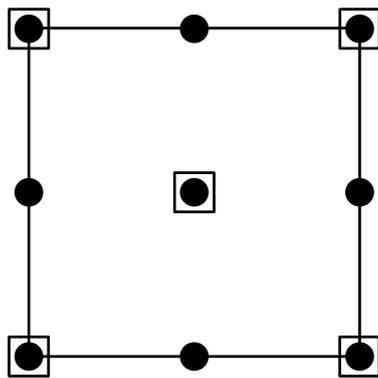
where \mathbf{M} is the so-called mass matrix:

$$\mathbf{M} = \int_{\Omega} \mathbf{N}_p^T \mathbf{N}_p d\Omega \quad (2.42)$$

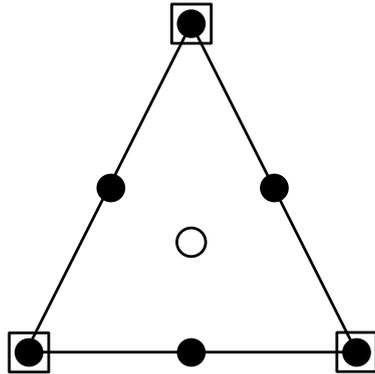
The choice of the penalty number λ is not a trivial decision. A low value of λ will not ensure near-incompressibility and it will introduce errors in the pressure field. On the other hand, λ has to be large enough so that it ensures the near-incompressibility of the flow, but the constrain term p/λ may dominate the system and result in a zero-velocity field if λ is excessively high. This phenomenon is often referred as *numerical locking* or *mesh locking*, and this occurs if there are too many incompressibility constrains compared to velocity unknowns. Locking is avoided by choosing elements that satisfy the so-called LBB (or inf-sup) condition. This condition states that not all the couples of the velocity and pressure functional spaces are stable. The mathematics behind the LBB condition are out of the scope of this thesis and more details can be found in Hughes [1987]; Donea and Huerta [2003], and references therein. For incompressible flow Zienkiewicz and Taylor [2005] recommends to use $\lambda = c\eta$, with $10^7 < c < 10^{10}$, for double precision computations.

There are different possible combinations of velocity-pressure functional spaces satisfying the LBB condition; some of the most popular choices for two-dimensional elements are shown in Fig. 2.1. LaCoDe uses Crouzeix-Raviart triangular elements, where the velocity field is approximated by seven nodal points and quadratic interpolation enhanced by a cubic bubble function in the baricenter of the element. Pressure is discontinuous with three nodal points and a linear interpolation [Crouzeix and Raviart, 1973].

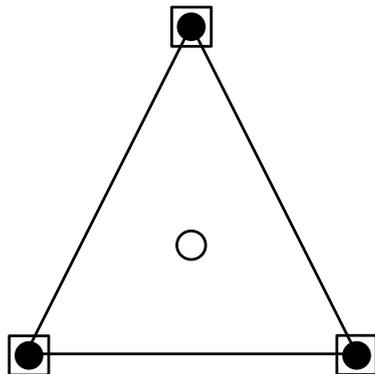
Locking can be alternatively avoided by under-integration of the penalty terms [Malkus and Hughes, 1978]. This method consists in using fewer integration points to calculate the numerical integrals that construct the element matrices than the number of integrations points required to exactly evaluate the integral. This results

**Q2Q1 (or Taylor-Hood) element:**

- Continuous biquadratic velocity
- Continuous bilinear pressure
- Quadratic convergence

**Crouzeix-Raviart element:**

- Continuous quadratic + cubic bubble function velocity
- Discontinuous linear pressure

**Mini element:**

- Continuous linear + cubic bubble function velocity
- Continuous linear pressure

□ Pressure node ● Velocity node ○ Bubble function node

Fig. 2.1 Examples of combinations of velocity and pressure spaces that satisfy the LBB condition [Donea and Huerta, 2003].

in a less accurate and lower order integral. By using a lower order integration scheme for the penalty term, the number of incompressibility constraints is effectively reduced.

High values of λ lead to a poorly conditioned stiffness matrix, which hinders its solvability with iterative schemes. Therefore, a lower value of λ is used in combination with a Powell-Hestenes iterative scheme [Powell, 1967; Hestenes, 1969]:

1. Choose $\tilde{\mathbf{p}}^0 = 0$.

2. Solve the velocity field:

$$\tilde{\mathbf{u}}^{k+1} = \left(\mathbf{A} + \lambda \mathbf{Q} \mathbf{M}^{-1} \mathbf{Q}^T - \mathbf{Q} \tilde{\mathbf{p}}^k \right)^{-1} \mathbf{f} \quad (2.43)$$

3. Calculate pressure correction:

$$\Delta \tilde{\mathbf{p}} = \lambda \mathbf{M}^{-1} \mathbf{Q}^T \tilde{\mathbf{u}}^{k+1} \quad (2.44)$$

4. Update pressure:

$$\tilde{\mathbf{p}}^{k+1} = \tilde{\mathbf{p}}^k + \Delta \tilde{\mathbf{p}} \quad (2.45)$$

5. Repeat steps 2-4 until $\Delta p < Tol$.

This algorithm enforces the incompressibility of the flow by correcting and updating the pressure field and the resulting forces. This solution scheme for the incompressible Stokes equations can also be understood as Augmented Lagrangian method, where the second row of the system of eqs. (2.41) is augmented by subtracting $\lambda^{-1} \mathbf{M} \tilde{\mathbf{p}}$ and adding the following iterative scheme:

$$\begin{pmatrix} \mathbf{A} & \mathbf{Q} \\ \mathbf{Q}^T & -\frac{1}{\lambda} \mathbf{M} \end{pmatrix} \cdot \begin{pmatrix} \tilde{\mathbf{u}} \\ \tilde{\mathbf{p}} \end{pmatrix}^{k+1} = \begin{pmatrix} \mathbf{g}_1 \\ -\frac{1}{\lambda} \mathbf{M} \tilde{\mathbf{p}}^k \end{pmatrix} \quad (2.46)$$

Upon convergence $\mathbf{M} \tilde{\mathbf{p}}^{k+1} = \mathbf{M} \tilde{\mathbf{p}}^k$ and the incompressibility constraint is satisfied.

2.1.4 Non-linear rheologies

The Stokes equations described in the previous section describe the thermo-mechanical behaviour of a linear visocous (or visco-elastic) flow. However, the deformation of tectonic events is better characterised by rheological laws such as diffusion creep, dislocation creep and/or plastic failure. These mechanisms depend on variables such as strain rate, temperature, pressure or differential stress, thus introducing different degrees of non-linearities into the Stokes equations. In this section, I present the mathematical description, implementation and solution scheme for these non-linear rheological laws. The plasticity presented in this chapter correspond to a non-dilatant formulation using the Prandtl-Reus flow law. The formulation of a dilatant plasticity is presented in Chapter 4.

Viscous creep

Two mechanisms for viscous deformation are included in the model: diffusion creep and dislocation creep [Poirier, 1985; Karato et al., 2001]. Diffusion creep occurs at low stress levels, when atoms diffuse through inside the crystal grains and along the grain boundaries, causing deformation of the rock. Deformation due to dislocation creep is caused by the migration of dislocations through the crystal lattice of the rock. Both creep mechanisms are strain rate-, temperature- and pressure- dependent:

$$\dot{\epsilon} = A \sigma_d^n \exp\left(\frac{E_a + pV_a}{nRT}\right) \quad (2.47)$$

where A is the pre-exponential parameter, E_a is the activation energy, V_a is the activation volume and R is the universal gas constant. The pre-exponential parameter $A = B d^{-m} f_{H_2O}^r \exp(\alpha \phi)$ takes in consideration the

Table 2.1 Mechanical parameters for dislocation and diffusion creep.

| Rheology | Creep | $\log_{10}(A)$ ($Pa^{-n}s^{-1}$) | n | E ($kJmol^{-1}$) | Reference |
|-----------------|-------------|------------------------------------|-----|--------------------|----------------------------|
| Wet quartzite | Dislocation | -28 | 4.0 | 223 | Gleason and Tullis [1995] |
| Mafic granulite | Dislocation | -21.05 | 4.2 | 445 | Gleason and Tullis [1995] |
| Dry olivine | Dislocation | -15.56 | 3.5 | 530 | Wilks and Carter [1990] |
| Wet olivine | Dislocation | -15.05 | 3.5 | 480 | Hirth and Kohlstedt [2003] |
| Dry olivine | Diffusion | -8.65 | 1.0 | 375 | Hirth and Kohlstedt [2003] |
| Wet olivine | Diffusion | -8.66 | 1.0 | 335 | Hirth and Kohlstedt [2003] |

grain size d , grain size exponent m , water fugacity f_{H_2O} , water fugacity exponent r , melt fraction factor α and melt fraction ϕ . Using the constitutive law of viscous flow:

$$\tau_{II} = 2\eta\dot{\epsilon}_{II} \quad (2.48)$$

the correspondent dislocation and diffusion viscosities can be computed substituting eq. (2.47) in (2.48) :

$$\eta_{dif} = \frac{1}{2}(A)^{-\frac{1}{n}}(\dot{\epsilon}_{II}^{dif})^{\frac{1}{n}-1} \exp\left(\frac{E_a + pV_a}{nRT}\right) \quad (2.49)$$

$$\eta_{dis} = \frac{1}{2}(A)^{-\frac{1}{n}}(\dot{\epsilon}_{II}^{dis})^{\frac{1}{n}-1} \exp\left(\frac{E_a + pV_a}{nRT}\right) \quad (2.50)$$

where n is the power-law exponent, the subindex II indicates the square root of the second invariant of an arbitrary tensor $a_{II} = \sqrt{\frac{1}{2}a_{ij}a_{ij}}$. The power-law exponent for diffusion creep is $n = 1$, thus yielding an expression for the diffusion viscosity that does depend on strain; on the other hand, theoretical values of the power-law exponent for dislocation are $n = 3 - 4$ and yields a non-linear equation. An effective creep viscosity is now built as:

$$\frac{1}{\eta} = \frac{1}{\eta_{dif}} + \frac{1}{\eta_{dis}} \quad (2.51)$$

In this way, the smallest viscosity will have the largest contribution to the effective viscosity, with deformation dominated by the mechanism that has the smallest activation stress. The viscous strain tensor is then $\dot{\epsilon}_{ij}^{visc} = \dot{\epsilon}_{ij}^{dif} + \dot{\epsilon}_{ij}^{dis}$ and, using the definitions (2.49) and (2.50), the diffusion and dislocation strain tensors are respectively computed as:

$$\dot{\epsilon}_{ij}^{dif} = \frac{\tau_{ij}}{2\eta_{dif}}; \quad \dot{\epsilon}_{ij}^{dis} = \frac{\tau_{ij}}{2\eta_{dis}} \quad (2.52)$$

Values of the parameters in eqs. (2.49) and (2.50) used in this thesis are summarised in Table 2.1. It must be noted that these values have been obtained in laboratory conditions from either uniaxial or triaxial experiments, meaning that the strain rate and differential stress at which they were calculated, is not equivalent to the square root of the second invariant of a given tensor (i.e. strain and deviatoric stress). The differential stress in these experiments is defined as:

$$\sigma_d = \sigma_1 - \sigma_3 \quad (2.53)$$

where σ_1 and σ_3 are the maximum and minimum eigenvalues (in other words, the maximum and minimum principal stresses) of the stress tensor and the conditions:

$$\sigma_1 \neq \sigma_2 = \sigma_3 \quad (2.54)$$

where 1 represents the vertical axis and 2 and 3 are the horizontal axis. For a uniaxial experiment $\sigma_2 = \sigma_3 = 0$, whereas σ_2 and σ_3 are the confining pressure for a triaxial experiment. The strain rate in these experiments is then the axial strain $\dot{\epsilon}_1$. Under these conditions, the deviatoric strain yields:

$$\tau = \begin{bmatrix} \sigma_1 - \frac{\sigma_1 - 2\sigma_3}{3} & 0 & 0 \\ 0 & \sigma_3 - \frac{\sigma_1 - 2\sigma_3}{3} & 0 \\ 0 & 0 & \sigma_3 - \frac{\sigma_1 - 2\sigma_3}{3} \end{bmatrix} \quad (2.55)$$

and the square root of its second invariant is:

$$\tau_{II} = \sqrt{\frac{1}{2} \left(\left[\frac{2}{3}\sigma_1 - \frac{2}{3}\sigma_3 \right]^2 + 2 \left[\frac{1}{3}\sigma_3 - \frac{1}{3}\sigma_1 \right]^2 \right)} \quad (2.56)$$

which simplifies to:

$$\tau_{II} = \frac{1}{\sqrt{3}} \sigma_d \quad (2.57)$$

If one assumes uniform deformation along the axis 2 and 3 and incompressibility:

$$-\frac{1}{2}\dot{\epsilon}_1 = \dot{\epsilon}_2 = \dot{\epsilon}_3 \quad (2.58)$$

and the square root of the second invariant of the strain rate tensor yields:

$$\dot{\epsilon}_{II} = \sqrt{\frac{1}{2} \left(\dot{\epsilon}_1^2 + 2 \left[-\frac{1}{2}\dot{\epsilon}_1 \right]^2 \right)} \quad (2.59)$$

which simplifies to

$$\dot{\epsilon}_1 = \frac{2}{\sqrt{3}} \dot{\epsilon}_{II} \quad (2.60)$$

The power law eq. (2.47) can be now generalized in terms of the square roots of the second invariants of stress and strain rate using eqs. (2.57) and (2.60):

$$\tau_{II} = C(A)^{\frac{1}{n}} (\dot{\epsilon}_{II})^{\frac{1-n}{n}} \exp\left(\frac{E + PV}{nRT}\right) \quad (2.61)$$

where C is a conversion factor:

$$C = (2)^{\frac{1}{n}} (3)^{-\frac{n+1}{2n}} \quad (2.62)$$

Plastic deformation

Materials undergo non-recoverable plastic deformation if the yield stress is exceeded. The stress at which a material fails is defined by the yield surface $\mathcal{F}(\tau_{ij}, q)$, a scalar function of the deviatoric stress and the softening parameter q , which limits the maximum stress possible within the material.

For points of the material where deformations are purely visco-elastic $\mathcal{F} < 0$, whereas $\mathcal{F} = 0$ at yield. If the stress field at any point of the domain is such that $\mathcal{F} > 0$, the stress must be brought back to the yield

surface. Plastic strain rate is defined by the plastic multiplier $\dot{\gamma} > 0$ and the plastic potential \mathcal{G} :

$$\dot{\epsilon}_{ij}^{(pl)} = \dot{\gamma} \frac{\partial \mathcal{G}}{\partial \tau_{ij}} \quad (2.63)$$

The addition of plastic strain rates to eq. (A.40) leads to the visco-elasto-plastic constitutive equation:

$$\dot{\epsilon}'_{ij} = \frac{\tau_{ij}}{2\eta} + \frac{1}{2G} \frac{D\tau_{ij}}{Dt} + \dot{\gamma} \frac{\partial \mathcal{G}}{\partial \tau_{ij}} \quad (2.64)$$

In this section we adopt the deviatoric, corner-free and non-associative ($\mathcal{F} \neq \mathcal{G}$) Prandtl-Reus flow rule (e.g. Zienkiewicz and Taylor [2005]). This flow rule takes the von Mises yield surface as the flow potential:

$$\mathcal{G} = \tau_{II} \quad (2.65)$$

and therefore

$$\frac{\partial \tau_{II}}{\partial \tau_{ij}} = \frac{\tau_{ij}}{2\tau_{II}} \quad (2.66)$$

Plastic volumetric strain can be included by using an associative flow rule ($\mathcal{F} = \mathcal{G}$). This topic is further described and discussed in Chapter 4. Using the Prandtl-Reus flow rule, and after some algebraic manipulations, we can obtain the following expression:

$$\tau_{ij} \left(\frac{1}{2\eta} + \frac{1}{2G\Delta t} + \frac{\dot{\gamma}}{2\tau_{II}} \right) = \dot{\epsilon}'_{ij} + \frac{\tau_{ij}^n}{2G\Delta t} + \frac{\tau_{ij}^{rot}}{2G} \quad (2.67)$$

where τ_{ij}^{rot} are the terms in eq. (A.41) associated to rigid body rotations. Assuming that at yield $\tau_{II} = \tau_y$:

$$\tau_{ij} = \eta_{vpl} \left(2\dot{\epsilon}'_{ij} + \frac{\tau_{ij}^n}{G\Delta t} + \frac{\tau_{ij}^{rot}}{G} \right) \quad (2.68)$$

where η_{vpl} is the effective visco-elasto-plastic viscosity given by:

$$\eta_{vpl} = \frac{\eta G \tau_y \Delta t}{G \tau_y \Delta t + \eta \tau_y + \eta \dot{\gamma} G \Delta t} \quad (2.69)$$

Or alternatively, the effective visco-elasto-plastic viscosity can be computed directly from eq. (2.68) as:

$$\eta_{vpl} = \frac{\tau_{ij}}{2\dot{\epsilon}'_{ij} + \frac{\tau_{ij}^n}{G\Delta t} + \frac{\tau_{ij}^{rot}}{G}} \quad (2.70)$$

In order to be consistent with our description of visco-elastic deformations, eq. (2.70) is rewritten in terms of η_{eff} and χ , giving the expression of the effective visco-elasto-plastic viscosity:

$$\eta_{vpl} = \frac{\tau_y}{2\dot{\epsilon}_{II} + \chi \widehat{\tau}_{II}} \quad (2.71)$$

Thus at yield, the constitutive visco-elasto-plastic constitutive law is given by

$$\tau_{ij} = 2\langle \eta \rangle \dot{\epsilon}_{ij} + \chi \widehat{\tau}_{ij} \quad (2.72)$$

and

$$\langle \eta \rangle = \begin{cases} \eta_{eff} & \text{for } \mathcal{F} \leq 0 \\ \eta_{vpl} & \text{for } \mathcal{F} > 0 \end{cases} \quad (2.73)$$

If one wishes to recover the plastic strain, $\dot{\gamma}$ can be computed as:

$$\dot{\gamma} = \tau_y \left(\frac{1}{\eta_{vpl}} - \frac{1}{\eta_{G\Delta t}} - 1 \right) \quad (2.74)$$

Plastic strain rate is then recovered after substitution of eq. (2.74) in (2.63):

$$\dot{\epsilon}_{ij}^{(pl)} = \frac{1}{2} \left(\tau_y \left[\frac{1}{\eta_{vpl}} - \frac{1}{\eta_{G\Delta t}} - 1 \right] \right) \frac{\tau_{ij}}{\tau_{II}} \quad (2.75)$$

Yield surface

The yield surface is a scalar function that defines the domain of admissible stresses, thus defining the yield stress. There is a vast range of proposed yield surfaces; however, in geodynamics it is common to use pressure-sensitive yield criterion. I therefore adopt the pressure-dependent Drucker-Prager [Drucker and Prager, 1952] yield criterium to describe the yield stress, for which \mathcal{F} is defined as:

$$\mathcal{F} = \tau_{II} - p \sin \phi - c \cos \phi \leq 0 \quad (2.76)$$

where ϕ is the friction angle and c is the cohesion. To avoid shear bands forming at 45° , we consider that the strength of the material depends on the total pressure [Kaus, 2010]. A common alternative to Drucker-Prager is the isotropic Mohr-Coulomb [Coulomb, 1773] yield surface:

$$\mathcal{F} = (\sigma_1 - \sigma_3) + (\sigma_1 + \sigma_3) \sin \phi - 2c \cos \phi \leq 0 \quad (2.77)$$

where σ_1 and σ_3 are the maximum and minimum principal stresses, respectively. The Drucker-Prager yield surface is represented by a cone in the space of principal stresses, whereas Moh-Coulomb is a hexagonal pyramid. This means that both functions have an apex aligned with the hydrostatic axis. The derivative at the apex is a singularity and should be computed very carefully. This issue can be bypassed by combining Drucker-Prager or Mohr-Coulomb for high stress/pressure values with the von Mises yield surface [Mises, 1913] for low stress/pressure values. Therefore, the following two-surface yield criterion is defined (Fig. 2.2):

$$\mathcal{F} = \begin{cases} \tau_{II} - p \sin \phi - c \cos \phi & \tau_{II} > c \\ \tau_{II} - c & \tau_{II} \leq c \end{cases} \quad (2.78)$$

for a mixed Drucker-Prager - von Mises yield surface, or:

$$\mathcal{F} = \begin{cases} \mathcal{F} = (\sigma_1 - \sigma_3) + (\sigma_1 + \sigma_3) \sin \phi - 2c \cos \phi & \tau_{II} > c \\ \tau_{II} - c & \tau_{II} \leq c \end{cases} \quad (2.79)$$

for a mixed Mohr Coulomb - von Mises yield surface. I must note that the Mohr-Coulomb yield surface is available in LaCode, however, it is not employed in the work presented in this thesis.

Strain softening

The accumulate plastic strain is commonly employing as the softening parameter [de Souza Neto et al., 2011]. Strain softening can be applied to both friction angle and cohesion; however, in this thesis, strain softening is only applied to the friction angle. The friction angle is then reduced as a linear function of the softening parameter, defined as competition between the plastic strain and a healing term [Moresi and

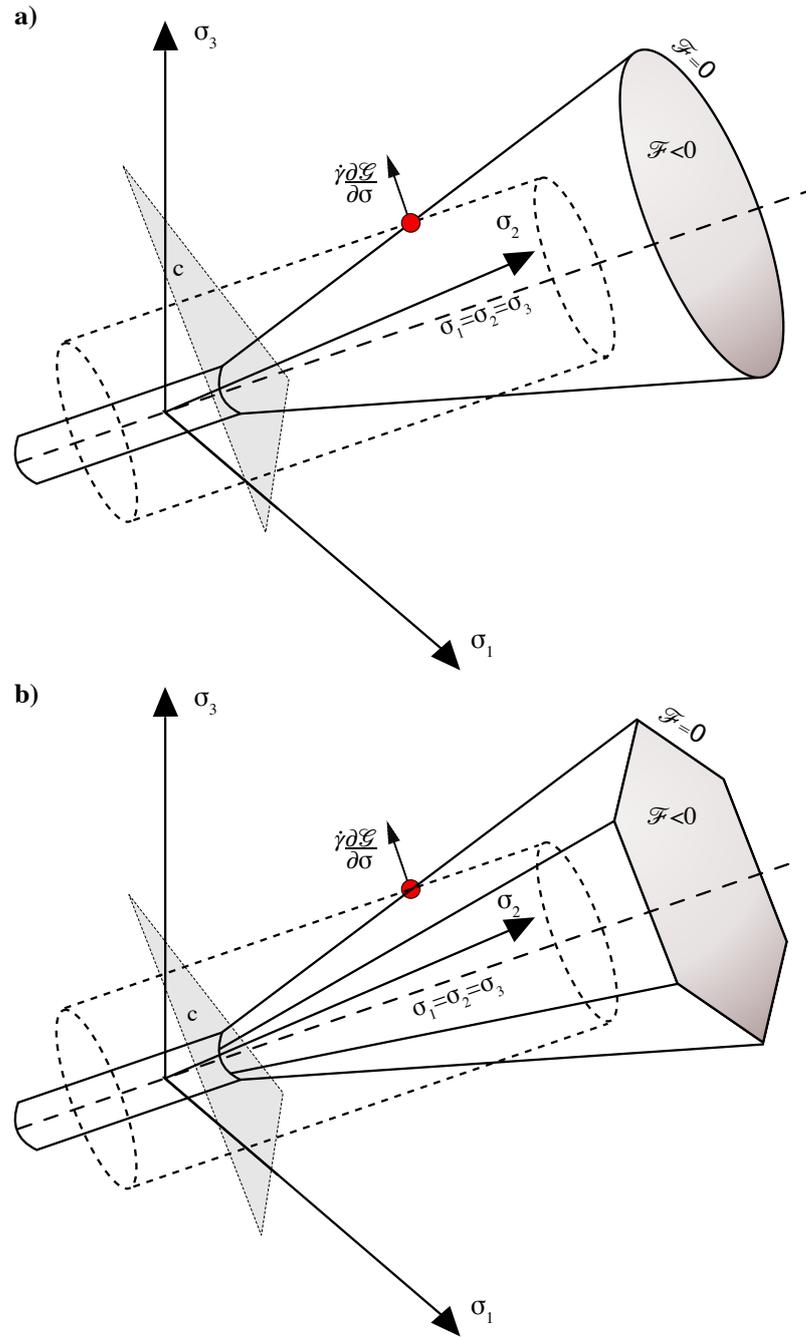


Fig. 2.2 Graphical representation of the mixed yield surface in the principal stress space: (a) Drucker-Prager and (b) Mohr-Coulomb yield surface define the strength envelope at high pressure/differential stress values, and Von Misses defines the strength envelope for stresses so that $\tau_{II} < c$. The dashed line represents the hydrostatic axis $\sigma_1 = \sigma_2 = \sigma_3$.

Mühlhaus, 2006]:

$$q = E_{ij}^{pl} = \int \left(\gamma \frac{\tau_{ij}}{2\tau_y} - q_h(\tau_{ij}, \eta) \right) dt \quad (2.80)$$

where the healing term $q_h(\tau, \eta)$ is a scalar function proportional to the background viscous strain:

$$q_h(\tau_{ij}, \eta) = \vartheta \frac{\tau_{II}}{\eta} \quad (2.81)$$

and $0 < \vartheta < 1$ is a scalar function of pressure and temperature. Unless specified, initial friction angles of $\phi_0 = 35^\circ$ and minimum angle of friction of $\phi_\infty = 15^\circ$ are used in the numerical experiments presented in

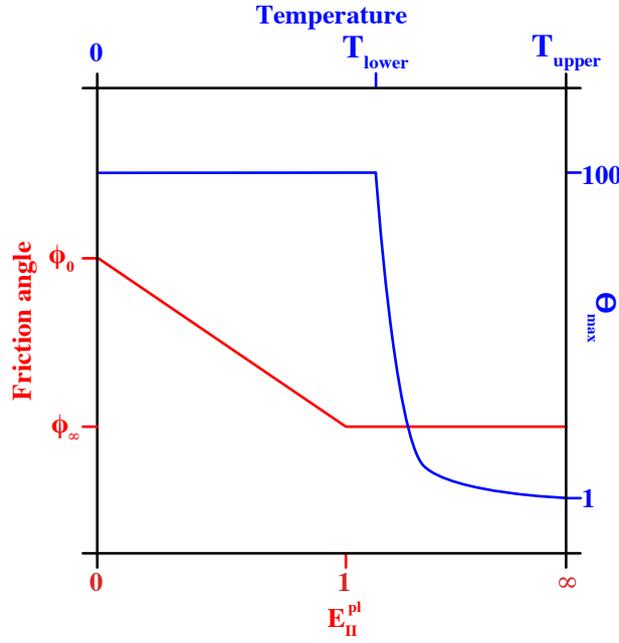


Fig. 2.3 In red: linear dependence of plastic softening on finite plastic strain. In blue: temperature dependency of the maximum viscous pre-exponential factor.

this thesis. The maximum amount of softening is reached at $E_{II} = 1$, thus no further softening is applied for larger deformations.

The factor Θ is introduced in the equations of diffusion and dislocation to weaken the viscous deformations due to grain size reduction by dislocation mechanism and crystallographic orientations [Karato and Wu, 1993; Hansen et al., 2012]:

$$\eta = \Theta \frac{1}{2} A^{(\frac{1}{n})} \dot{\epsilon}_{II}^{(\frac{1}{n}-1)} \exp\left(\frac{E + PV_a}{nRT}\right) \quad (2.82)$$

The pre-exponential Θ factor depends linearly on the second invariant of the strain tensor. The pre-exponential factor (Fig. 2.3) takes values of $1 \leq \Theta \leq \Theta_{max}$, with Θ_{max} being a scalar function of temperature:

$$\Theta_{max}(T) = \begin{cases} 100 & \text{if } T \leq T_{lower} \\ \exp\left(\frac{T - T_{upper}}{T_{lower}}\right) & \text{if } T_{lower} < T < T_{upper} \\ 1 & \text{if } T > T_{upper} \end{cases} \quad (2.83)$$

where T_{lower} and T_{upper} are the lower and upper limits that define the exponential decay, respectively.

2.1.5 Non-linear iterations

In geodynamical problems, rheological non-linearities are typically present (e.g. dislocation creep, plastic deformation, temperature-dependent density). These non-linearities are treated by nesting the linear solver within a set of Picard iterations. Non-linear iterations are terminated when the value of the residual of the velocity is below a certain tolerance. The residual is defined as:

$$R = \frac{\|u^i - u^{i-1}\|_{\infty}}{\|u^i\|_{\infty}} \leq Tol \quad (2.84)$$

where $\|\cdot\|_{\infty}$ is the infinity norm and i is the non-linear iteration index.

2.2 Advection scheme

Let us first define the material time derivative as:

$$\frac{D\psi}{Dt} = \frac{\partial\psi}{\partial t} + u \cdot \nabla\psi \quad (2.85)$$

where $\partial(\cdot)/\partial t$ indicates the partial time derivative of the function (\cdot) , and $\psi = \psi(x, y, z)$. As described previously in this chapter, LaCoDe solves the governing equations of a compressible and incompressible viscous flow using a Lagrangian frame of reference. Under this assumption, any fluid particle is followed by the observer as it moves through space and time. For this reason, the advection term $u \cdot \nabla\psi$ in eq. vanishes, and the material time derivative is equivalent to the partial time derivative. On the other hand, the observer is fixed and does not follow the flow particles as they move through time and space under a Eulerian frame of reference. In the Eulerian formulation of the Stokes equations for a visco-elastic flow, the advection term must be included in the equations of conservation of momentum, conservation of mass, and in the Jaumann derivative of the elastic stress.

In the context of the Lagrangian FEM, the elements are deformed, at the end of every time step, once the velocity and temperature fields are computed, by simply performing the following calculation for every node of the FEM mesh (Fig. 2.4a):

$$x_i = x_i + u_i dt \quad (2.86)$$

Since LaCoDe uses 7-nodes triangular elements, the edges of the elements may not remain completely straight and the inner node may not be located at the barycentre of the element, after the mesh is advected (Fig. 2.4b). This effect might lead to highly distorted elements after few time steps. To avoid an excessive use of remeshing techniques, the position of the nodes located at the centre of the edges is recalculated so that the edges are straight, and the 7th node inside the element is brought back to the barycentre (Fig. 2.4c).

2.3 Remeshing

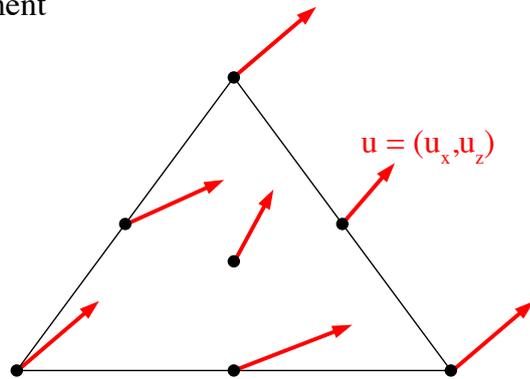
One of the drawbacks of using a Lagrangian formulation is that large deformation of the mesh eventually leads to highly distorted elements. This issue is overcome by mapping the necessary fields (i.e. temperature, density, accumulated strain) onto a newly generated high-quality mesh. To reduce the associated computational cost and interpolation errors, the new mesh is generated only when the quality of the mesh is below a given threshold. Therefore, the remeshing algorithm is called only if $q_n < .25$, $\alpha < 7^\circ$ and $\beta > 170^\circ$, for at least one element. Where α and β are the smallest and largest angles, respectively, and q_n is a quality factor defined as:

$$q_n = \frac{4\sqrt{3}A}{||ab||^2 + ||ac||^2 + ||bc||^2} \quad (2.87)$$

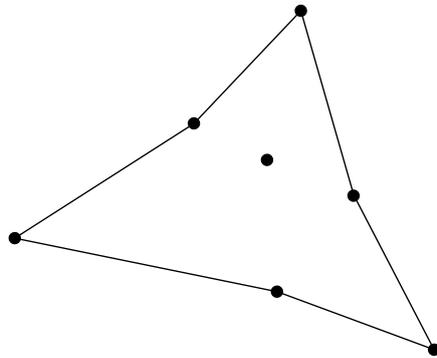
where A is the area, and a , b and c are the vertices of the triangle. The remeshing scheme and mapping of the fields onto the new mesh and the accuracy of the remeshing scheme are described in Chapter 3. The remeshing algorithm currently implemented in LaCoDe is optimised to work with the mesh generator Triangle [Shewchuk, 1996] for perfect body-fitting meshes, and with an adaptive mesh generator [Liu et al., 2018] for non-body-fitting meshes.

In the first case, the interface between two different bodies (e.g. the contact between two rheological phases) is tracked through time. When remeshing is necessary, the nodes at this interface are used as a boundary condition in order to generate the new mesh. Perfect body-fitting meshes are extremely convenient in the context of the FEM to model the mechanical behaviour of a composite body. However, this approach might

a) Original element

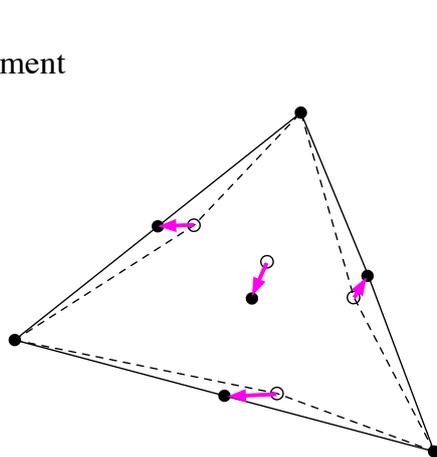


b) Advected element



advection:
 $x_i = x_i + u_i dt$

c) Corrected element



straighten edges and
 relocate inner node

Fig. 2.4 Advection scheme: (a) Undeformed 7-node Crouzeix-Raviart triangular element. The red arrows represent the velocity vectors. (b) Deformed element after applying the advection scheme. (c) The element is corrected by straightening its edges and relocating the central node back to its baricenter. The dashed triangle and empty circles represent the element and the nodes pre-correction, respectively.

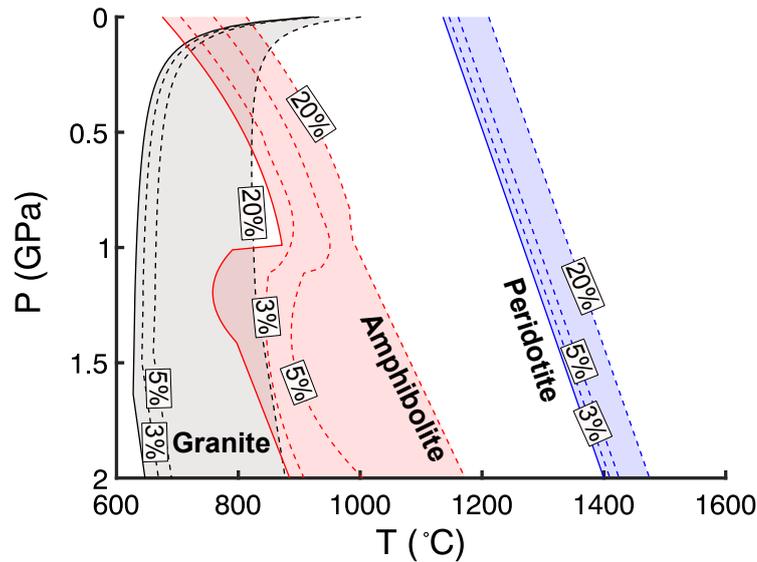


Fig. 2.5 Solidus of a hydrated granite [Boettcher and Wyllie, 1968] (in black), fluid-absent MORB-derived amphibolite [López and Castro, 2001] (in red) and fertile peridotite [Phipps Morgan, 2001b] (in blue). Dashed lines represent the solidus at different degrees of melt fraction.

be problematic if two different interfaces cross each other. For example, during continental break-up, the mesh comprising the upper crust, lower crust and mantle lithosphere is subject to extensive extension and thinning, which may result in the crossing of the interfaces between these rheological bodies nearby the spreading centre. This problem is fixed in LaCoDe by imposing a minimum distance between layers of around 100 and 1000 m, depending on the spatial resolution of the model. This issue can also be bypassed in LaCoDe by using an adaptive mesh generator. In this case, the interfaces between two distinct bodies do not perfectly match their contact, instead they are defined as higher resolution areas, and the code generates a cloud of randomly generated tracers (at least 6 tracers per element) that store the rheological phase. In every time step, the velocity field is interpolated onto the tracers, which are advected following the same procedure as for the FEM mesh. Then the rheological phase of a single element is defined as the median of the tracers inside the element. Employing this approach, rheological layers that were laterally continuous at the beginning of the simulation can actually break-up, thus avoiding problems of crossing interfaces.

2.4 Other processes

The code includes two additional features of significant relevance to geological processes: melt generation and serpentinisation. Their correspondent parametrisation is briefly described in the following sections.

2.4.1 Partial melting

The production of partial melt is calculated following Phipps Morgan [2001b]. The mantle solidus temperature T^s is defined as:

$$T^s = T_o^s + \left(\frac{\partial T^s}{\partial p} \right)_F p + \left(\frac{\partial T^s}{\partial F} \right)_P F \quad (2.88)$$

where T_o^s is the solidus temperature at the surface, $\partial T^s / \partial P$ is the solidus-pressure gradient, $\partial T^s / \partial F$ is the solidus-depletion gradient and F is melt fraction. Melting is produced in a parcel of the model if $T > T^s$, and two mechanisms are responsible for the production partial melting: 1) temperatures above the solidus,

Table 2.2 *Thermodynamic properties for mantle material. Values from Phipps Morgan [2001b]*

| Lithology | T_o^s (°C) | $\partial T^s/\partial p$ (°C/GPa) | $\partial T^s/\partial f$ (°C) | ΔH |
|-----------------------|-----------------|---------------------------------------|-----------------------------------|------------|
| Fertile peridotite | 1081 | 132 | 250 | 550 |
| Refractory peridotite | 1136 | 132 | 250 | 550 |

and 2) decompression. The decompression melt productivity for a lithology i within a lithology j is given by [Phipps Morgan, 2001b]:

$$-\frac{\partial F_i}{\partial p} = \frac{\frac{\partial T_i^s}{\partial p} - \frac{\alpha T}{\rho C_p} + \frac{T}{C_p} \phi_j \Delta S_j \left(\frac{\partial T_i^s}{\partial p} - \frac{\partial T_j^s}{\partial p} \right)}{\frac{T}{C_p} \phi_i \Delta S_j \left(\frac{\partial T_i^s/\partial F_i}{\partial T_j^s/\partial F_j} \right) + \frac{\partial T_i^s}{\partial F_i}} \quad (2.89)$$

where ΔS is the entropy of the solid-melt phase change, which can be related to the latent heat of melting ΔH , for a pure substance, as $\Delta H = T \Delta S$. A single component melting is considered in this thesis and the amount of decompression melting is defined as:

$$dF_{pressure} = dP \left(-\frac{\frac{\partial T^s}{\partial p}}{\frac{\Delta H}{C_p} + \frac{\partial T^s}{\partial F}} \right) \quad (2.90)$$

where the adiabatic term is missing because the temperatures are potential temperature. The temperature change during decompression melting is given by:

$$\frac{dT}{dP} = \frac{\partial T_s^i}{\partial p} + \frac{\partial T_s^i}{\partial F_i} \frac{dF_i}{dP} \quad (2.91)$$

The amount of melt under isobaric conditions is given by [Nielsen and Hopper, 2004]:

$$dF_{temp} = \frac{T^m - T^s}{\frac{\Delta H}{C_p} + \frac{\partial T^s}{\partial F}} \quad (2.92)$$

where T^m is the mantle temperature and the total amount of instantaneous melt is $dF = dF_{pressure} + dF_{temp}$. The total amount of melting produced in a parcel is then the summation of dF over time:

$$F = \Delta t \sum_{t=1}^n dF^t \quad (2.93)$$

where the superscript t is the time step and n is the total number of time steps. For undepleted mantle, the wet solidus (fertile peridotite in Table 2.2) is used initially, and the dry solidus (refractory peridotite in Table 2.2) is used after 2% melting [Braun et al., 2000]. Partial melting of the crustal is calculated in the same manner as melting of the mantle; however crustal T^s have been parametrised from solidus-liquidus curves obtained with experimental studies. For the work correspondent to Chapter 4, two different source of crustal melting are considered (see Figure 2.5): 1) a hydrated granite [Boettcher and Wyllie, 1968]; and 2) a fluid-absent MORB-derived amphibolite [López and Castro, 2001].

Since buoyancy forces due to melt production are relatively small, they can be included in the equation of state under the Boussinesq approximation:

$$\rho(T, P) = \rho_o(1 + \alpha(T - T_o) - \beta F) \quad (2.94)$$

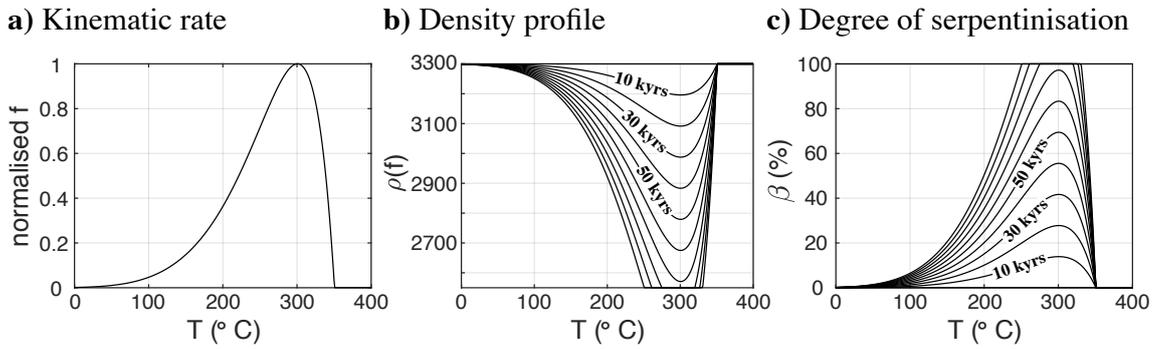


Fig. 2.6 (a) Kinematic rate as a function of temperature. Evolution of the (b) density and (c) degree of serpentinisation with temperature and time.

where ρ_o and T_o are the density and temperature at the surface temperature and zero pressure, respectively, α is the thermal expansivity, F is the depletion and β is defined as:

$$\beta = 1 - \frac{\rho_{molten}}{\rho_{solid}} \quad (2.95)$$

where ρ_{solid} and ρ_{molten} are the reference densities of the rock in its solid and molten states. The density of molten crust is taken $\rho = 2400 \text{ kg/m}^3$ and $\rho = 2900 \text{ kg/m}^3$ for molten mantle material [Gerya and Meilick, 2011].

2.4.2 Serpentinisation

Serpentinisation reactions occur when cold lithospheric mantle rocks react with seawater within the temperature limits (< 350 °C) of the serpentine group minerals (Fig. 2.6) and represent a relevant chemico-mechanical process that takes place in some tectonic events such as continental break-up and subducting slabs. The low friction angle and volumetric strain associated to these reactions is known to weaken the strength of the lithosphere [Escartin et al., 1997, 2001] and theorised to control the development of decollements at the crust-mantle boundary in slow-spreading oceanic ridges [Pérez-Gussinyé and Reston, 2001]. Additionally, serpentinisation occurring at oceanic transform faults associated to slow-spreading oceanic ridges may have a significant impact on global marine biogeochemical cycles (Rüpke and Hasenclever [2017], and references therein). Serpentinisation have been also identified between the outer rise and the trench of subducting slabs [Ranero et al., 2003], and the volumetric strain of associated to the formation of serpentine group minerals is inferred to enhance the bending of the subducting slab [Phipps Morgan, 2001a]. Moreover, it is commonly accepted that the deep (at ~ 250 - 300 km) dehydration of the slab is responsible for triggering arc melting [Rüpke et al., 2004], and the associated decrease of volyme may aid the unbending of the slab.

Two different mechanism are responsible of bringing seawater into contact with mantle rocks: i) exhumation of ultra mafic rocks, and ii) active faults that cut through the crust and reach the mantle lithosphere, resulting in the formation of conduits of seawater that reaches and reacts with mantle lithospheric rocks. The later mechanism has been observed under thinned continental crust in slow-spreading oceanic ridges [e.g. Pérez-Gussinyé and Reston, 2001; Rüpke et al., 2013; Rüpke and Hasenclever, 2017], and the amount of seawater reaching the mantle is thought to be controlled by the amount of displacement along the faults [Bayrakci et al., 2016]. Serpentinisation occurring at a crustal scale in subducting slabs has also been linked to active normal faults related to the bending slab [Ranero et al., 2003].

In LaCoDe, the degree of serpentinisation of the mantle is calculated in every time step for those parcels of the model under brittle failure (i.e. points of the model where $\tau = \tau_y$) and within the serpentinisation pressure and temperature stability conditions. The serpentinisation reaction is implemented assuming a temperature-dependent kinetic rate [Malvoisin et al., 2012]:

$$f(T) = C_o A \exp\left(-\frac{b}{T}\right) \left(1 - \exp\left[-c \left(\frac{1}{T} - \frac{1}{T_o}\right)\right]\right) \quad (2.96)$$

with $A = 808.3$, $b = 3640$ K, $T_o = 623.6$ K, $c = 8759$ K. The rate of density change due to serpentinization is then:

$$\frac{\partial \rho}{\partial t} = -f \rho_o \quad (2.97)$$

The phase change is then incorporated to a pressure and temperature dependant equation of state:

$$\rho(p, T, f) = \rho_o (1 - \alpha(T - T_o) + K^{-1} p - f \Delta t) \quad (2.98)$$

where ρ_o and T_o are the reference density and temperature, respectively, and K is the bulk modulus. Serpentinisation is an exothermic reaction, thus a term Q_{serp} that represents the rate heat generation by the reaction of serpentinisation is added to the equation of conservation on energy [Emmanuel and Berkowitz, 2006]:

$$Q_{serp} = H_{serp} \frac{\partial \rho_o}{\partial t} \quad (2.99)$$

where H is the thermal energy released during the hydration (or dehydration) reaction per unit mass of serpentinised mineral. It must be noted that density changes due to serpentinisation reactions may become larger than 10% with respect the reference density. Therefore, one should be extremely cautious (and drop the pressure term) if the eq. (2.98) is used under the incompressible Boussinesq approximation.

2.5 Code structure

The incompressible thermo-mechanical problem described in this chapter is solved using the code LaCoDe. This code is written in MATLAB and uses the optimised approach described in [Dabrowski et al., 2008] to build the block matrices that constitute the system of linear equations. Previous versions of LaCoDe included non-Newtonian flow and elastic deformation for an incompressible material. For this thesis, I have enhanced incompressible version of LaCoDe by adding plastic deformation, plastic softening and shear heating. I have also written a different version of LaCoDe that includes a fully compressible formulation (see Chapter 3 for a description and discussion). The global workflow of LaCoDe for the incompressible Boussinesq approximation is summarised in Fig. 2.7. The structure of the code can be sub-divided in three parts:

Pre-processor

In this part the thermo-mechanical properties and geometry defining the problem is defined. A triangular mesh is generated using the mesh generator Triangle [Shewchuk, 1996] or an adaptive mesh generator [Liu et al., 2018]. The velocity and temperature boundary conditions are also prescribed in this section of the code.

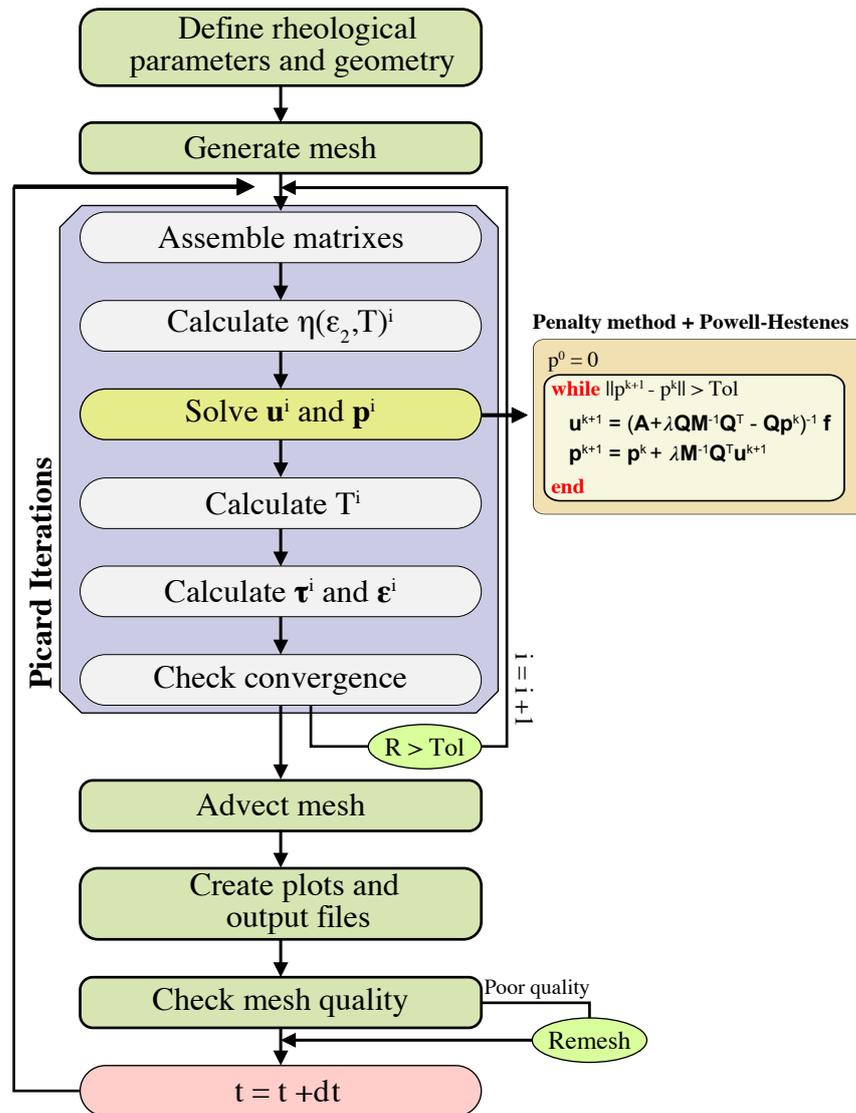


Fig. 2.7 Global work flow of the LaCoDe for the incompressible Boussinesq approximation, and details of the linear solver (combination of penalty method and Powell-Hestenes iterations).

Processor

This part of the codes solves the incompressible Stokes equations and thermal diffusion to obtain the velocity, pressure and temperature fields. After advection of the mesh, a remeshing algorithm is called if it is too distorted (this procedure is described in Chapter 3).

Post-processor

Other variables are here calculated from the velocity, pressure and temperature; for example: stress and strain fields or partial melting. Visualization algorithms are called to produce plots of the results of the models.

3 | LaCoDe: a Lagrangian two-dimensional thermo-mechanical code for large strain compressible viscoelastic geodynamical modelling

Albert de Montserrat, Jason P. Morgan and Jörg Hasenclever. LaCoDe: a Lagrangian two-dimensional thermo-mechanical code for large strain compressible viscoelastic geodynamical modelling. Planned for submission to *Tectonophysics*.

Authors contribution

AdM and JM designed the mathematical description of a compressible visco-elastic flow and the numerical implementation was introduced by AdM in collaboration with JM and JH. AdM designed and analysed the results of the numerical experiments in discussion with JM and JH. AdM wrote the manuscript under the supervision of JM and JH.

LaCoDe: a Lagrangian two-dimensional thermo-mechanical code for large strain compressible visco-elastic geodynamical modeling

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Abstract

We present the numerical code LaCoDe (Lagrangian Compressible Deformation) for two-dimensional problems for mantle-lithosphere geodynamic modeling. Unlike a great number of mantle modeling codes that are based on the incompressible Boussinesq approximation, LaCoDe adopts a compressible continuity equation, including the volumetric strains linked to elastic compression. LaCoDe is a finite element method code that typically uses a Lagrangian frame of reference. It solves the Stokes equations for a non-Newtonian visco-elastic rheology. A remeshing algorithm is implemented to track and transfer the physical parameters of the material from a heavily distorted mesh into to an updated one. In this paper, we first describe the equations governing the deformation of Earth materials with detailed description of the algorithm and its numerical implementation. We then benchmark the accuracy of LaCoDe by comparing numerical results with analytical solutions for bending of a thin elastic beam under a constant uniform load, viscous inclusions, Rayleigh-Taylor instabilities, stress build-up in a visco-elastic Maxwell body, and Couette flow with viscous flow. The Rayleigh-Taylor test is further used to demonstrate the accuracy of the remeshing algorithm. Finally, we show the importance of including volumetric strain in some crust-lithospheric deformation cases, such as magma-free slow-spreading ridges and subducting slabs. Additionally, we find that the ex-

tra non-linearities introduced by the compressible Stokes equations are better solved using nested sets of Picard iterations.

Keywords:

, Modeling of large-strain visco-elastic deformation, Compressible formulation

1 Introduction

Rocks are exposed to thermal, mechanical and chemical processes that induce volumetric changes. Obvious examples are mechanical compression and decompression, thermal expansion, and phase changes resulting from partial melting and serpentinization. Even though stresses related to compressibility may play an important role in rock deformation and failure, the incompressible Boussinesq approximation of the governing equations is the most common approach used in geodynamical modeling of coupled lithospheric-mantle systems. This approximation is considered to be reasonably valid under lithospheric conditions and offers a simple and straightforward numerical implementation, hence its popularity. The Boussinesq approximation is considered to be appropriate if: 1) the size of the domain is shorter than any physical scale height (i.e. $D_f = |f_m^{-1}df_o/dz|^{-1}$, where f is any state variable, f_m is the space average of f , and f_o is the variation in the absence of motion (Spiegel and Veronis, 1959)); 2) the density of the material does not change more than 10% with respect to its reference value (Gray and Giorgini, 1976); and, 3) volume-change-related stresses are small with respect to the lithostatic pressure and deformation-linked stresses.

These approximations are usually valid for lithospheric-scale models, but may be violated in certain scenarios. For instance, it is well known that metamorphic phase changes occurring at crustal conditions can induce significant changes in density in localised regions that far exceed the maximum density changes thought to be appropriate for the Boussinesq approximation. In the case of partial serpentinization, for example, density can be reduced by up to 18%, and the associated volumetric strains can cause rocks to fail. This

26 mechanism potentially reduces the strength of the lithosphere by 30% (Escartin
27 et al., 1997), or even more when intact rock is replaced by a serpentinized fault.
28 Volume-change-linked stresses related to phase changes may therefore have a
29 significant influence on the localisation of deformation when brittle failure is an
30 important rheological feature.

31 The first studies proposing a compressible formulation for mantle deforma-
32 tion (Jarvis and McKenzie, 1980; Quarenì et al., 1986; Yuen et al., 1987) made
33 use of the so-called anelastic approximation. These studies were aimed at un-
34 derstanding the behaviour of deep mantle convection, while implications for
35 lithospheric failure and deformation were not considered. In the last decades
36 numerous studies focused on the development of numerical tools to investigate
37 lithospheric and upper mantle geodynamical processes (e.g. Christensen, 1987;
38 Braun and Sambridge, 1994; Fullsack, 1995; Schmalholz et al., 2001; Moresi
39 et al., 2003; Petrunin and Sobolev, 2006; Gerya and Yuen, 2007; von Tschärner
40 and Schmalholz, 2015). However, all of these studies employed the Boussinesq
41 incompressible approximation. To date, relatively little effort has been made to
42 include and discuss the effects of volumetric strains at the lithospheric scale. To
43 our knowledge, SLIM3D (Popov and Sobolev, 2008) and DynEarthSol2D (Choi
44 et al., 2013) are the only available numerical models that include elastic com-
45 pressibility. However, these studies do not assess its implications for lithospheric
46 scale processes.

47 We propose a new compressible formulation that has been implemented in
48 the 2-D geodynamic code LaCoDe (Hasenclever, 2010; Hasenclever et al., 2011).
49 LaCoDe solves for visco-elastic deformation, thermal convection and melting
50 processes. It is written in MATLAB and uses an optimized assembly based
51 on the 'blocking' and vectorization approaches described in Dabrowski et al.
52 (2008). Stokes equations are solved using a Lagrangian mixed velocity-pressure
53 approach with the Finite Element Method (FEM). An additional feature of
54 LaCoDe, not discussed here, is a free-surface algorithm (Andrés-Martínez et al.,
55 2015) that allows the tracking of the evolution of topographic relief.

56 The purpose of this paper is to assess the stability of the numerical im-

57 plementation of a visco-elastic rheology that does not assume the incompress-
58 ible Boussinesq approximation and emphasize its relevance for some geologi-
59 cal events at a lithospheric scale. We first describe the new formulation and
60 its numerical implementation. Then we test the accuracy of our code with
61 a series of benchmarks for viscous and elastic deformation: 1) bending of an
62 elastic cantilever under a uniform load, 2) deformation around a viscous in-
63 clusion, 3) Rayleigh-Taylor instability, 4) steady-state thermal convection, 5)
64 build-up of stress in a visco-elastic Maxwell body, 6) Couette-flow of a fluid with
65 temperature-dependent viscosity and viscous heating effects. We then present
66 two examples of tectonic processes where compressibility might be an important
67 mechanism: 1) volumetric strain linked to phase changes, and 2) comparison
68 of a compressible and incompressible subducting slab. Finally, we prove that
69 when non-linear rheologies are employed with a compressible formulation, it is
70 more convenient to use nested Picard iterations.

71 2. Governing equations for compressible flow

72 Mantle-lithosphere deformation are considered to be a thermo-mechanical
73 process described by the equations of continuity, conservation of momentum
74 and conservation of energy, respectively, in a domain Ω :

$$\frac{\partial \sigma_{ij}}{\partial x_j} = -\rho g_i \quad (1)$$

75

$$\frac{D\rho}{Dt} + \rho \frac{\partial u_i}{\partial x_i} = q_m \quad (2)$$

76

$$\rho C_p \frac{DT}{Dt} = \frac{\partial}{\partial x_i} \left(\kappa \frac{\partial T}{\partial x_i} \right) + \alpha T \frac{Dp}{Dt} + H_q + H_{sh} \quad (3)$$

77 where ρ is density, x_i are the spatial coordinates with the indexes i, j referring
78 to the directions x and z in a two-dimensional Cartesian coordinate system,
79 respectively, u_i are the velocity components, σ_{ij} is the Cauchy stress tensor,
80 g_i is the gravitational acceleration, C_p is heat capacity, T is temperature, κ is
81 thermal conductivity, α is thermal expansivity, H_q is a heat source, and shear

82 heating is defined as the energy released by the inelastic work $H_{sh} = \sigma_{ij}\varepsilon_{ij}^{inel}$.
83 The function $q_m = q(x, t)$ in eq. (2) describes the rate of mass being added
84 (local source of mass: $q_m > 0$) or subtracted (local sink of mass: $q_m < 0$) from
85 a region, with dimensions of mass per unit volume and unit time. Note that,
86 when a Lagrangian frame of reference is adopted, the material time derivative
87 $D(\cdot)/Dt$ is equal to the partial time derivative $\partial(\cdot)/\partial t$.

88 The set of equations (1), (2) and (3) describe the thermo-mechanical be-
89 haviour of a compressible viscous flow. Several approximations of these equa-
90 tions have been widely employed to address the compressibility of the mantle,
91 such as the anelastic approximation (ALA) or the truncated anelastic approx-
92 imation (TALA) (e.g. Jarvis and McKenzie, 1980; Bercovici et al., 1992; King
93 et al., 2010). On the other hand, models studying geodynamic processes at a
94 lithospheric scale (e.g. from rifting of continental crust to subduction zones)
95 widely employ the so-called incompressible Boussinesq approximation, where
96 the continuity equation is simplified as divergence-free. In the (T)ALA ap-
97 proximations the dynamic pressure is assumed negligible with respect the hy-
98 drostatic pressure ($p_{dyn} \ll p_{total}$), leading to a depth-dependent density. We
99 propose a formulation where the continuity equation is directly computed using
100 its Lagrangian form, employing an equation of state that depends on the total
101 pressure:

$$\rho(T, p) = \rho_o [1 - \alpha(T - T_o) + K^{-1}(p - p_o)] \quad (4)$$

102 where ρ_o , T_o , p_o are the reference density, temperature and pressure, K is the
103 bulk modulus and p is total pressure. It is convenient to define a the reference
104 density, for example, using the Adams-Williamson equation, or an approxima-
105 tion of the hydrostatic pressure, so that the volumetric changes are with respect
106 to the reference state. If one wishes, density changes due to phase changes can
107 be incorporated to the equation of state. The density time derivative in the
108 continuity equation is then computed in an implicit manner, so that eq. (2) is
109 approximated as:

$$\frac{\partial u_i^{n+1}}{\partial x_i} = \frac{1}{\rho^{n+1}} \left(q_m - \frac{\rho^{n+1} - \rho^n}{\Delta t} \right) \quad (5)$$

110 where the superscript n indicates the time step iteration, and Δt is the time
 111 step. The time derivative of the density introduces a non-linearity in the system
 112 of equations and eq. (2) can also be solved either in an explicit manner. A
 113 comparison between both approaches has been discussed in Heister et al. (2017)
 114 and, a priori, it is not obvious whether one approach is numerically more stable
 115 and/or more efficient than the other. By definition, the explicit approach would
 116 require less non-linear iterations than the implicit approach; however, Heister
 117 et al. (2017) concluded that both approaches yield equally accurate results at
 118 similar computational time requirements.

119 2.1. Mixed formulation

120 The implementation of a mixed formulation to solve the Stokes equations
 121 splits the Cauchy stress tensor into its deviatoric and hydrostatic components:

$$\sigma_{ij} = \tau_{ij} - p\delta_{ij} \quad (6)$$

122 where τ_{ij} is the deviatoric stress tensor, δ_{ij} is the Kroenecker delta and the
 123 pressure is the mean of the principal stresses $p = -\sigma_{kk}/3$. Using eq. (6),
 124 the conservation of momentum is written in terms of the deviatoric stress and
 125 pressure:

$$\frac{\partial \tau_{ij}}{\partial x_j} - \frac{\partial p}{\partial x_i} = -\rho g_i \quad (7)$$

126 2.2. Constitutive equation of a visco-elastic fluid

127 The viscous constitutive law is conveniently expressed in terms of deviatoric
 128 stress τ_{ij} and deviatoric strain rate $\dot{\epsilon}_{ij}$:

$$\tau_{ij} = 2\eta\dot{\epsilon}_{ij} \quad (8)$$

129 where η is the shear viscosity, and the deviatoric strain rate tensor is defined
 130 as:

$$\dot{\epsilon}_{ij} = \frac{1}{2} \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) - \frac{1}{3} \frac{\partial u_k}{\partial x_k} \delta_{ij} \quad (9)$$

131 Elastic deformation is incorporated by adopting a Maxwell material model,
 132 where the visco-elastic deviatoric strain rate is the sum of the viscous and elastic

133 strain rates:

$$\dot{\epsilon}_{ij} = \dot{\epsilon}_{ij}^{visc} + \dot{\epsilon}_{ij}^{el} = \frac{\tau_{ij}}{2\eta} + \frac{\check{\tau}_{ij}}{2G} \quad (10)$$

134 where G is the shear modulus and $\check{\tau}_{ij}$ is the objective deviatoric stress rate (e.g.
135 Hashiguchi and Yamakawa, 2012). The Zaremba-Jaumann derivative is used to
136 compute the objective deviatoric stress rate in eq. (10):

$$\check{\tau}_{ij} = \frac{\partial \tau_{ij}}{\partial t} - \omega_{ik}\tau_{kj} + \tau_{ik}\omega_{kj} \quad (11)$$

137 where $\omega_{ij} = 1/2(\partial u_i/\partial x_j - \partial u_j/\partial x_i)$ is the spin tensor associated with the rigid
138 body rotation. Following the implementation of large-strain elastic deformation
139 described by Moresi et al. (2003) and Kaus (2010), $\check{\tau}_{ij}$ is approximated by an
140 implicit discretisation of the time derivative:

$$\check{\tau}_{ij} \approx \frac{\tau_{ij}^{n+1} - \tau_{ij}^n}{\Delta t} - \omega_{ik}^n \tau_{kj}^n + \tau_{ik}^n \omega_{kj}^n \quad (12)$$

141 Substitution of eq. (12) into eq. (10) with subsequent rearrangement of the
142 terms leads to the visco-elastic constitutive law:

$$\tau_{ij} = 2\eta_{eff}\dot{\epsilon}_{ij} + \chi\hat{\tau}_{ij} \quad (13)$$

143 where

$$\eta_{eff} = \frac{1}{\frac{1}{\eta} + \frac{1}{G\Delta t}} \quad (14)$$

144

$$\chi = \frac{1}{1 + \frac{G\Delta t}{\eta}} \quad (15)$$

145

$$\hat{\tau}_{ij} = \tau_{ij}^n + (\omega_{ik}^n \tau_{kj}^n - \tau_{ik}^n \omega_{kj}^n)\Delta t \quad (16)$$

146 were the “real” viscosity has been substituted by an effective viscosity η_{eff} that
147 includes the elastic terms. A pure viscous rheology is recovered if $\Delta t \rightarrow \infty$.
148 Note that the visco-elastic deformation obtained per time step depends on the
149 size of the time step. However, the deformation after a certain simulation time
150 has to be independent of the chosen time step.

151 *2.3. Viscous creep*

152 Two mechanisms for viscous deformation are included in our model: diffusion
 153 creep and dislocation creep (Poirier, 1985; Karato et al., 2001). Diffusion creep
 154 occurs at low stress levels, when atoms diffuse inside the crystal grains and
 155 along the grain boundaries, resulting deformation of the rock. Deformation
 156 due to dislocation creep is caused by the migration of dislocations through the
 157 crystal lattice of the rock. Both creep mechanisms are strain rate-, temperature-
 158 and pressure- dependent:

$$\eta_{dif} = \frac{1}{2}(A)^{-\frac{1}{n}}(\dot{\varepsilon}_{II}^{dif})^{\frac{1}{n}-1} \exp\left(\frac{E_a + pV_a}{nRT}\right) \quad (17)$$

159

$$\eta_{dis} = \frac{1}{2}(A)^{-\frac{1}{n}}(\dot{\varepsilon}_{II}^{dis})^{\frac{1}{n}-1} \exp\left(\frac{E_a + pV_a}{nRT}\right) \quad (18)$$

160 where A is the pre-exponential parameter, n is the power-law exponent (with
 161 $n = 1$ for diffusion creep and, theoretically, $n \approx 3$ for dislocation creep), $\dot{\varepsilon}_{II} =$
 162 $\sqrt{(1/2)\dot{\varepsilon}_{ij}\dot{\varepsilon}_{ij}}$ is the square root of the second invariant of the deviatoric strain
 163 rate tensor, E_a is the activation energy, V_a is the activation volume and R is the
 164 universal gas constant. We now build an effective creep viscosity, using diffusion
 165 and dislocation viscosities in parallel:

$$\frac{1}{\eta} = \frac{1}{\eta_{dif}} + \frac{1}{\eta_{dis}} \quad (19)$$

166 In this way, the smallest viscosity will have the largest contribution to the
 167 effective viscosity, with deformation dominated by the mechanism that has the
 168 smallest activation stress. The viscous strain tensor is then $\dot{\varepsilon}_{ij}^{visc} = \dot{\varepsilon}_{ij}^{dif} + \dot{\varepsilon}_{ij}^{dis}$
 169 and, using the definitions (17) and (18), the diffusion and dislocation strain
 170 tensors are respectively computed as:

$$\dot{\varepsilon}_{ij}^{dif} = \frac{\tau_{ij}}{2\eta_{dif}}; \quad \dot{\varepsilon}_{ij}^{dis} = \frac{\tau_{ij}}{2\eta_{dis}} \quad (20)$$

171 **3. Numerical implementation**

172 LaCoDe solves the resulting set of governing equations of the thermo-mechanical
 173 problem using the FEM to generate the system of matrix equations (e.g. Hughes,

174 1987; Zienkiewicz and Taylor, 2005). Discretizing the domain into elements, the
 175 primary variables u , p and T are approximated using the shape functions N_u
 176 for velocity, N_p for pressure and N_T for temperature:

$$u(x, y) \approx \sum_{a=1}^n N_u^a(x, y) \tilde{u}_a \quad (21)$$

177

$$p(x, y) \approx \sum_{a=1}^n N_p^a(x, y) \tilde{p}_a \quad (22)$$

178

$$T(x, y) \approx \sum_{a=1}^n N_T^a(x, y) \tilde{T}_a \quad (23)$$

179 where the subscript a is the nodal index and n is the number of nodes in the
 180 element. Employing the Galerkin procedure, the governing eqs. (1), (2) and
 181 (3) are transformed into their weak forms using the shape functions as trial
 182 functions.

183 The choice of the approximation space for the coupled velocity-pressure prob-
 184 lem has to be taken carefully so that the so-called LBB (or *inf-sup*) condition
 185 is satisfied. Some combinations of approximation spaces for velocity and pres-
 186 sure will violate such condition and result in spurious pressure modes and/or
 187 non-converged flow solutions. In LaCoDe, the LBB condition is satisfied by us-
 188 ing Crouzeix-Raviart triangular elements (Crouzeix and Raviart, 1973), where
 189 the velocity field is approximated by seven nodal points and quadratic inter-
 190 polation enhanced by a cubic bubble function in the baricenter of the element
 191 (Fig. .1). Pressure is discontinuous with three nodal points describing a linear
 192 interpolation within each element.

193 In the following sections we detail the strong forms of the Stokes and thermal
 194 diffusion equations as well as their numerical implimentation, where we drop the
 195 $\tilde{\cdot}$ from the approximated fields in order to simplify the notation. The reader is
 196 referred to FEM textbooks (e.g. Hughes, 1987; Zienkiewicz and Taylor, 2005) for
 197 more details on the method and the description of the weak formulation of the
 198 Stokes and thermal diffusion equations.

199 *3.1. FEM formulation of thermal diffusion*

200 The time derivatives in eq. (3) are approximated using a backward Euler
201 discretisation:

$$\rho C_p \left(\frac{T^{n+1} - T^n}{\Delta t} \right) = \frac{\partial}{\partial x_i} \left(k \frac{\partial T^{n+1}}{\partial x_i} \right) + \alpha T^{n+1} \frac{p^{n+1} - p^n}{\Delta t} + H_r + H_{sh} \quad (24)$$

202 Using FEM for the spatial discretization in the space and rearranging eq. (24),
203 we can express it in a compact matrix notation:

$$\mathbf{K}_T \mathbf{T} = \mathbf{f}_T \quad (25)$$

204 where the stiffness matrix is:

$$\begin{aligned} \mathbf{K}_T = \int_{\Omega} \nabla \mathbf{N}_T k \nabla \mathbf{N}_T d\Omega + \frac{1}{\Delta t} \int_{\Omega} \mathbf{N}_T^T \rho^{n+1} C_p \mathbf{N}_T d\Omega + \\ \frac{1}{\Delta t} \int_{\Omega} \mathbf{N}_T^T \alpha \mathbf{N}_u (p^{n+1} - p^n) \mathbf{N}_T d\Omega \end{aligned} \quad (26)$$

205 and the right-hand-side vector:

$$\mathbf{f}_T = \frac{1}{\Delta t} \int_{\Omega} \mathbf{N}_T^T \rho^{n+1} C_p T^n \mathbf{N}_T d\Omega + \int_{\Omega} \mathbf{N}_T H_r d\Omega + \int_{\Omega} \mathbf{N}_T H_{sh} d\Omega \quad (27)$$

206 We use the same shape functions for temperature as velocity, i.e. $N_T = N_u$.

207 *3.2. FEM formulation of Stokes equations*

208 The motion of a compressible visco-elastic flow is described by the Stokes
209 equations (1) and (2). The weak forms of the Navier-Stokes equations can be
210 expressed in matrix form as:

$$\int_{\Omega} \mathbf{B}^T \mathbf{D} \mathbf{B} \mathbf{u}^{n+1} d\Omega - \int_{\Omega} \mathbf{B}^T \mathbf{m} \mathbf{N}_p p^{n+1} d\Omega = \int_{\Omega} \mathbf{N}_u^T \rho \mathbf{g} d\Omega - \int_{\Omega} \mathbf{B}^T \chi \hat{\tau} d\Omega \quad (28)$$

211

$$\int_{\Omega} \mathbf{N}_p^T \mathbf{m}^T \mathbf{B} \mathbf{u}^{n+1} d\Omega = \int_{\Omega} \mathbf{N}_p^T \left(\frac{1}{\rho^{n+1}} \left(q_m - \frac{\rho^{n+1} - \rho^n}{\Delta t} \right) \right) d\Omega \quad (29)$$

212 The elemental matrix \mathbf{B}^e represents the strain-displacement matrix and \mathbf{D}^e
213 is the rheology matrix that relates strain rates to deviatoric stresses:

$$\mathbf{B}^e \mathbf{u}^e = \begin{bmatrix} \frac{\partial N_u}{\partial x} & 0 \\ 0 & \frac{\partial N_u}{\partial z} \\ \frac{\partial N_u}{\partial z} & \frac{\partial N_u}{\partial x} \end{bmatrix} \begin{bmatrix} u_x \\ u_z \end{bmatrix} = \begin{bmatrix} \dot{\epsilon}_{xx} \\ \dot{\epsilon}_{zz} \\ \dot{\epsilon}_{xz} \end{bmatrix} \quad (30)$$

$$\mathbf{D}^e = \eta_{eff} \begin{bmatrix} C_1 & C_2 & 0 \\ C_2 & C_1 & 0 \\ 0 & 0 & 1 \end{bmatrix} \quad (31)$$

$$\mathbf{m}^T = [1 \quad 1 \quad 0] \quad (32)$$

214 The \mathbf{m}^T vector is necessary when the cross derivatives in the last row in of
 215 the matrix \mathbf{B} are not necessary. In the compressible case the coefficients in the
 216 rheology matrix \mathbf{D}^e take values of $C_1 = 4/3$ and $C_2 = -2/3$. The weak forms
 217 (28) and (29) can then be written in a compact matrix notation as:

218 [Figure 1 about here.]

$$\begin{pmatrix} \mathbf{A} & \mathbf{G} \\ \mathbf{G}^T & \mathbf{0} \end{pmatrix} \cdot \begin{pmatrix} \mathbf{u} \\ \mathbf{p} \end{pmatrix} = \begin{pmatrix} \mathbf{f}_1 \\ \mathbf{f}_2 \end{pmatrix} \quad (33)$$

219 where:

$$\mathbf{A} = \int_{\Omega} \mathbf{B}^T \mathbf{D} \mathbf{B} d\Omega \quad (34)$$

$$\mathbf{G} = - \int_{\Omega} \mathbf{B}^T \mathbf{m} \mathbf{N}_p d\Omega \quad (35)$$

$$\mathbf{f}_1 = \int_{\Omega} \mathbf{N}_u^T \rho \mathbf{g} d\Omega - \int_{\Omega} \mathbf{B}^T \chi \hat{\tau} d\Omega \quad (36)$$

$$\mathbf{f}_2 = \int_{\Omega} \mathbf{N}_p^T \left(\frac{1}{\rho^{n+1}} \left(q_m - \frac{\rho^{n+1} - \rho^n}{\Delta t} \right) \right) d\Omega \quad (37)$$

220 and the right-hand-side vector f_2 contains the non-zero divergence terms related
 221 to density changes.

222 3.3. Solution scheme of Stokes equations

223 The expression (33) mathematically describes the so-called *saddle point*
 224 problem. Numerical complications arise due to the presence of the diagonal
 225 zero-block in the full matrix, which makes the matrix positive semi-definite,
 226 and thus it cannot be solved directly with numerical algorithms such as Con-
 227 jugate Gradient or Cholesky factorization. LaCoDe solves the Stokes equation

228 using the Augmented Lagrangian method (Rockafellar, 1974), which consists of
 229 subtracting $\lambda^{-1}\mathbf{M}\mathbf{p}$ from the left- and right-hand-side of the continuity equa-
 230 tion, and introducing the following iterative scheme:

$$\begin{pmatrix} \mathbf{A} & \mathbf{G} \\ \mathbf{G}^T & -\lambda^{-1}\mathbf{M} \end{pmatrix} \cdot \begin{pmatrix} \mathbf{u} \\ \mathbf{p} \end{pmatrix}^{k+1} = \begin{pmatrix} \mathbf{f}_1 \\ \mathbf{f}_2 - \lambda^{-1}\mathbf{M}\mathbf{p}^k \end{pmatrix} \quad (38)$$

231 where k is the iteration counter, λ is an artificial compressibility term penalising
 232 the new pressure term in the second row of the global matrix and has units of
 233 dynamic viscosity, and \mathbf{M} is the mass matrix defined as:

$$\mathbf{M} = \int_{\Omega} \mathbf{N}_p^T \mathbf{N}_p d\Omega \quad (39)$$

234 The choice of λ is not trivial, as the block matrix might become ill-posed or
 235 numerical locking might occur if λ is either too high or too low. A value of $\lambda =$
 236 $\max(\eta)$ has been proven to work well in our benchmarks. Upon convergence,
 237 $\mathbf{p}^{k+1} = \mathbf{p}^k$ and the system of equations (33) is recovered. The new system of
 238 equations (38) allows the elimination of the pressure field, and the first and
 239 second rows of the system are solved in a segregated manner. Rearranging the
 240 second equation we obtain the expression for the updated pressure:

$$\mathbf{p}^{k+1} = \mathbf{p}^k + \mathbf{M}^{-1}(\lambda\mathbf{G}^T \mathbf{u}^{k+1} - \mathbf{f}_2) \quad (40)$$

241 and after substitution of eq. (40) into the first equation in the system (38) we
 242 obtain the following linearised expression for the velocity field:

$$\mathbf{u}^{k+1} = \mathbf{K}^{-1}\mathbf{f}^{k+1} \quad (41)$$

243 where the stiffness matrix \mathbf{K} is defined as:

$$\mathbf{K} = \left(\mathbf{A} + \mathbf{G}\lambda\mathbf{M}^{-1}\mathbf{G}^T \right) \quad (42)$$

244 and the force vector in the right-hand-side is:

$$\mathbf{f}^{k+1} = \mathbf{f}_1 + \mathbf{G}(\lambda\mathbf{M}^{-1}\mathbf{f}_2 - \mathbf{p}^k) \quad (43)$$

245 The expression (40) is clearly non-linear because the density in \mathbf{f}_2 depends
 246 on the pressure via the equation of state. We treat this non-linearity by adding

247 a set of Picard iterations and freezing the density during the Powell-Hestenes
 248 iterations:

$$\begin{aligned}
 & \overbrace{\nabla \cdot \mathbf{u}^{k+1} + \frac{1}{\lambda} \mathbf{p}^{k+1}}^{\text{new velocity and pressure}} = \frac{1}{\underbrace{\rho(\mathbf{P}^m, \mathbf{T}^m)}_{\text{previous Picard iteration}}} \quad . \\
 & \left(q_m - \frac{\overbrace{\rho(\mathbf{P}^m, \mathbf{T}^m)}^{\text{previous Picard iteration}} - \overbrace{\rho(\mathbf{P}^n, \mathbf{T}^n)}^{\text{previous time step}}}{\Delta t} \right) = \mathbf{f}_2^m \quad (44)
 \end{aligned}$$

249 where the superscripts k , m and n are the counters of the Powell-Hestenes,
 250 Picard and time iterations, respectively. Eqs. (40) and (41) are thus solved
 251 iteratively combining Powell-Hestenes and Picard iterations in the following
 252 scheme (Fig. .2):

- 253 1. $\mathbf{p}^0 = 0$ for $n = 1$, and $\mathbf{p}^0 = \mathbf{p}^{n-1}$ for $n > 1$.
- 254 2. Calculate: \mathbf{K} .
- 255 3. Calculate: \mathbf{f}_2^m
- 256 4. Calculate: \mathbf{f}^{k+1}
- 257 5. Solve: $\mathbf{u}^{k+1} = \mathbf{K}^{-1} \mathbf{f}^{k+1}$
- 258 6. Update pressure: $\mathbf{p}^{k+1} = \mathbf{p}^k + \mathbf{M}^{-1}(\lambda \mathbf{G}^T \mathbf{u}^{k+1} - \mathbf{f}_2^m)$
- 259 7. Check convergence of the continuity equation. If $\| -\mathbf{Q}^T \mathbf{u} - \mathbf{g}_2 \|_\infty > \text{Tol}$,
 260 and repeat steps 4 and 7.
- 261 8. If $\| \mathbf{f}_2^m - \mathbf{f}_2^{m+1} \|_\infty > \text{Tol}$, repeat steps 3 to 7.

262 where $\| \cdot \|_\infty$ is the infinity norm. We note that for $\mathbf{p}^0 = 0$, the equations
 263 are equivalent to the penalty method. The solution scheme presented here is
 264 equivalent to the resulting schemes from Uzawa iterations (Arrow et al., 1958;
 265 Zienkiewicz, 1985) and later extended in the context of optimization indepen-
 266 dently by Hestenes (Hestenes, 1969) and Powell (Powell, 1967).

267 [Figure 2 about here.]

268 *3.4. Iteration scheme for non-linear rheology*

269 The problem described in Section 3.3 becomes even more non-linear if tem-
 270 perature and/or a non-Newtonian rheology are also considered. We propose two
 271 different approaches to tackle highly non-linear problems (Fig. .2): i) all the
 272 non-linearities are treated within a single loop of Picard iterations (Approach
 273 1); and, ii) the rheological and density non-linearities are split into two levels of
 274 nested Picard iterations (Approach 2). While Approach 2 is likely to increment
 275 the total number of linear and non-linear iterations for a single time step, the
 276 rheological non-linearities are performed in a presumably better converged flow
 277 solution. The rheology iterations are stopped when the residual R is below a
 278 given tolerance:

$$R = \frac{\|u^{i+1} - u^i\|_\infty}{\|u^{i+1}\|_\infty} \leq Tol \quad (45)$$

279 where i is the rheology iteration counter, and we take a typical value of $Tol =$
 280 10^{-3} . We note that this iterative scheme is able to handle other kinds of rheology
 281 non-linearities not included in this paper, such as plastic deformation. The
 282 efficiency of both methods is compared in Section 5.2.

283 **4. Remeshing**

284 One of the drawbacks of using a Lagrangian formulation is that large de-
 285 formation of the mesh may lead to highly distorted elements. This issue is
 286 overcome by mapping the necessary variable fields onto a newly generated high
 287 quality mesh. One could perform a remeshing after every time step, but to
 288 reduce the associated computational cost and interpolation errors, a new mesh
 289 is generated only when the quality of the mesh is below a given threshold. Let
 290 us define a triangle with the area A , vertices a , b and c , and the smallest and
 291 largest angles α and β , respectively. We define the quality factor of the triangle
 292 to be:

$$q_n = \frac{4\sqrt{3}A}{\|ab\|^2 + \|ac\|^2 + \|bc\|^2} \quad (46)$$

293 where q_n is a measurement of how close a triangle is to be equilateral. The
 294 remeshing algorithm is called only if one (or several) triangular element has
 295 $q_n < Tol_q$, $\alpha < Tol_\alpha$ or $\beta > Tol_\beta$. Unless specified, we use values of $Tol_{q_n} =$
 296 0.25 , $Tol_\alpha = 7^\circ$ and $Tol_\beta = 170^\circ$.

297 For fields that are computed at the nodes (i.e. temperature), the 6-node
 298 elements are split into 3-nodes elements and the fields are linearly interpolated
 299 into the new nodal positions. The information of the fields associated with the
 300 elements (i.e. stress, density) is stored at the integration points of the elements
 301 and they are mapped onto the new mesh using the following procedure:

- 302 1. Find the element of the old mesh containing the new integration point us-
 303 ing the quick search algorithm *tsearch2* (*Mutils* package: <http://milamin.sourceforge.net/downloads>).
- 304 2. Calculate local coordinates of the new integration point with respect to
 305 the element in the old mesh.
- 306 3. The field $\Psi(x, y)$ is mapped element-to-element onto the old nodes of using
 307 linear shape functions:

$$\Psi_a(x, y) = (N^a(\xi, \eta))^{-1} \Psi(x', y') \quad (47)$$

308 where a is the nodal index, ξ and η are the local coordinates of the shape
 309 function and x' and y' are the coordinates of the integration point of the
 310 old mesh.

- 311 4. The nodal values of target field $\Psi_a(x, y)$ are mapped onto the new inte-
 312 gration point using the shape functions:

$$\Psi(x^*, y^*) = \sum_{a=1}^n N^a(\xi, \eta) \Psi_a(x, y) \quad (48)$$

313 where ξ and η are the local coordinates of the shape function and x^* and
 314 y^* are the coordinates of the integration point of the new mesh.

315 While this scheme works particularly well for perfect body-fitting meshes,
 316 for which each element of the new and old meshes belongs to a single material
 317 phase, other approaches may be better suited for non-body-fitting meshes. The
 318 accuracy of this remeshing scheme is demonstrated in Section 5.1.3.

319 [Figure 3 about here.]

320 5. Results

321 We present a set of benchmarks and numerical experiments to test the im-
322 plementation of the formulation described above. We first demonstrate the
323 accuracy of LaCoDe, comparing the results of these experiments with analyti-
324 cal solutions and results from previously published studies. These benchmarks
325 are: i) bending of a thin beam under a distributed load (Turcotte and Schubert,
326 2014); ii) deformation around a viscous inclusion (Schmid and Podladchikov,
327 2003); iii) Rayleigh-Taylor instability (van Keken et al., 1997); iv) stress build-
328 up in a visco-elastic Maxwell body (Gerya and Yuen, 2007); and v) solution of
329 a Couette-flow with viscous heating and temperature-dependent viscosity (Tur-
330 cotte and Schubert, 2014). Then, we investigate the effectiveness of the two
331 approaches to solve problems with non-linear rheologies described in Section
332 3.4. Finally, two tectonic scenarios where the effect of compressibility effects is
333 relevant are presented: i) an example of volumetric strain produced by phase
334 changes; ii) subduction of a compressible slab.

335 5.1. Benchmarks

336 5.1.1. Cantilever beam under a uniform load

337 In this benchmark we compare the numerical results of a bending elastic
338 thin plate, clamped at one end, against an analytical solution for a perfectly-
339 elastic material (Turcotte and Schubert, 2014). We also use this benchmark
340 to compare the accuracy of the non-linearised and linearised formulations in
341 resolving elastic problems. The ratio between the thickness and length of the
342 cantilever is taken to be 1/10 in order to satisfy the thin beam hypothesis. The
343 density of the beam is $\rho = 150 \text{ kg/m}^3$ (an approximate value for the density
344 contrast between the upper and lower crust) and the shear modulus is $G = 36$
345 GPa . The analytical solution for the maximum deflection ω is,

$$\omega = \frac{3}{24} \frac{\rho g h L^4}{D} \quad (49)$$

346 where h and L are the height and length, respectively, and D is the so-called
347 flexural rigidity of the plate. The latter can be expressed in terms of the Youngs

348 modulus E and the Poisson ratio ν : $D = Eh^3/12(1 - \nu^2)$. The maximum
 349 horizontal stress in the cantilever is given by:

$$\sigma_{xx}^{max} = \frac{3gL^2}{h} \quad (50)$$

350 [Figure 4 about here.]

351 To test the mesh-dependence and the accuracy of our code we use structured
 352 meshes with different configurations of triangular elements, see Fig. .4a. We
 353 use triangles with a ratio height/length of 1 and we run the model for differ-
 354 ent numbers of elements in the vertical direction. The deformed beam and the
 355 resulting stress field of the beam with $\nu = 0.25$ are shown in Fig. .4b. The max-
 356 imum deflection of the cantilever (Fig. .4c) is well-resolved for different degrees
 357 of elastic compressibility ($0.25 \leq \nu \leq 0.4999$). Convergence to the analytical
 358 solution is achieved with only 8 elements in the vertical direction with relative
 359 errors $e_\omega < 1\%$ for all the Poisson ratios and different mesh configurations.
 360 Maximum horizontal stresses show high relative errors for coarse meshes but
 361 rapidly converge to the analytical solution with $e_{\sigma_{xx}} < 2\%$ for meshes with 10
 362 elements in the vertical direction. A good accuracy of the solver is demonstrated
 363 in both the compressible or incompressible limits. Relative errors for $\nu < 0.45$
 364 are consistent with the results obtained employing quadrilateral elements with
 365 4 nodes by Popov and Sobolev (2008) and 8 nodes by Quinteros et al. (2009).

366 5.1.2. Viscous inclusion

367 The model set-up (Fig. 4a) consists of a circular viscous inclusion with
 368 radius $R = 0.1$ embedded in a homogeneous matrix under pure shear boundary
 369 conditions in a square domain $\Omega = [-1, 1] \times [-1, 1]$. The aim of this numerical
 370 experiment is to assess the accuracy of the pressure and velocity fields in cases
 371 with strong viscosity jumps. The dimensionless viscosity of the inclusion is $\eta_1 =$
 372 10^3 and $\eta_2 = 1$ for the matrix. The domain is discretised using an unstructured
 373 mesh of triangular elements. The edges of the elements match with the interface
 374 between the inclusion and the matrix, resulting in elements belonging either to
 375 the inclusion or to the matrix. This near-perfect body-fitting mesh is the most

376 accurate way for the FEM to model this test (Deubelbeiss and Kaus, 2008).
 377 Velocity boundary conditions are imposed on the edges of the domain. These
 378 are obtained from the analytical solution for the velocity field (Schmid and
 379 Podladchikov, 2003) with a background strain rate $\dot{\epsilon}_b = 1$ (Appendix B).

380 [Figure 5 about here.]

381 Total root-mean-square (rms) errors are calculated to assess the numerical ac-
 382 curacy of this test:

$$e_p^t = \sqrt{\frac{\int_{\Omega} (P - P^{ana})^2 d\Omega}{\int_{\Omega} (P^{ana})^2 d\Omega}} \quad (51)$$

$$e_u^t = \sqrt{\frac{\int_{\Omega} (u_x - u_x^{ana})^2 + (u_z - u_z^{ana})^2 d\Omega}{\int_{\Omega} ||u^{ana}||^2 d\Omega}} \quad (52)$$

$$e_p^{t*} = \sqrt{\int_{\Omega} (P - P^{ana})^2 d\Omega} \quad (53)$$

$$e_u^{t*} = \sqrt{\int_{\Omega} (u_x - u_x^{ana})^2 + (u_z - u_z^{ana})^2 d\Omega} \quad (54)$$

386 where the superscript *ana* denotes the analytical values. Pressure errors de-
 387 crease with increasing numerical resolution (Fig. Appendix Db), with minimum
 388 values of rms error of $e_p^t = 2.3 \cdot 10^{-2}$ for high resolution meshes with $DOF \lesssim 10^5$.
 389 The velocity field is accurately calculated even for coarse meshes ($DOF = 10^3$)
 390 and shows little dependence in the number of DOF, with minimum errors of
 391 $e_u^t = 3.6 \cdot 10^{-4}$ in the finest mesh ($DOF \lesssim 10^5$). Figs. Appendix D c-d show the
 392 pressure and velocity along the horizontal plane $y = 0$ for different numerical
 393 resolutions. Coarse meshes with low number of DOFs show accurate pressure
 394 solutions in the matrix, whereas near the inclusion there is an evident drop in
 395 the accuracy of the numerical solution. High spatial resolutions ($DOF \lesssim 10^4$)
 396 lead to smoother pressure solutions around the inclusion. The velocity along the
 397 same plane displays higher levels of accuracy, with a smooth solution around
 398 the viscosity jump even for low numerical resolutions.

399 Fig. Appendix D shows the analytical and numerical solutions for pressure
400 and velocity, as well as the e_p^{t*} and e_u^{t*} distribution for the pressure and velocity
401 fields with respect to the analytical solution. As discussed above, the highest
402 pressure errors are located around the the contact between the inclusion and
403 the matrix and the minimum pressure errors are distributed along the four
404 diagonals of the domain and within the inclusion. Velocity errors are smoothly
405 distributed over the matrix and the minimum error values occur inside the
406 inclusion. The maximum numerical values of pressure and velocity show a
407 difference of 2.4450% and 0.2559%, respectively, with respect to the analytical
408 solution. These results are comparable with previous numerical benchmarks
409 (e.g. Deubelbeiss and Kaus, 2008; von Tscharner and Schmalholz, 2015).

410 [Figure 6 about here.]

411 5.1.3. Rayleigh-Taylor instability

412 The purpose of this test is to benchmark viscous deformation due to con-
413 vection driven by density contrasts (van Keken et al., 1997). The large defor-
414 mation produced in this experiment provide an excellent way to validate not
415 only the viscous deformation, but also the implementation of the remeshing
416 algorithm. Both fluids are assumed to be isoviscous with equal viscosity but
417 different density. In this test we use the dimensionless equation of conservation
418 of momentum:

$$\frac{\partial \tau_{ij}}{\partial x_j} + \frac{\partial P}{\partial x_j} = R_b \Gamma n_j \quad (55)$$

419 [Figure 7 about here.]

420 where n_j is the unit vector in the direction j and R_b is the compositional
421 "Rayleigh number" $R_b = \Delta \rho g h^3 / \kappa \eta_r$, where η_r is the reference viscosity. Γ
422 is a step function with $\Gamma = 1$ for the layer at the bottom and $\Gamma = 0$ for the top
423 layer. The domain consists of a box of height h and width λ . The thickness
424 of the bottom layer is 0.2 with an initial perturbation between the two phases
425 given by:

$$\omega = 0.02 \cos\left(\frac{\pi x}{\lambda}\right) \quad (56)$$

Table 1: Values of growth rate, maximum rms velocity and its corresponding time. $\delta\gamma$ and δu_{rms}^{max} denote the difference in % of the growth rate and maximum rms velocity with respect to the 'best' results from van Keken et al. (1997) (PvK code with 80x80 C1 elements).

| initial num. els. | DOF | γ | $\delta\gamma$ | u_{rms}^{max} | δu_{rms}^{max} | t^{max} |
|-------------------|--------|----------|----------------|-----------------|------------------------|-----------|
| 1808 | 10754 | 0.01221 | 0.28% | 0.003110 | 0.61% | 215 |
| 7093 | 42592 | 0.01222 | 0.20% | 0.003080 | 0.36% | 212 |
| 17960 | 107468 | 0.01222 | 0.20% | 0.003075 | 0.52% | 211 |

426 The aspect ratio of the domain ($\lambda = 0.9142$) is chosen such that a harmonic
427 perturbation with wavelength 2λ is the most unstable, giving the largest growth
428 rate. Displacements are restricted at the bottom and top boundaries and tan-
429 gential free-slip is allowed along the lateral boundaries (Fig. .7a).

430 We consider only an isoviscous case with $\eta_r/\eta_o = 1$ and $\rho_r/\rho_o = 1.3$.
431 Throughout the evolution of the flow we calculate the evolution with time of
432 root-mean-square velocity:

$$u_{rms}(t) = \sqrt{\frac{1}{h\lambda} \int_0^\lambda \int_0^1 ||u||^2 dx dz} \quad (57)$$

433 We use the 'best' results from van Keken et al. (1997) as a reference (Pvk code
434 with 80x80 C1 finite elements) to validate the results obtained with LaCoDe.
435 The Rayleigh-Taylor instability shows the same evolution (Fig. .7a-e) as the
436 one shown by the reference results. Only a few discrepancies are found in the
437 geometry of the secondary and tertiary diapirs in the late stages of the flow
438 evolution. Models with coarse meshes are able to predict accurate values of the
439 maximum rms velocity, but predict maximum rms velocities for the secondary
440 diapir that are 13% higher than the values obtained with a finer mesh (Fig. .7f).
441 The growth rate of the instability γ at $t = 0$ and the maximum rms velocity
442 (Table 5.1.3) are in agreement with the reference values, with errors smaller
443 than 1%. The increase in the difference of the maximum u_{rms} for the case with
444 17960 elements is due to a numerical resolution 2.8 times higher than the one
445 employed in the reference case, presumably leading to a more accurate solution.

446 The remeshing algorithm is called when the quality of any element or ele-

447 ments of the mesh is below the quality threshold. The two fluids are discretised
448 in the space such that their interface represents a sharp contact, with individual
449 elements belonging to a single phase. The interface is tracked with time, and it
450 is used to define the geometry of the new mesh. The interface between the two
451 fluids undergoes a high amount of stretching during the evolution of the flow
452 and it is refined during the remeshing so that its spatial resolution is constant
453 (i.e. a new node is added to the interface if the space between two consecutive
454 nodes is larger than a specified distance), producing a considerable increase of
455 the number of elements in later stages.

456 In this numerical experiment it is sufficient to generate a new mesh and
457 there is no actual need to transfer information from the new mesh into the
458 new one. However for benchmarking purposes, we perform the mapping of the
459 second invariant of the accumulated strain E_{II} onto the new high quality mesh.
460 Fig. .7g,h shows an accurate mapping of E_{II} from the old mesh onto the new
461 mesh. The quality of the remeshing algorithm is assessed by comparing the
462 finite strain field before and after remeshing. In order to compare the pre- and
463 post-remeshing results, both fields are sampled in high-resolution rectangular
464 grid of 1000 by 1000 nodal points, where the root-mean-square error of the
465 mapped field is computed (Fig. .7i).

466 5.1.4. Stress build up in a visco-elastic Maxwell body

467 Visco-elastic deformation is demonstrated by repeating the numerical ex-
468 periment of build-up of stress in a Maxwell body under pure shear deformation
469 (Gerya and Yuen, 2007). A constant background strain rate $\dot{\epsilon} = 10^{-15} \text{ s}^{-1}$ is
470 prescribed at the boundaries of a body with a 100 by 100 km domain (Fig. .8a).
471 The mechanical parameters are: $G = 10 \text{ GPa}$, $\eta = 10^{22} \text{ Pa} \cdot \text{s}$ and gravity is
472 switched off. We take $\nu = 0.4999$ in order to approximate an incompressible
473 material. The build-up of the stress is described by the following analytical
474 expression:

$$\tau = 2\dot{\epsilon}_{II}\left(1 - \exp\left(-\frac{Gt}{\eta}\right)\right) \quad (58)$$

475 The analytical and numerical time-stress curves overlap (Fig. .8b,c), demon-
476 strating the high accuracy of the implementation of the Maxwell rheology.

477 [Figure 8 about here.]

478 *5.1.5. Couette flow with viscous heating and temperature dependent viscosity*

479 [Figure 9 about here.]

480 The aim of this test is to demonstrate the accuracy of the numerical solution of
481 thermal diffusion and the coupling of the Stokes equations with the conservation
482 of energy for fluids with temperature-dependent viscosity and shear heating.
483 The set-up of the model is consists of the Couette flow in a rectangular channel
484 (Fig..9a). The motion of the flow is driven by shear along the top boundary
485 of the channel with the following boundary conditions: no-slip motion ($u(z =$
486 $0) = 0$) and constant temperature ($T(z = 0) = T_0$) at the lower boundary, zero
487 vertical pressure gradient ($\partial P/\partial z = 0$), constant shear stress and $\partial T/\partial x = 0$ at
488 the lateral boundaries of the model. The size of the model is $\Omega = [0, 90] \times [0, 12]$
489 km. This length-to-depth ratio is sufficiently large to avoid errors in the flow
490 due to boundary effects. The model is started with T_0 across the whole domain.
491 The analytical solution of this problem is described in the Appendix Appendix
492 D.

493 The dependence of the maximum non-dimensional temperature change in
494 the channel θ with the Brinkman number Br is used to compare the analytical
495 solution with the numerical results, taking values of $E_a = 150$ J/mol, $R =$
496 8.35 , $A = 10^{15}$ Pa·s, $K = 2$ W/m/K and $T_0 = 1000$ K. The results obtained
497 with LaCoDe show an excellent agreement with the analytical solution (Fig..9),
498 demonstrating the capability of the code to model coupled thermo-mechanical
499 problems with non-linear rheologies and shear heating.

500 *5.2. Non-linear rheology iterations: single vs nested Picard iterations*

501 We test the accuracy and efficiency of these two solution schemes with two
502 different numerical experiments: A) a visco-elastic rectangular body under pure

503 shear with a non-Newtonian rheology including diffusion and dislocation creep;
504 and, B) a set-up for a subduction problem with a non-Newtonian visco-elastic
505 rheology. In both problems, we keep track and compare the number of linear and
506 non-linear iterations, residual velocity and computational time during the first
507 five time steps for Test A, and six time steps for the Test B (this corresponds
508 to the number of time steps before remeshing is required). Details of the model
509 set-up, boundary conditions and thermo-mechanical parameters are found in
510 Appendix Appendix A.

511 [Figure 10 about here.]

512 Results from Test A (Fig. .10a) show that, as expected, Approach 2 leads to
513 a higher number of Powell-Hestenes iterations compared to dealing with all non-
514 linearities in the same loop as in Approach 1, resulting in typically ~ 1.5 times
515 more linear iterations $\sim 25\%$ more computational time per iteration. Despite
516 being somewhat more expensive, Approach 2 yields a better-converged solution.

517 The efficiency of Approach 1 and 2 is further checked with the more realistic
518 Test B, where a rheologically layered domain adds new degrees of complexity
519 to the problem. In this case we have capped the maximum number of the outer
520 level of Picard iterations to 60. Approach 2 converges typically within 17-30
521 outer Picard iterations, whereas Approach 1 constantly reaches the maximum
522 allowed number of iterations and results in a poorly-converged solution (Fig.
523 .10b). In this case, every time step using Approach 2 needs to perform about 2
524 or 3 times the number of linear iterations performed by Approach 1; however,
525 approximately half of the rheological non-linear iterations are required, yielding
526 a slightly cheaper solution scheme.

527 Considering these results, we infer that treating all the non-linearities in one
528 level of Picard iterations (Approach 1) is more efficient in terms of total number
529 of iterations; however, this approach yields larger residuals of the velocity field
530 (Fig. .10). Approach 2 also becomes substantially cheaper than Approach 1 as
531 the complexity of the problem increases because a lower number of outer Picard
532 iterations is required. We therefore recommend to use the solution scheme as

533 in Approach 2 for complex and highly non-linear problems.

534 5.3. Numerical experiments with a compressible crust and mantle

535 5.3.1. Volumetric strain induced by serpentinization

536 The phase change from peridotite to serpentinite is accompanied by a con-
537 siderable reduction in density. In this experiment, we simulate a visco-elastic
538 oceanic lithosphere in which serpentinization occurs to different degrees. The
539 transformation of mantle peridotites to serpentinite occurs within a specific
540 range of pressure and temperature and with an inflow of sea water into the ma-
541 terial. However, in the model shown here, we simplify this process by imposing
542 a rate of density change in a target region, at a rate that reaches the maximum
543 degree of serpentinization after 1 Myr. This experiment is designed to explore
544 the impact of the sudden reduction of density and change of volume on the
545 stress and strain fields.

546 The model is 300 km long by 100 km deep and is stretched under pure shear
547 boundary conditions, with a full extension rate of $u_{ext} = 1$ mm/yr. Serpentiniza-
548 tion occurs within the 40 km by 10 km rectangular area located at the centre
549 of the model. The rheology is visco-elastic with $\eta = 10^{23}$ Pa s, $G = 36GPa$
550 and $\nu = 0.3$. The density of the serpentinized material is calculated as a linear
551 function of β (Escartin et al., 2001):

$$\rho(\beta) = \rho_{serp} \left(1 - \frac{\beta}{100} \right) (\rho_o - \rho_{serp}) \quad (59)$$

552 where β is the percent of serpentinization. We take a $\rho_o = 3300$ kg/m³ char-
553 acteristic of mantle material and $\rho_{serp} = 2550$ kg/m³. We run a set of models
554 with different values of degree of serpentinization ($\beta = 0, 20$ and 40%).

555 [Figure 11 about here.]

556 It is known that at these values of serpentinization, significant weakening
557 of the lithosphere might occur (Escartin et al., 1997). Considering a pressure
558 dependent failure criterion such as Drucker-Prager, $\tau_y = p \sin(\phi) + C \cos(\phi)$,
559 and assuming a friction angle $\phi = 30^\circ$ and cohesion $C = 30$ MPa (dashed

560 line in Fig..11b), it becomes evident that the stress linked to the volumetric
561 increased caused by serpentinization reactions can easily exceed the yield stress
562 at shallow depths (at ~ 2 km for $\beta = 20\%$ and ~ 10 km for $\beta = 20\%$; Fig..11b),
563 thus localising, or enhancing, inelastic deformation in faults and shear bands.
564 Topographic expressions in the sea-floor could also be linked to the production of
565 serpentinite at shallow depths (Fig..11c). Our models predict topographic highs
566 from 0.3 km and 0.7 km for a partially serpentinized material for $\beta = 20\%$ and
567 $\beta = 40\%$, respectively.

568 For comparison, we include a model with $\beta = 40\%$ using the incompressible
569 Boussinesq approximation (i.e. the continuity equation is approximated as $\nabla \cdot$
570 $\mathbf{u} = 0$). The incompressible approximation is not able to resolve the volumetric
571 strains and the flow solution only accounts for the buoyancy forces produced by
572 the serpentinization. Therefore, the strain field is barely affected by the phase
573 change and the stress field is incorrect, showing even lower stresses than for
574 $\beta = 0\%$ (Fig..11b). Furthermore, the pressure dependence of the density in this
575 model is switched off or it would become unstable after few time steps.

576 Even though the model considered here is very simple, and more realistic set-
577 ups and conditions might change the values of the effect of serpentinization (e.g.
578 plastic deformation, rheological layering, etc.), it serves as an example for how
579 the volumetric strain produced by a phase change can potentially weaken the
580 crust and localise brittle deformation. Therefore, weakening by serpentinization
581 may play a crucial role to shape the kinematics of magma-poor margins and the
582 bending/unbending of subducting plates (Phipps Morgan, 2001). This numer-
583 ical example also shows that the incompressible Boussinesq approximation is
584 not able to deal with large density changes and predicts unrealistic strain and
585 stress fields. Instead, a compressible formulation should be used.

586 *5.3.2. Subduction of a compressible slab*

587 [Figure 12 about here.]

588 In subduction zones, the cold subducting plate is rapidly buried to great depths.
589 Hence the subducting slab is subject to considerable pressure changes that imply

590 large variations of the density. In this test, we investigate how large these density
591 variations can be for a compressible mantle and lithosphere, and whether they
592 eventually become large enough ($> 10\%$) so that the Boussinesq approximation
593 becomes inaccurate. We employ a non-Newtonian visco-elastic rheology and
594 the mechanical parameters, set-up and boundary conditions for subduction are
595 described in Appendix A.2. The thermal ages of the oceanic and continental
596 lithospheres are 70 Ma and 400 Ma, respectively. For completeness, we compare
597 results of obtained with compressible ($\nu = 0.30$) and incompressible mantle-
598 lithosphere. In the latter, incompressibility is approximated by using a Poisson
599 ratio of $\nu = 0.4999$. In the compressible case, ridge push boundary conditions
600 are applied until 4 Ma. At this moment, the tip of the slab is dense enough for
601 slab-pull to become effective, and no additional forces are required to sustain
602 the subduction of the oceanic lithosphere. The density in the incompressible
603 case is lower, and ridge push boundary conditions need to be prescribed until 5
604 Ma.

605 At 3.5 Ma, while ridge push is still active, the compressible oceanic litho-
606 sphere has subducted 297 km and the dip at its tip is 60° (Fig. .12a). After
607 slab-pull becomes effective, the trench starts to retreat and the slab rolls-back.
608 At 7.1 Ma, the pressure at the tip of the slab is high enough to produce density
609 variations with respect to the reference state that exceed the accuracy thresh-
610 old of the Boussinesq approximation (Fig. .12a). At this point the trench has
611 retreated 114 km, the slab is 14° steeper, and has further subducted down to
612 477 km depth (Fig. .12a).

613 In the incompressible case, the oceanic lithosphere has subducted to a depth
614 comparable to the compressible case. However, the dip of the incompressible slab
615 is 10° less. Furthermore, at 7.1 Ma the incompressible slab will subduct only
616 another 73 km (even if ridge push lasts an additional million year), whereas the
617 compressible slab subducts extra 180 km with respect to the depth at 3.5 Ma.

618 This simple numerical experiments illustrates how compressibility is a me-
619 chanical feature that is certainly important to account for in models of subduct-
620 ing slabs. The enormous pressures that build up at the tip of the slab lead to

621 density variations of more than 10% that affects the timing and effectiveness of
622 slab pull, and the dynamics of subduction.

623 **6. Discussion and summary**

- 624 1. An implicit approach of the general compressible Stokes equation can be
625 well resolved using iterative solvers such as the Augmented Lagrangian
626 Method.
- 627 2. The dependency on density of the compressible continuity equation intro-
628 duces an additional non-linearity into the problem, with respect to the
629 incompressible approximation, thus increasing the total number of iter-
630 ations per time step. We find that for non-Newtonian rheologies, one
631 could treat all the non-linearities within one Picard loop. However, as
632 the complexity of the problem increases, it becomes convenient to split
633 the non-linearities with a rheological nature from the ones raising from
634 the continuity equation into two levels of Picard iterations, as it leads to
635 faster convergence rates and better resolved solutions. Even if not consid-
636 ered in this paper, the latter scheme holds if other non-linear rheological
637 features are incorporated in the model, such as plastic deformation.
- 638 3. While the Boussinesq approximation is a valid hypothesis for simple mod-
639 eling of crustal deformation, more complex models that aim to study pro-
640 cesses such as phase changes or subduction of oceanic lithosphere will
641 require a modification of the Boussinesq approximation to accommodate
642 the effects of volumetric strains and volume-change-linked stresses.
- 643 4. Benchmarks for elastic deformation and stresses show that the formulation
644 presented here is able to model elasticity both for compressible materials
645 and in the incompressible limit.
- 646 5. The accuracy of LaCoDe for viscous deformation has been demonstrated.
647 The velocity and pressure fields from the viscous inclusion test are consis-
648 tent the analytical solutions. The benchmarks of compositional convection
649 is also in agreement with previous benchmarks.

- 650 6. The agreement of the numerical and analytical solution of a Couette flow
651 with viscous heating and temperature dependent viscosity demonstrates
652 the accuracy of LaCoDe to solve thermo-mechanical problems.
- 653 7. The inclusion of a self-consistent volume change source term is a powerful
654 tool that opens an opportunity to study the effects of overpressure caused
655 by the inflow and outflow of mass into geological features (e.g. serpen-
656 tinization and melt extraction). Exploring these processes will be the goal
657 of future work.

658 **Appendix A. Model set-up and boundary conditions for tests in Sec-** 659 **tion 3.4**

660 *Appendix A.1. Test A: Pure shear deformation of a non-Newtonian visco-elastic* 661 *body*

662 The initial size of the models is a 500 km by 400 km rectangular box with
663 an initial temperature profile as shown in (Fig. .13a). We use a non-Newtonian
664 visco-elastic with the thermo-mechanical parameters of wet olivine (Table Ap-
665 pendix A.2). Pure shear far-field boundary conditions are prescribed in the
666 boundaries of the model (i.e. half and full extension rate are prescribed at the
667 lateral and bottom boundaries of the domain, respectively), the boundaries of
668 the model are thermally insulated and tangential free slip condition are pre-
669 scribed at the lateral and bottom boundaries. Temperature is fixed at 0 °C and
670 1300 °C at the surface and bottom of the model. A free-surface algorithm is em-
671 ployed to calculate the dynamic response of the topography (Andrés-Martínez
672 et al., 2015). The domain of the model is discretised by an unstructured mesh
673 of 13828 triangular elements (42271 DOFs).

674 *Appendix A.2. Test B: Subduction initiation*

675 The set-up of Test B correspond to a subduction problem with a size of 3000
676 km by 1500 km. The oceanic and continental lithosphere are 80 km and 140 km
677 thick, respectively. The motion of the bottom and lateral sides is fixed, and con-
678 vergence boundary velocity conditions are prescribed in a vertical profile along

679 the oceanic lithosphere 500 km before the trench. We use a non-Newtonian
680 visco-elastic rheology with a wet quartzic crust, dry olivine continental litho-
681 sphere and wet olivine for the oceanic lithosphere and asthenosphere. All the
682 boundaries except the surface are thermally insulating; bottom and top tem-
683 peratures are constant at 0 °C and 1300 °C at the surface; and free surface
684 boundary conditions are prescribed at the top of the model. The initial thermal
685 structure is given by continental lithosphere with a thermal age of 500 Ma and
686 an oceanic lithosphere with a thermal age of 75 Ma. To ease the subduction
687 initiation, we introduce a weak layer between the oceanic and continental litho-
688 spheres with a constant viscosity of $5 \cdot 10^{19}$ Pa.s. The domain of the model is
689 discretised by an unstructured mesh of 17927 triangular elements (55107 DOFs).

690 [Figure 13 about here.]

Table A.2: Rheological parameters. Wet quartzite from Gleason and Tullis (1995) and dry olivine and wet olivine from Hirth and Kohlstedt (2003), respectively.

| Parameter | Units | Wet Olivine | Dry Olivine | Wet Quartzite |
|------------------|-----------------|-------------------|-------------------|---------------------|
| c | MPa | 20 | 20 | 20 |
| ρ | kgm^{-3} | 3300 | 3300 | 2850 |
| G | GPa | 74 | 74 | 36 |
| α | - | $3 \cdot 10^{-5}$ | $3 \cdot 10^{-5}$ | $2.4 \cdot 10^{-5}$ |
| H_Q | Wm^{-3} | 0 | 0 | $0.2 \cdot 10^6$ |
| K | $Wm^{-3}K^{-3}$ | 3.3 | 3.3 | 2.5 |
| $\log_{10}(A)$ | $Pa^{-n}s^{-1}$ | -15.56 | -15.56 | -28 |
| E | $KJmol^{-3}$ | 480 | 530 | 223 |
| $\log_{10}(V_o)$ | m^3mol^{-3} | -6 | -6 | 1 |
| n_{dis} | - | 3.5 | 3.5 | 4 |
| n_{dif} | - | 1 | 1 | 0 |

691 **Appendix B. Analytical solution for a thin beam under uniform load**

692 The general equation describing the deflection ω of an elastic cantilever of
 693 length L and thickness h is given by:

$$D \frac{d^4 \omega}{dx^4} = q(x) - p \frac{d^2 \omega}{dx^2} \quad (\text{B.1})$$

694 where $q(x)$ is the load and p is the pressure. Considering $p = 0$ and a constant
 695 and uniform load, eq. (B.1) yields:

$$\frac{d^4 \omega}{dx^4} = \frac{q}{D} \quad (\text{B.2})$$

696 Eq. (B.2) can be integrated using the following boundary conditions: 1) $\omega = 0$
 697 at $x = 0$ (fixed end); 2) $d\omega/dx = 0$ at $x = 0$; 3) $d^2\omega/dx^2 = 0$ at $x = L$; and,
 698 4) $dM/dx = V$, where M is the bending momentum and V is the shear force.
 699 After some algebra, the solution can be written as:

$$\omega = \frac{qx^2}{D} \left(\frac{x^2}{24} + \frac{Lx}{6} + \frac{L^2}{4} \right) \quad (\text{B.3})$$

700 with the q being the gravitational load $q = g\rho Lh$. The horizontal stress along
 701 the cantilever is given by the expression:

$$\sigma_{xx} = \frac{E}{1 - \nu^2} \varepsilon_{xx} \quad (\text{B.4})$$

702 the horizontal strain is given by:

$$\varepsilon_{xx} = -z \frac{d^2 \omega}{dx^2} \quad (\text{B.5})$$

703 and the bending momentum at $x = 0$ is:

$$M = -\frac{qL^2}{h} \quad (\text{B.6})$$

704 The maximum bending stress at $x = 0$ in a cantilever, centred at $z = 0$, occurs
 705 at $z = \pm h/2$ and it is obtained combining eqs. (B.4), (B.5) and (B.6):

$$\sigma_{xx}^{max} = \frac{3qL^2}{h^2} \quad (\text{B.7})$$

706 **Appendix C. Analytical solution for a viscous inclusion**

707 The analytical solution of a viscous inclusion within a homogeneous matrix is
 708 based on Muskhelishvili's complex variable stress-function method and solution
 709 (Muskhelishvili, 1953) for 2D elasticity. Here we present a brief description
 710 with the solution under pure shear conditions. A more detailed description
 711 in the geological literature is found in Schmid and Podladchikov (2003). The
 712 coordinates are expressed in the complex plane:

$$z = x + iy \quad (\text{C.1})$$

713 where $i = \sqrt{-1}$. For a slow incompressible viscous flow in plane strain, the
 714 velocity field can be expressed in terms of the complex functions $\phi(z)$ and $\psi(z)$:

$$u_x + iu_z = \frac{\phi(z) - z\overline{\phi'(z)} - \overline{\psi(z)}}{2\eta} \quad (\text{C.2})$$

715 where the overbar refers to the complex conjugate and the prime refers to the
 716 derivative with respect to z . Under pure shear boundary conditions the func-
 717 tions $\phi(z)$ and $\psi(z)$ in the matrix are given by:

$$\phi_m(z) = -\frac{2\dot{\epsilon}Ar_c^2}{z} \quad (\text{C.3})$$

719

$$\psi_m(z) = -2\dot{\epsilon}\eta_m z - \frac{2\dot{\epsilon}Ar_c^4}{z^3} \quad (\text{C.4})$$

720 with

$$A = \frac{\eta_m(\eta_c - \eta_m)}{\eta_c + \eta_m} \quad (\text{C.5})$$

721 where r_c is the radius of the inclusion and η_m and η_c are the viscosities of the
 722 matrix and the inclusion, respectively. Inside the inclusion:

$$\phi_c(z) = 0 \quad (\text{C.6})$$

723

$$\psi_c(z) = -4\dot{\epsilon}\frac{\eta_c\eta_m}{\eta_c + \eta_m}z \quad (\text{C.7})$$

724 Substitution of eqs. (C.3) and (C.4) into (C.2) yields the analytical solution for
 725 the velocity field in the matrix:

$$u_x + iu_z = \frac{\dot{\epsilon}Ar_c^2}{\eta_m} \left[-\frac{1}{z} + \frac{z}{z^2} - \frac{1}{z^3} - \frac{\bar{z}\eta_m}{Ar_c^2} \right] \quad (\text{C.8})$$

726 Substitution of (C.6) and (C.7) into (C.2) give the analytical solution for the
 727 velocity inside the inclusion:

$$u_x + iu_z = -\frac{4\dot{\epsilon}}{2\eta_c} \frac{\eta_c\eta_m}{\eta_c + \eta_m} \bar{z} \quad (\text{C.9})$$

728 The general expression of the pressure field is given by:

$$p = -2Re(\phi'(z)) \quad (\text{C.10})$$

729 with $Re(\cdot)$ denoting the real part of (\cdot) . Under pure shear boundary conditions
 730 the pressure field in the inclusion is $p_c = 0$ and the pressure in the matrix is
 731 given by:

$$p_m = -2Re\left(\frac{2\dot{\epsilon}Ar_c^2}{z^2}\right) \quad (\text{C.11})$$

732 **Appendix D. Analytical solution for a Couette flow with viscous heat-** 733 **ing and temperature dependent viscosity**

734 The non-Newtonian viscosity of the flow is controlled by the following equa-
 735 tion (Turcotte and Schubert, 2014):

$$\eta = A \exp\left[\frac{E_a}{RT_0} \left(1 - \frac{T - T_0}{T_0}\right)\right] \quad (\text{D.1})$$

736 where E_a is the activation energy, R is the gas constant and A is a pre-
 737 exponential factor that depends on the material. The analytical solution of
 738 the temperature field of the flow is described by the following set of equations
 739 (Turcotte and Schubert, 2014):

$$x = \frac{L}{B} \ln \left[\frac{(D - B)(C - B)}{(D - B)(C + B)} \right] \quad (\text{D.2})$$

740

$$B = \ln \left[\frac{1 + \left(1 - \frac{2Br}{B^2}\right)^2}{1 + \left(1 + \frac{2Br}{B^2}\right)^2} \right] \quad (\text{D.3})$$

741

$$C = \sqrt{2(\phi_1 - \phi(x))Br} \quad (\text{D.4})$$

742

$$D = \sqrt{2(\phi_1 - 1)Br} \quad (\text{D.5})$$

743

$$\phi(x) = \exp(\theta(x)) \quad (\text{D.6})$$

744

$$\theta(x) = \frac{E_a T(x) - T_0}{RT_0^2} \quad (\text{D.7})$$

745

$$\phi_1 = \frac{B^2}{2Br} = \exp(\theta_1) \quad (\text{D.8})$$

746

$$\theta_1 = \frac{E_a(T_1 - T_0)}{RT_0^2} \quad (\text{D.9})$$

747

$$Br = \frac{(\sigma_{xz1}L)^2 E_a}{KART_0^2} \exp\left(-\frac{E_a}{RT_0}\right) \quad (\text{D.10})$$

748 where Br is the non-dimensional Brinkman number, θ is the non-dimensional
 749 temperature change, σ_{xz1} is the shear stress at the top boundary, K is the
 750 thermal conductivity and T_1 is the temperature at the top boundary. If non-
 751 negative values of B are chosen, the Brinkman number can be calculated as
 752 (Gerya, 2009):

$$Br = \frac{B^2}{2} \left[1 - \left(\frac{\exp(B) - 1}{\exp(B) + 1} \right) \right] \quad (\text{D.11})$$

753 For a given σ_{xz} the solution is non-unique and two flows with different temper-
 754 ature and velocity exist. However, a unique solution exists if a given velocity is
 755 prescribed at the upper boundary. Therefore, we prescribe a constant horizon-
 756 tal velocity boundary u^* at the upper boundary instead of imposing a constant
 757 shear stress.

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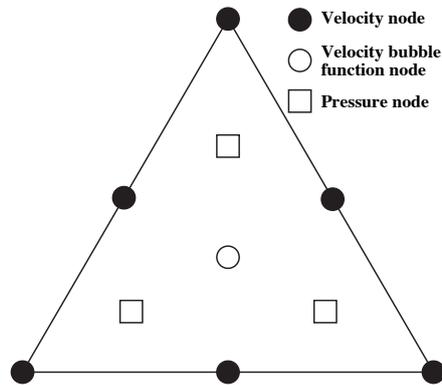


Figure .1: Crouzeix-Raviart triangular element. These elements are characterised by continuous quadratic velocities with cubic bubble function in the baricenter of the triangle and discontinuous linear pressure and show quadratic convergence.

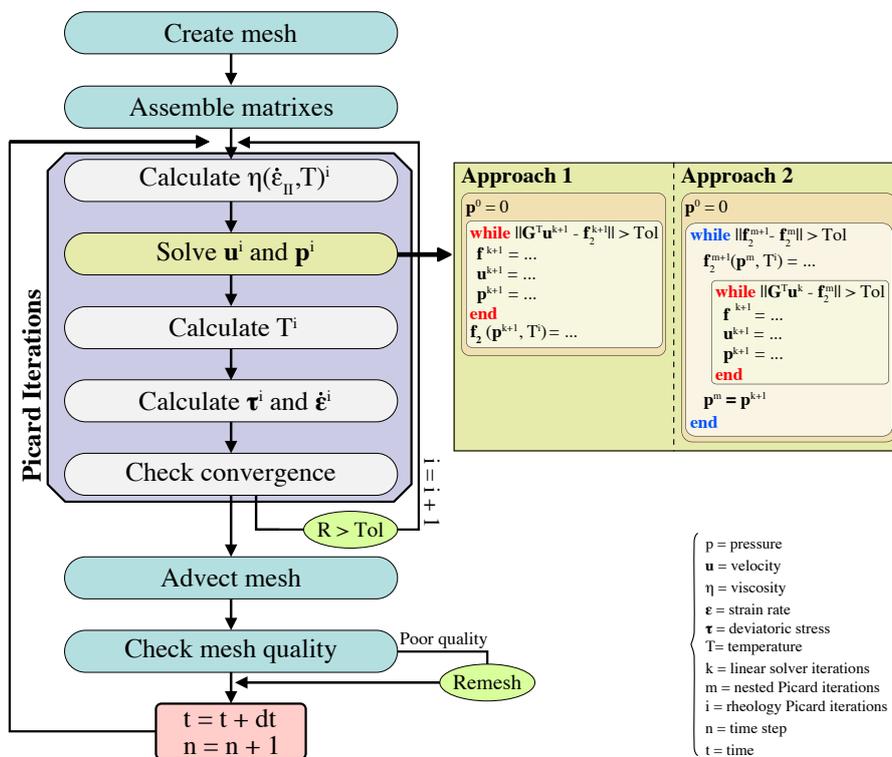


Figure .2: Global work flow of the code.

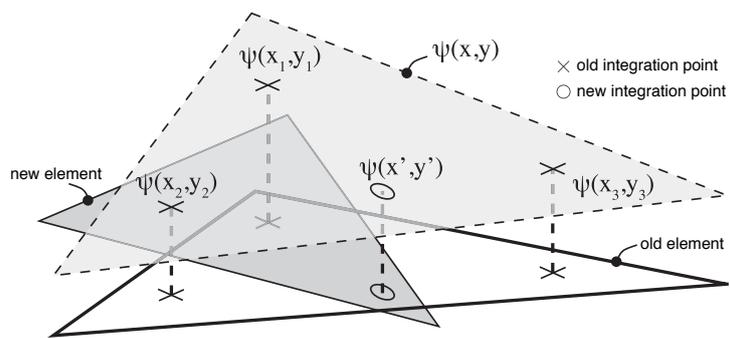


Figure .3: The information stored at the integration points of the elements of the old mesh is mapped into the new elements using the shape functions as interpolation functions. For simplicity, the field $\Psi(x, y)$ depicted in this sketch is assumed to be linear.

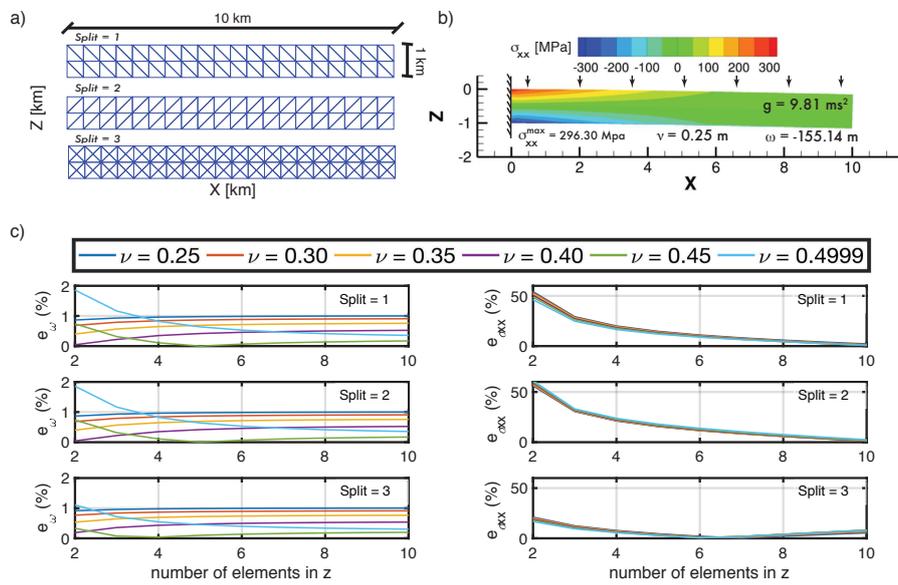


Figure 4: **a)** Structured triangular meshes with different element orientations. **b)** Set-up for the cantilever problem and flexure and stress field after loading for $\nu = 0.25$. **c)** Relative errors of the maximum deflection and bending stress for a thin beam embedded in one side and subjected to a uniform loading.

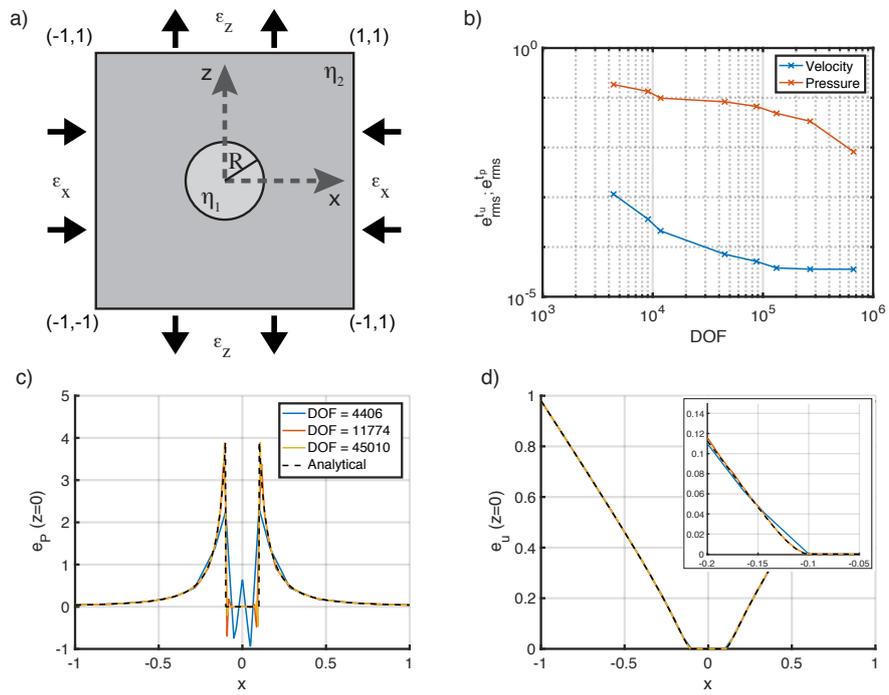


Figure 5: **a)** Set-up of a viscous inclusion with radius R and viscosity η_1 within a homogeneous matrix with viscosity η_2 and equal height and width. **b)** The domain is discretised with triangular elements so that the edges of the elements near-perfectly fit the boundary between the inclusion and the matrix. Comparison of the analytical solution for **c)** pressure and **d)** velocity along the plane $z = 0$ for numerical solutions with increasing numbers of DOF. The inset in **d)** shows the smooth transition in the velocity field with increasing number of DOF.

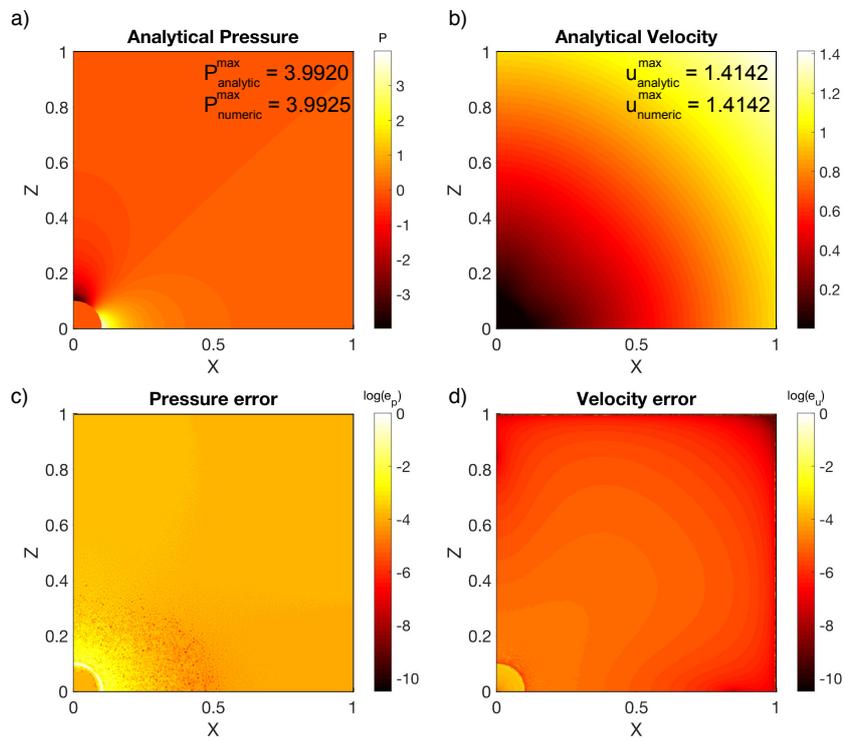


Figure .6: Numerical solutions of the **a)** pressure and **b)** velocity fields; and distribution of the logarithmic rms error of **c)** pressure and **d)** velocity. The zoom-in in **d)** shows the zero velocity error in the boundaries of the domain. Due to the symmetry of the pressure and velocity fields, only the upper-right corner of the domain ($\Omega' = [0, 1]x[0, 1]$) is shown in this figure. The results shown here correspond to a mesh with $6.65 \cdot 10^5$ DOF.

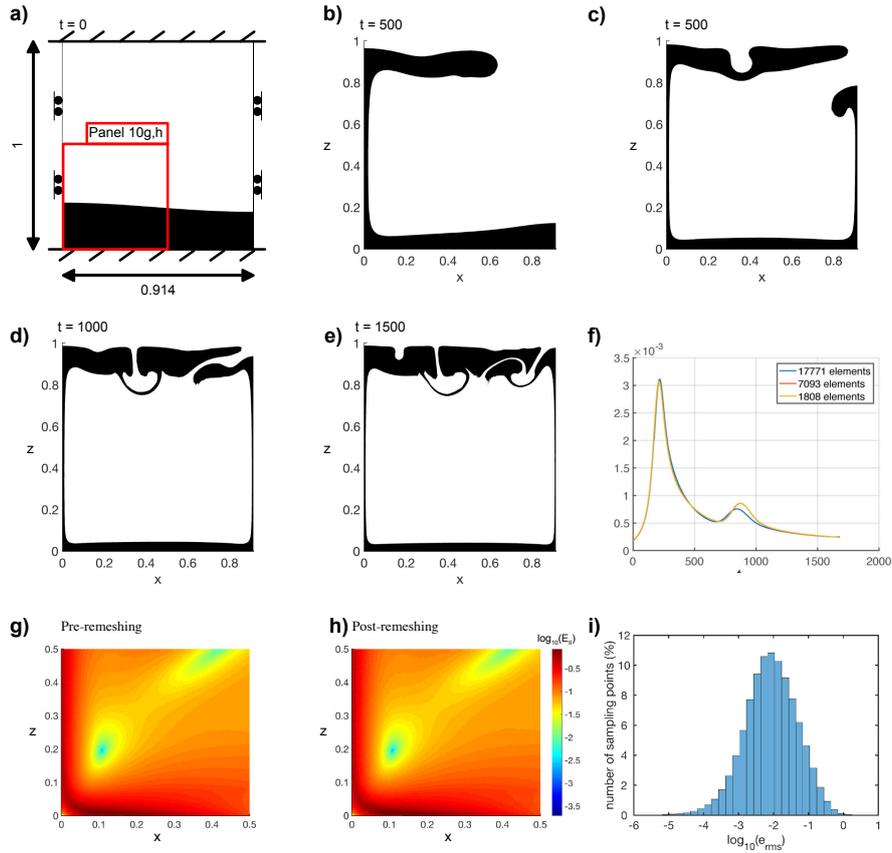


Figure .7: **a-e)** Temporal evolution of the Rayleigh-Taylor instability. **g)** Evolution of u_{rms} . Remeshing of the domain is necessary when the mesh becomes highly distorted. Note that the red lines is overlapped by the blue line. **g-h)** Comparison between the second invariant of strain E_{II} field in a mesh with heavily distorted elements and the accumulated square root of second invariant of the strain rate interpolated into a new mesh. **i)** Histogram showing the logarithm of the error between the accumulated square root of second invariant of the strain rate, pre and post remeshing

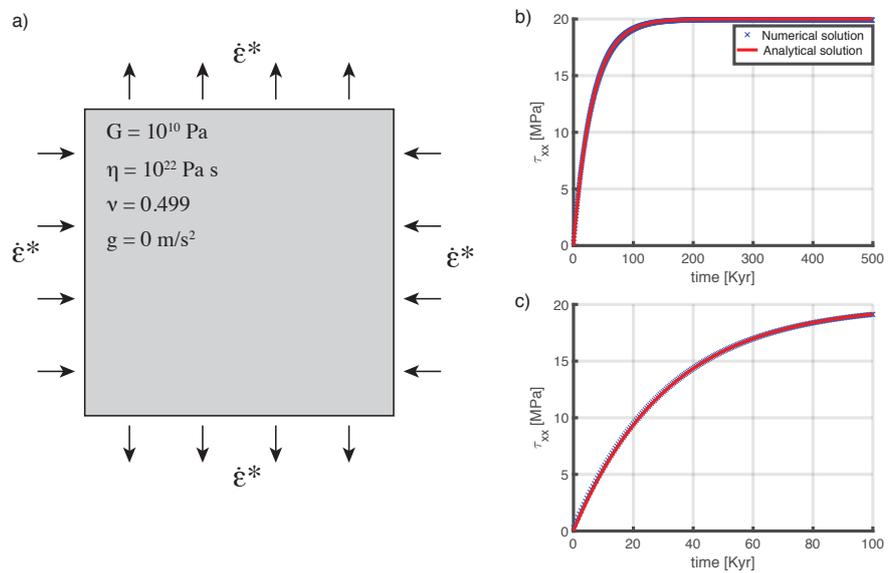


Figure .8: **a)** Set-up for the stress build-up experiment: a rectangular body is deformed with a constant background strain rate under pure shear boundary conditions. **b)** Comparison of the stress between the analytical solution and the numerical results. **c)** Zoom in the stress-time curve in the visco-elastic regime.

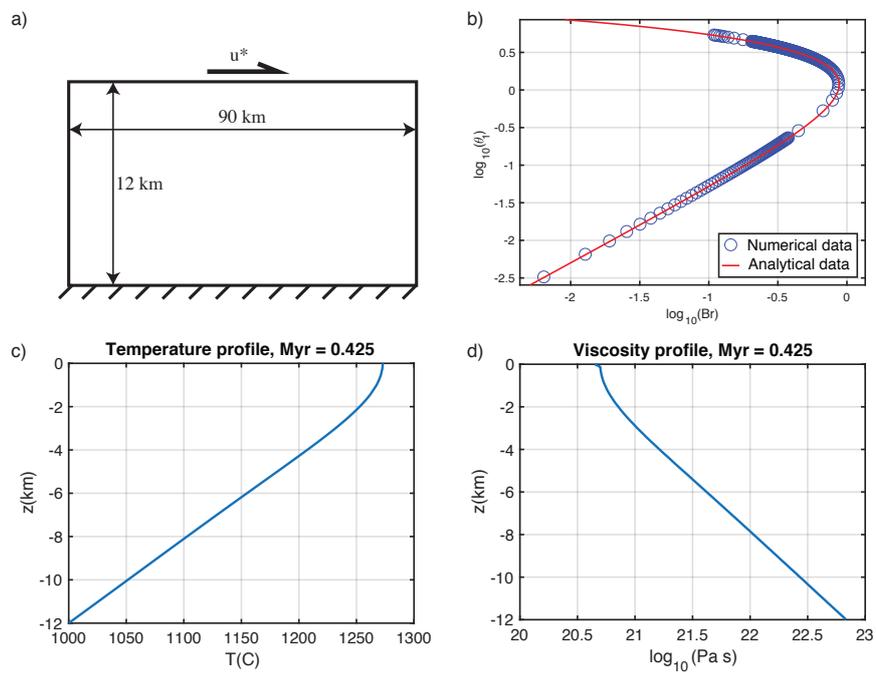
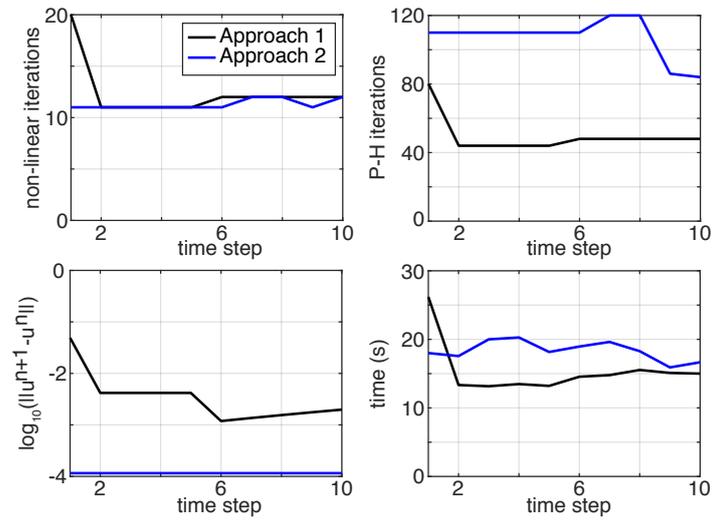


Figure .9: **a)** Set-up for Couette flow: the velocity at the bottom is $u=0$ and constant velocity u^* is prescribed at the top boundary. **b)** Analytical and numerical relationship between the Brinkman number and the non-dimensional temperature at the top of the Couette flow. Vertical **c)** temperature and viscosity **d)** profiles after 0.425 Myrs.

a) Test 1: Non-newtonian body under pure shear



b) Test 2: Subduction problem

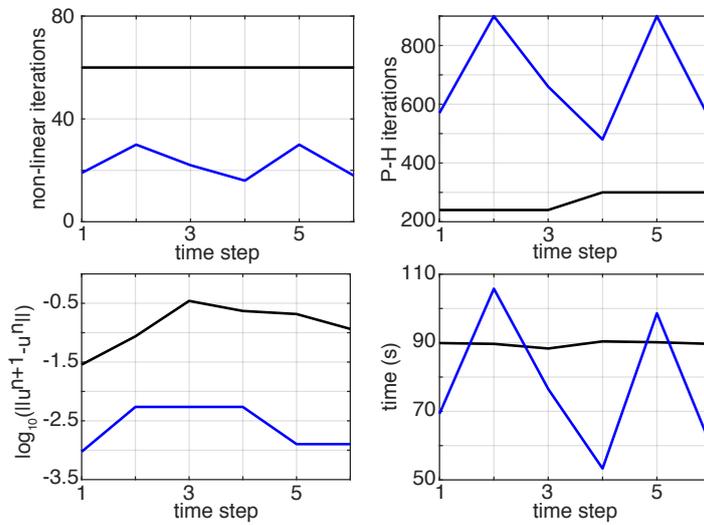


Figure .10: Comparison of the number of non-linear and linear iterations, residual velocity and computational time between Approach 1 and Approach 2 for a) Test A and b) Test B. The average computational times per time iteration for Test A are 15.43 s for Approach 1 and 18.34 s for Approach 2, whereas Test B yields average times of 89.68 s and 77.20 s for Approach 1 and 2, respectively.

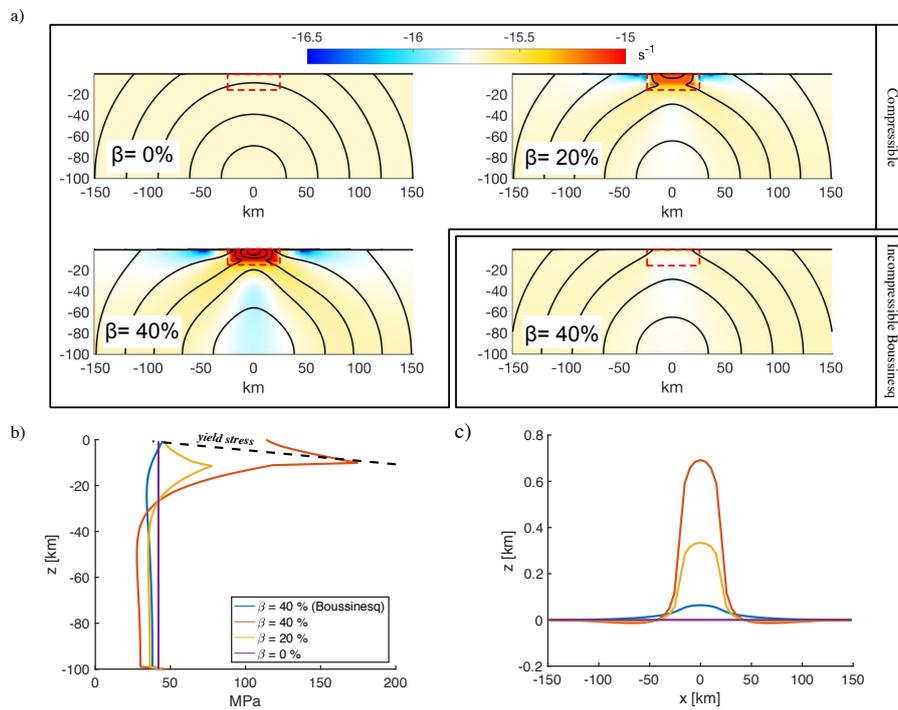


Figure .11: **a)** Results for different values of β . The density depends linearly on the degree of serpentinization: $\beta = 0, 20$ and 40 kg/m^3 . The color maps represent the square root of the second invariant of the stress and the thick black lines are isolines of the velocity field. The change of density occurs within area delimited by the dashed red rectangle. **b)** Vertical profile of τ_{II} at $x = 0$; the dashed line represents the yield stress given by a pressure dependent yield surface: $\tau_y = p \sin \phi + C \cos \phi$. **c)** Comparison of the topographic relief for different degrees of serpentinization. All the results shown here correspond to $t = 1 \text{ Myr}$.

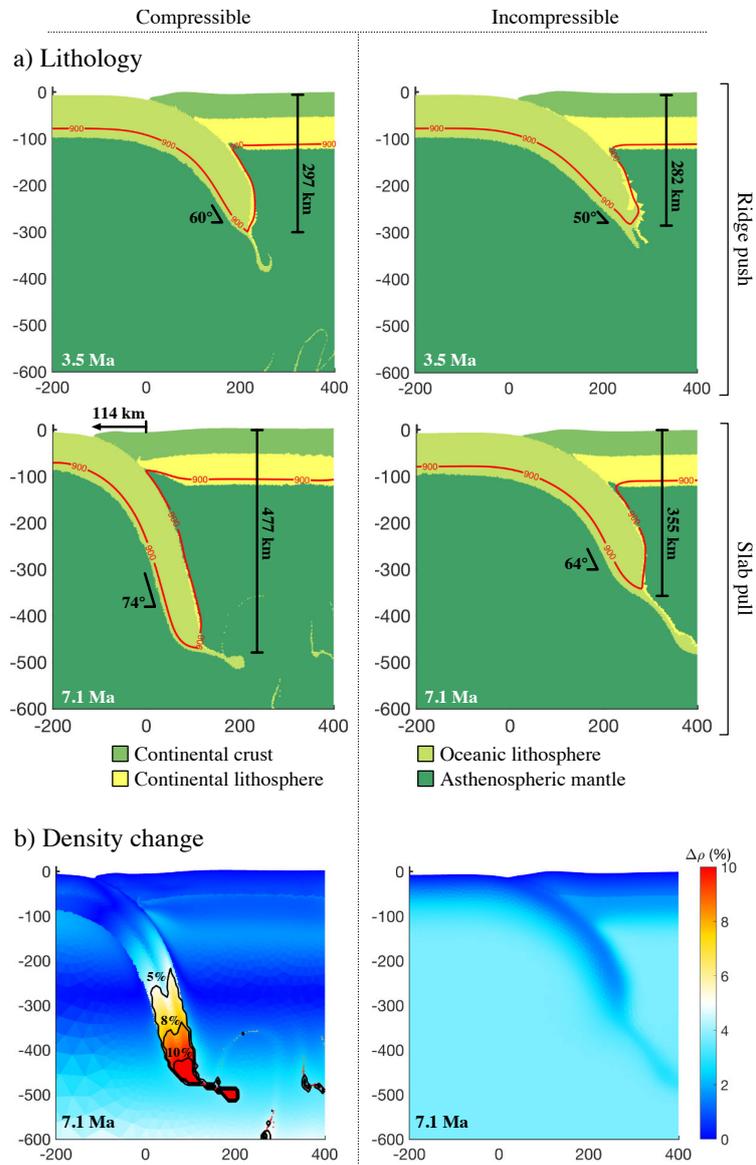


Figure .12: **a)** Snapshots of the subducting slab at 4.0 and 7.1 Ma for compressible ($\nu = 0.3$) and incompressible materials ($\nu \approx 0.5$). The red line represents the 900°C isotherm. **b)** Density variations with respect the reference state at 7.1 Ma.

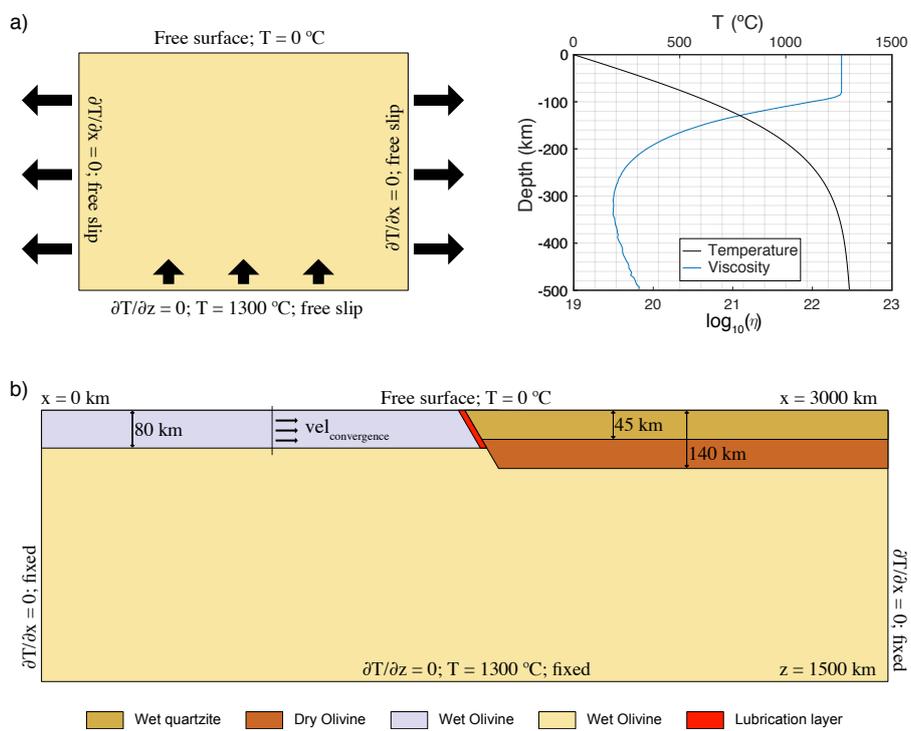


Figure .13: a) Model set-up, boundary conditions and vertical temperature and viscosity profiles of Test A. b) Model set-up and boundary conditions of Test B.

4 | Effects of dilatant pressure-dependent plasticity in geodynamic models

Albert de Montserrat and Jason P. Morgan. Effects of dilatant pressure-dependent plasticity in geodynamic models. Planned for submission to *Tectonophysics*.

Authors contribution

AdM designed the research in discussion with JM. AdM derived and implemented the new constitutive visco-elasto-plastic equation into the numerical code. AdM first designed the numerical experiments to validate the proposed formulation and its implementation. In discussion with JM, AdM programmed and interpreted the tectonic models to further test the implications of plastic dilation in tectonic processes. AdM wrote the manuscript in collaboration with JM.

Effects of dilatant pressure-dependent plasticity in geodynamic models

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Abstract

Volumetric expansion at microscopic and macroscopic scales upon brittle failure has been long accepted to occur in most of the rocks that comprise the continental crust. At greater depths, pressure-solution and melt-wall rock reactions could lead to net mass removal along an active shear zone. However, the mathematical description of models that attempt to mimic the development of faults as observed in geomaterials is often simplified by omitting dilation effects. With this model simplification, different numerical studies have successfully replicated shear bands that are oriented within the range of stable angles provided by analytical solutions; at the same time, they have failed to provide a unified answer, as the orientation of the fault has been proven to be highly sensitive to the employed numerical technique and spatial resolution. We show that, with dilation, an associated flow law combined with a Drucker-Prager failure criterion efficiently generates orientation angles tightly bound to analytical solutions without any compromise in computational cost. In addition, we modify the resulting set of equations to accommodate the possibility of compaction within shear zones as could occur by geological mechanisms such as pressure solution or wall rock reactions during fluid migration along a shear zone. Should chemically linked volume reduction occur along a fault zone, this would influence lithosphere deformation by leading to more rapid localisation of lithosphere scale shear zones.

Keywords:

1 **1. Introduction**

2 The Earth exhibits strong bimodal mechanical behaviour. While the defor-
3 mation and convection of the deep and hot mantle is governed by power-law
4 ductile flow (e.g. Karato and Wu, 1993), tectonic processes such as subduction
5 zones, continental rifting or mountain building, involve the deformation of cold
6 and shallow lithospheric rocks and are characterised by a combination of elastic
7 deformation and brittle failure. The development of constitutive models that
8 capture ductile-brittle deformation and yield results compatible with geologi-
9 cal and laboratory observations has been a key technical hurdle to overcome
10 for better model-based understanding of geological processes at both micro and
11 macro scales.

12 A wide spectrum of brittle failure models is commonly employed in engineer-
13 ing problems in order to simulate the non-recoverable deformation of metals,
14 rocks, concrete, soils and other granular materials. In spite of the discontinu-
15 ous nature of fractures and shear zones, continuum approaches such as the Fi-
16 nite Element Method (FEM) or Finite Difference Method (FDM) are the most
17 widespread numerical tools used in geodynamics to tackle the brittle failure of
18 geomaterials during tectonic processes. The deformation of any given material
19 is described by its so-called constitutive law, which defines the response of the
20 material to external forces. Due to its relative simplicity, an isotropic visco-
21 elasto-plastic constitutive law based on a Maxwell model has been established
22 as a powerful instrument to study coupled ductile-brittle deformation at geo-
23 logical time and length scales (e.g. Braun et al., 2008; Buck, 1991; Buitter et al.,
24 2006; Choi et al., 2013; Fullsack, 1995; Gerya and Yuen, 2007; Huismans and
25 Beaumont, 2007; Kaus, 2010; Lemiale et al., 2008; Moresi et al., 2007; Popov
26 and Sobolev, 2008; Brune et al., 2014; Ros et al., 2017). Another reason to
27 favour this model is that it requires of only two well-constrained physical pa-
28 rameters, namely the material's cohesion and friction angle; while there is a
29 lack of constraints on the mechanical parameters, at geological scales, that de-

30 fine other constitutive models (for example, the fracture energy required damage
31 models to define the amount of inelastic work a material can withstand before
32 it fractures).

33 Plastic deformation is controlled by a scalar function, the *yield surface*
34 or *yield criterion*, that limits the amount of stored stress in any given mate-
35 rial. Although there have been a large number of yield-surfaces proposed for
36 different materials, the yield stress in geodynamics is typically defined by a
37 pressure-sensitive Mohr-Coulomb (Coulomb, 1773) or Drucker-Prager (Drucker
38 and Prager, 1952) yield surface. While there is a noteworthy amount of re-
39 search focused on the numerical implementation of visco-elasto-plastic models
40 that mimic the observed angles and length-scales of shear bands and fractures
41 (e.g. Moresi et al., 2007; Lemiale et al., 2008; Buiter et al., 2006; Buiter, 2012;
42 Kaus, 2010; Popov and Sobolev, 2008; Spiegelman et al., 2016), current results
43 exhibit wide variability. For example, Moresi et al. (2007) observed that shear
44 bands were initiating at 45° using the Drucker-Prager model, independent of
45 friction angle. It was later pointed out that the numerical resolution has to be
46 sufficiently fine (Lemiale et al., 2008) and that mechanical heterogeneities have
47 to be sufficiently well-resolved (Kaus, 2010) for analytical Coulomb angles to
48 be recovered. In addition, the community benchmark presented by Buiter et al.
49 (2006) illustrates that, even though the results obtained from different codes are
50 comparable, there is a non-unique solution for visco-plastic flow, that strongly
51 depends on the choice of numerical technique, which in turns hinders the re-
52 producibility of these numerical results. Furthermore, even the convergence of
53 visco-plastic models containing a dynamic pressure-sensitive yield criterion is
54 under debate (Spiegelman et al., 2016).

55 It is known that plastic yielding in granular materials and rocks is (with
56 some exceptions such as serpentinite (Escartin et al., 1997)) accompanied by a
57 volumetric increase referred to as *dilatancy* (Brace and Byerlee, 1966; Vermeer
58 and De Borst, 1984; Scholz, 2002). This mechanism was neglected in the studies
59 previously mentioned, and has received little attention to date. Interestingly,
60 Moresi et al. (2007) attributes the invariability of the orientation of the shear

61 bands to the suppression of volumetric changes, and Choi and Petersen (2015)
62 conclude that an associated Mohr-Coulomb models yields shear band orienta-
63 tions tightly bounded to the predicted Coulomb angles. However, a thorough
64 analysis of the potential role of plastic volume changes in tectonic processes is
65 still lacking.

66 The aim of this paper is to investigate the effects of plastic dilation in strain
67 localisation and shear band formation and discuss the implications of these ef-
68 fects for macro-deformation at the scale of faults and lithospheric processes.
69 We start by describing a visco-elasto-plastic constitutive law for geodynamic
70 problems that includes a non-elastic volume change via a so-called associated
71 Drucker-Prager flow law. We then demonstrate the viability of this formula-
72 tion with shear-induced dilatancy to reproduce shear band orientations in good
73 agreement with the analytical predictions from bifurcation analyses (Vermeer
74 and De Borst, 1984; Rudnicki and Olsson, 1998), in model conditions that do
75 not require a spatial resolution as fine as that required in previous numerical
76 studies.

77 We further extend our analysis to a rift-like scenario to illustrate poten-
78 tial feedbacks between plastic dilation and tectonic processes (or even fault-
79 slip linked "contraction"). Our results suggest that dilatant and non-dilatant
80 plasticity lead to rifted conjugate margins with a relatively comparable final
81 geometry and faulting history, but the volume increase in associated plastic-
82 ity results in the strengthening of the lithosphere and delays crustal break-up
83 and mantle exhumation. In contrast, if shear zone motions are accompanied by
84 fault-normal contraction or dissolution, the lithosphere experiences enhanced
85 fault localisation and more rapid break-up. In all cases, the thermal evolution
86 is also altered.

87 **2. Plasticity model**

88 Materials undergo non-recoverable plastic deformation if the stress at any
89 material point is such that the yield stress is exceeded. The deviatoric plastic

90 strain rate is then defined as:

$$\varepsilon^{plastic} = \lambda \frac{\partial \mathbb{G}}{\partial \tau} \quad (1)$$

91 where $\lambda \geq 0$ is a plastic multiplier, \mathbb{G} is the plastic potential, τ is the deviatoric
 92 stress tensor and $\varepsilon^{plastic}$ is the plastic component of the deviatoric strain rate
 93 tensor:

$$\dot{\varepsilon} = \frac{1}{2}(\nabla \mathbf{u} + (\nabla \mathbf{u})^T) - \frac{1}{3}(\nabla \cdot \mathbf{u}) \quad (2)$$

94 The elastic domain is defined by the yield surface $\mathbb{F}(\sigma, \xi, h)$, which limits the
 95 maximum stress possible and is a scalar function of the Cauchy stress tensor
 96 σ , a set of material parameters ξ , and the hardening/softening parameter h .
 97 The choice of h is not trivial and will be discussed later on. In domains where
 98 deformation is pure elastic $\mathbb{F} < 0$, whereas $\mathbb{F} = 0$ at yield. If the stress field
 99 at any point of the domain is such that $\mathbb{F} > 0$, the stress needs to be corrected
 100 and brought back to the yield surface ($\mathbb{F} = 0$). In this study we employ a two-
 101 surface yield surface, combining the von Mises (Mises, 1913) and Drucker-Prager
 102 (Drucker and Prager, 1952) criteria:

$$\mathbb{F} = \begin{cases} \alpha(\sigma) - p(\sigma) \sin(\phi(h)) - c \cos(\phi(h)) & \text{if } c \leq \tau_{II} \quad (\text{Drucker-Prager}) \\ \alpha(\sigma) - H & \text{if } c > \tau_{II} \quad (\text{von Mises}) \end{cases} \quad (3)$$

103 where ϕ is the friction angle, c is the cohesion, the pressure is defined as
 104 $p = -tr(\sigma)\mathbf{I}/3$, τ is the deviatoric stress tensor and the subscript II denotes
 105 the square root of the second invariant (i.e. $\mathbf{C}_{II} = \sqrt{(1/2)\mathbf{C} : \mathbf{C}}$, where \mathbf{C} is any
 106 given tensor). The Drucker-Prager criterion is a pressure dependant yield sur-
 107 face and is a smooth (corner free) approximation of the Mohr-Coulomb surface.
 108 On the other hand, the von Mises law is pressure-insensitive and states that
 109 plastic flow begins whenever $\alpha(\sigma)$ reaches a critical value H . In the principal
 110 stress space, Drucker-Prager has a conic shape, whereas von Mises is described
 111 by a cylinder (Fig.1). The parameter $\alpha(\sigma)$ is commonly taken to be the square
 112 root of the second invariant of the deviatoric stress tensor. We also take $H = c$,

113 thus assuming that the yield stress at low stress levels is ultimately defined by
 114 the cohesion of the material.

115 2.1. Flow law

116 Plasticity models can be classified in two types according to the choice of the
 117 plastic potential. If the yield surface is taken to be the plastic potential ($\mathbb{F} = \mathbb{G}$),
 118 this is commonly referred as *associative* plasticity and the material follows the
 119 so-called normality rule: the plastic strain increment vector is normal and moves
 120 outwards from the failure surface. If the plastic potential and yield surface are
 121 different ($\mathbb{F} \neq \mathbb{G}$), this is known as *non-associative* plasticity and the normality
 122 rule is not obeyed.

123 In the geodynamic community, plasticity is often assumed to be non-associative
 124 and the Prandtl-Reus flow rule is the preferred tool to solve plastic deformation
 125 for Maxwell visco-elasto-plastic bodies. An extensive bibliography describing
 126 this formulation is available and the reader is referred to, for example, Moresi
 127 et al. (2003); Gerya and Yuen (2007); Popov and Sobolev (2008); Kaus (2010);
 128 Buiter (2012) for further details. On the contrary, in this paper we focus on the
 129 use of an associated Drucker-Prager flow law model to solve plastic deformation
 130 for problems of mantle-lithosphere deformation.

131 Taking $\mathbb{F} = \mathbb{G}$, the derivative of the plastic potential with respect to the
 132 stress field yields:

$$\frac{\partial \mathbb{G}}{\partial \tau} = \begin{cases} \frac{\tau}{2\tau_{II}} + \frac{1}{3} \sin \phi \mathbf{I} & \text{if } c \leq \tau_{II} \\ \frac{\tau}{2\tau_{II}} & \text{if } c > \tau_{II} \end{cases} \quad (4)$$

133 where \mathbf{I} is the identity matrix. Because \mathbb{F} is defined to be a continuous piecewise
 134 function, \mathbb{G} is continuous piecewise derivable and we do not have to deal with
 135 the derivative at the apex of the Drucker-Prager yield surface. If one wishes
 136 to use a model with only the Drucker-Prager yield surface for the whole stress
 137 domain, the derivative at the apex must be done carefully. Due to the pressure
 138 dependence of the Drucker-Prager yield surface, the plastic strain rate is no
 139 longer purely deviatoric and the stress derivative of the Drucker-Prager yield

140 surface results in both volumetric and deviatoric components:

$$\dot{\varepsilon}^{plastic} = \begin{cases} \lambda(\Upsilon^v + \Upsilon^d) & \text{if } c \leq \tau_{II} \\ \lambda\Upsilon^d & \text{if } c > \tau_{II} \end{cases} \quad (5)$$

141 where $\Upsilon^v = (1/3) \sin \psi \mathbf{I}$ and $\Upsilon^d = \tau/2\tau_{II}$. If $c > \tau_{II}$, the corresponding
 142 yield surface is von Mises, which results in zero volumetric plastic strain. The
 143 volumetric plastic deformation predicted by associated flow rule is often found
 144 to be excessively high and the friction angle in the volumetric plastic strain rate
 145 component is commonly replaced by the so-called dilatancy angle $\psi \leq \phi$. It is
 146 obvious to see that if $\psi = 0^\circ$ non-associated flow rule is recovered and $\Upsilon^v = 0$
 147 for $c \leq \tau_{II}$.

148 [Figure 1 about here.]

149 **3. Rheological model for a visco-elasto-plastic body with an associ-** 150 **ated Drucker-Prager flow rule**

151 Adopting a Maxwell model to describe the deformation of a visco-elasto-
 152 plastic body, the total deviatoric strain rate is given by the summation of its
 153 elastic and inelastic components:

$$\dot{\varepsilon} = \dot{\varepsilon}^{viscous} + \dot{\varepsilon}^{elastic} + \dot{\varepsilon}^{plastic} \quad (6)$$

154 which yields the following the visco-elasto-plastic constitutive equation:

$$\dot{\varepsilon} = \frac{\tau}{2\eta} + \frac{1}{2G} \frac{D\tau}{Dt} + \lambda \frac{\partial \mathbb{G}}{\partial \sigma} \quad (7)$$

155 where η is the viscosity and G is the shear modulus. The time derivative of the
 156 deviatoric stress tensor is computed by approximating the Jaumann derivative
 157 in an implicit manner (e.g. Kaus, 2010):

$$\frac{D\tau}{Dt} \approx \frac{\tau - \tau^o}{\Delta t} - \omega^o \tau^o + \tau^o \omega^o \quad (8)$$

158 where the super-script o refers to the previous time step, Δt is the time step,
 159 and $\omega = 1/2(\nabla \mathbf{u} - (\nabla \mathbf{u})^T)$ is the skew symmetric part of the velocity gradient

160 tensor, commonly known as spin tensor, that rotates the stored stress tensor.
 161 After substitution of eq. 4 into 7, the general visco-elasto-plastic constitutive
 162 equation reads:

$$\dot{\varepsilon} = \frac{\tau}{2\eta} + \frac{1}{2G} \frac{D\tau}{Dt} + \lambda \left(\frac{\tau}{2\tau_{II}} + \frac{1}{3} \sin \psi \mathbf{I} \right) \quad (9)$$

163 To calculate the unknown plastic multiplier, we rearrange the previous equation
 164 using eq. 8, consider that at yield $\tau = \tau_y$, and take the second invariants of $\dot{\varepsilon}$
 165 and $\hat{\tau}$:

$$\lambda = \frac{2\dot{\varepsilon}_{II} + \frac{1}{G\Delta t} \hat{\tau}_{II} - \tau_y \left(\frac{1}{\eta} + \frac{1}{G\Delta t} \right)}{1 + \frac{2}{3} \sin \psi \mathbf{I}} \quad (10)$$

166 where $\hat{\tau} = \tau^o + (\omega^o \tau^o - \tau^o \omega^o) \Delta t$. Now that λ is known, we can rearrange eq. 9
 167 to obtain the following stress-strain relationship:

$$\tau = \eta_{vpl} \left(2\dot{\varepsilon} + \theta \hat{\tau} - \frac{2}{3} \lambda \sin \psi \mathbf{I} \right) \quad (11)$$

168 where $\theta = (G\Delta t)^{-1}$. The effective visco-elasto-plastic viscosity can be computed
 169 directly from eq. 11 assuming yielding conditions:

$$\eta_{vpl} = \frac{\tau_y}{2\dot{\varepsilon}_{II} + \theta \hat{\tau}_{II} - \frac{2}{3} \lambda \sin \psi \mathbf{I}} \quad (12)$$

170 Knowing that $\psi = 0^\circ$ for $c > \tau_{II}$, we can build a piecewise effective viscosity
 171 that covers the whole stress domain:

$$\eta_{vpl} = \begin{cases} \tau_y / (2\dot{\varepsilon}_{II} + \theta \hat{\tau}_{II} + \frac{2}{3} \lambda \sin \psi \mathbf{I}) & \text{if } c \leq \tau_{II} \\ \tau_y / (2\dot{\varepsilon}_{II} + \theta \hat{\tau}_{II}) & \text{if } c > \tau_{II} \end{cases} \quad (13)$$

172 where the effective viscosity for $c > \tau_{II}$ is equivalent to the expression obtained
 173 by employing the Prandtl-Reus flow rule. Finally, we can build a general consti-
 174 tutive relationship for visco-elasto-plastic materials as:

$$\tau = 2\eta^* \dot{\varepsilon} + \theta \eta^* \hat{\tau} - \langle 1 \rangle \frac{2}{3} \lambda \sin \psi \mathbf{I} \quad (14)$$

175 where $\eta^* = \eta_{vpl}$ and $\langle 1 \rangle = 1$ for $\tau > \tau_y$, and $\eta^* = \eta_{eff}$ and $\langle 1 \rangle = 0$ for $\tau \leq \tau_y$.
 176 The effective viscosit is defined as (e.g. (Kaus, 2010)):

$$\eta_{eff} = \frac{1}{\frac{1}{\eta} + \frac{1}{G\Delta t}} \quad (15)$$

177 *3.1. Strain softening*

178 To mimic the post-peak stress drop often seen in granular materials and
 179 rocks, strain softening, as a function of the softening parameter h , is applied to
 180 some of the mechanical properties that define the strength of the material. As
 181 commonly used in geodynamics (e.g. Huismans and Beaumont, 2002; Buitert,
 182 2012; Choi and Petersen, 2015), we take the accumulated plastic strain as the
 183 softening parameter, and the initial values of ϕ and ψ are linearly reduced with
 184 with increasing accumulated plastic strain. The latter is defined as:

$$\mathbf{E}^{pl} = \int \sqrt{\frac{1}{2} \dot{\boldsymbol{\varepsilon}}^{plastic} : \dot{\boldsymbol{\varepsilon}}^{plastic}} dt \quad (16)$$

185 where the colon represents a dyadic contraction of the plastic strain-rate tensor.
 186 The friction angle is chosen so that $\phi(\mathbf{E}^{pl} = 0) = \phi_0$ and $\phi(\mathbf{E}^{pl} \geq \mathbf{E}_{max}^{pl}) = \phi_\infty$.
 187 Strain softening is also applied to the dilatancy angle following Choi and Pe-
 188 tersen (2015), who inferred that ϕ and ψ should have initially their maximum
 189 *undamaged* value so that a shear band can form at a Coulomb angle. An-
 190 other reasons in favour of reducing the dilatancy angle are that, with increas-
 191 ing slip: 1) faults do not have an everlasting expansive behaviour (Detour-
 192 nay, 1986) and will evolve from associated plastic state towards non-associative
 193 plastic deformation; and 2) mass (i.e. fluids) within the dilatant shear band
 194 can diffuse into the new pore spaces and thereby decrease the ambient pore
 195 pressure Rudnicki (1988). These effects are coarsely simulated by linearly re-
 196 ducing the dilatancy angle with increasing accumulated plastic strain so that
 197 $\psi(\mathbf{E}^{pl} \geq \mathbf{E}_{max}^{pl}) = \psi_\infty = 0^\circ$, and a non-associated state is recovered. There
 198 is a general lack of studies to constrain the values of \mathbf{E}_{max}^{pl} and ϕ_∞ ; however,
 199 it is common to adopt values within the range of $0.25 \leq \mathbf{E}_{max}^{pl} \leq 1.25$ and
 200 $2 \leq \phi_\infty \leq 15$. In the models further presented in this paper, we adopt $\mathbf{E}_{max}^{pl} = 1$
 201 and $\phi_\infty = 15$. Alternatively, one could use the accumulated plastic work as to
 202 define the softening curve (de Souza Neto et al., 2011). However, we do not
 203 investigate this option here.

204 **4. Numerical code**

205 [Figure 2 about here.]

206 We use the LaCoDe code (de Montserrat et al., 2018), based on a Lagrangian
 207 formulation of the Finite Element Method, to solve the coupled equations of
 208 conservation of momentum, conservation of mass and conservation of energy:

$$\nabla \sigma = \rho \mathbf{g} \quad (17)$$

$$\rho \frac{D\rho}{Dt} + \nabla \cdot \mathbf{u} = 0 \quad (18)$$

$$\rho C_p \frac{DT}{Dt} = k \nabla^2 T + H_Q + H_{sh} \quad (19)$$

209 where ρ is the density, \mathbf{g} is the gravitational acceleration, \mathbf{u} is the velocity field,
 210 C_p is specific heat, T is temperature, κ is thermal conductivity, H_Q is radioactive
 211 heating and H_{sh} is shear heating. Density corrections are calculated using a
 212 temperature and pressure dependent equation of state:

$$\rho(T, p) = \rho_o \left(1 - \alpha (T - T_{ref}) + \frac{p}{K} \right) \quad (20)$$

213 where ρ_o is the reference density, α is the thermal expansivity, T_{ref} is the
 214 reference temperature and K is the bulk modulus. We use a non-Newtonian
 215 rheology to describe diffusion and dislocation creep, where the viscosity is given
 216 by a strain rate- and temperature- dependent power law (Poirier, 1985; Karato
 217 et al., 2001):

$$\eta = \frac{1}{2} (A)^{-\frac{1}{n}} (\dot{\epsilon}_{II})^{\frac{1}{n}-1} \exp \left(\frac{E_a + pV_a}{nRT} \right) \quad (21)$$

218 where A is a pre-exponential parameter, n is a power-law exponent, E_a is ac-
 219 tivation energy, V_a is activation volume and R is the universal gas constant.
 220 The second invariant of the deviatoric strain rate in eq. 21 corresponds to ei-
 221 ther the diffusion or dislocation creep deviatoric strain rate tensor. A resultant
 222 composite viscosity is obtained as:

$$\eta = \frac{1}{\frac{1}{\eta_{dif}} + \frac{1}{\eta_{dis}}} \quad (22)$$

223 where deformation is dominated by the mechanism with the smallest activation
 224 stress.

225 Eq. 19 describes the conservation of mass for a compressible material where
 226 volumetric deformation is linked to density changes. When at yield, eq. 19 is
 227 modified to include the volumetric deformation from the dilatant plasticity:

$$\frac{D\rho}{Dt} + \rho \nabla \cdot \mathbf{u} = \rho \lambda \Upsilon^v \quad (23)$$

228 For a (nearly) incompressible material, the material time derivative of the den-
 229 sity in eq. 23 vanishes, and the extended Boussinesq approximation will be
 230 recovered:

$$\nabla \cdot \mathbf{u} = \lambda \Upsilon^v \quad (24)$$

231 Using a mixed formulation to solve the Stokes equations, the Cauchy stress is
 232 split into its deviatoric and volumetric components:

$$\boldsymbol{\sigma} = \boldsymbol{\tau} + p \quad (25)$$

233 and using eqs. 11 and 25, the equation of conservation of momentum yields:

$$\nabla (2\eta^* \dot{\boldsymbol{\varepsilon}}) + \nabla \mathbf{p} = \rho \mathbf{g} - \nabla (\theta \eta^* \hat{\tau}_{II}) + \nabla \left(\langle 1 \rangle \eta^* \frac{2}{3} \lambda \sin \phi \mathbf{I} \right) \quad (26)$$

234 Eqs. 23 and 26 are solved combining Powell-Hestenes iterations with the penalty
 235 method (e.g. Dabrowski et al., 2008). Rheological (non-Newtonian viscous creep
 236 and plasticity) and density (compressible continuity equation) non-linearities are
 237 treated with Picard iterations.

238 5. Shear band initiation

239 [Figure 3 about here.]

240 The implementation of the associated and non-associated flow rules is tested
 241 with a numerical experiment modelling shear band initiation. Similar tests
 242 can be found in, for example, Popov and Sobolev (2008); Kaus (2010); Choi
 243 and Petersen (2015); Spiegelman et al. (2016). The domain of this numerical

244 experiment is a 20 by 10 km rectangular visco-elasto-plastic box with a density
 245 of 2700 kg/m^3 that simulates the uppermost part of the crust. The mechanical
 246 parameters are $G = 30 \text{ GPa}$, $c = 30 \text{ MPa}$, $\nu \approx 0.5$, and a uniform viscosity
 247 $\eta = 10^{23} \text{ Pa}\cdot\text{s}$. An extension rate of 1 cm/year is prescribed at the right hand-
 248 side edge of the domain, and free-slip boundary conditions are prescribed at
 249 the bottom and left boundaries. The surface of the model is traction-free (i.e.
 250 a free surface). A mechanical heterogeneity of $1 \text{ by } 1 \text{ km}$ is introduced at the
 251 bottom-centre of the model, so that shear bands will nucleate around it. The
 252 friction angle of this weak region is set to be $\phi = 0^\circ$, while $\phi_0 = 30^\circ$ elsewhere.
 253 We perform a suite of tests with the arbitrary values of the dilatancy angle
 254 $\psi_0 = \phi_0$, $\psi_0 = \phi_0/2$ and $\psi_0 = 0^\circ$, with strain softening applied to the friction
 255 ($\phi_\infty = 15^\circ$) and dilatancy angle ($\psi_\infty = 0^\circ$). The models are run only until the
 256 mesh is so distorted that a remeshing algorithm would be necessary.

257 To avoid any geometrical bias derived from a preferred orientation of the
 258 mesh, the domain is discretised into triangular elements that are regularly and
 259 symmetrically distributed along the domain. This mesh is constructed by 1)
 260 dividing the domain in a regular grid formed by rectangles, and 2) splitting
 261 the rectangles into four triangles with a common vertex at the centre of the
 262 rectangle. The effect of spatial resolution on associated plasticity is assed by
 263 running the model with different numerica resolutions: fine ($4 \cdot 10^4$ triangular
 264 elements), intermediate ($2.25 \cdot 10^4$ triangular elements) and coarse (10^4 triangu-
 265 lar elements). The aspect ratio (height/base) of the triangular elements is 0.25
 266 for all meshes considered.

267 *5.1. Results*

268 [Figure 4 about here.]

269 From bifurcation analysis using a Mohr Coulomb failure criterion, Vermeer and
 270 De Borst (1984) deduced three mechanically stable angles for shear bands: 1)
 271 $\theta = 45^\circ + \frac{\psi}{2}$, Roscoe (1970); 2) $\theta = 45^\circ + \frac{\phi}{2}$, Coulomb (1773) ; and 3) $\theta =$
 272 $45^\circ + \frac{\phi+\psi}{4}$, Arthur et al. (1977). Again from a bifurcation analysis, but this time

273 employing a Drucker-Prager criterion, Rudnicki and Olsson (1998) obtained the
274 alternative expression for a stable shear band orientation:

$$\theta = 45^\circ + \frac{1}{2} \arcsin \alpha \quad (27)$$

275 where $\alpha = (\phi + \psi)/2$ in the incompressible case under extension and pure shear
276 boundary conditions.

277 All the models are able to consistently produce a single pair of well-converged
278 conjugate shear bands that initiate at the mechanical heterogeneity (Fig .3).
279 Depending on the spatial resolution, shear bands initiate at approximately 57.5–
280 60.5° for associated ($\psi = \phi$) models (Table 5.1). On the other hand, non-
281 associated ($0 < \psi < \phi$) models yields shear bands with orientations around
282 55.5 – 57.5°, whereas non-dilatant models yield shear bands at 51.5 – 53.5°.
283 In spite of some minor angle deviations for $\psi \neq \phi$, these values show that
284 the plastic formulation presented here using a associated Drucker-Prager flow
285 rule yields shear bands in close agreement with the orientations predicted by
286 Rudnicki (1988), rather than the Coulomb, Arthur or Roscoe angles (Fig .4a).

287 The build up of volumetric plastic strain within the shear bands translates
288 into a relative structural hardening with increasing dilatancy angle. This can be
289 inferred from the stress-strain curves (Fig .4b, where the vertical axis represents
290 the integrated stress along the right edge of the model): while for $\psi_0 = 0^\circ$ the
291 model shows an immediate post-peak strain-softening behaviour with a sharp
292 drop of stress, dilatant models are able to sustain near-peak stress conditions
293 shortly after yielding, and their stress drop is more sustained and progressive.
294 Strain softening is not observed in the case of $\psi_0 = 0^\circ$ because the models
295 are run only until the quality of the mesh becomes too poor, therefore, in this
296 particular case, strain softening has not kicked in enough to become noticeable
297 in the stress-displacement curve.

298 The plastic strain within the shear bands increases with the dilatancy angle
299 as a result of a non-zero volumetric plastic strain rate. This enhances strain
300 localisation within the shear bands, resulting in higher strain rate in the yielded
301 material. One would expect that higher plastic strain rates will be reflected

Table 1: Shear band orientation as a function of the dilatancy angle ψ for different spatial resolutions.

| ψ ($^{\circ}$) | Shear band orientation ($^{\circ}$) | | | Mohr-Coulomb ($^{\circ}$) | | | Drucker-Prager ($^{\circ}$) |
|-----------------------|---------------------------------------|------------|------------|-----------------------------|--------|--------|-------------------------------|
| | 10000 els. | 22500 els. | 40000 els. | Coulomb | Arthur | Roscoe | Rudnicki |
| 30 | 55 | 59.5 | 60.5 | 60 | 60 | 60 | 60.8 |
| 15 | 54 | 57.5 | 57.5 | 60 | 56.25 | 52.5 | 56.6 |
| 0 | 51.5 | 53 | 53.5 | 60 | 52.5 | 45 | 52.6 |

Table 2: Vertical fault displacement and width of the graben. Values are normalised with respect the case $\psi = 0^{\circ}$.

| ψ ($^{\circ}$) | Normalised vertical fault displacement | | |
|-----------------------|--|------------|------------|
| | 10000 els. | 22500 els. | 40000 els. |
| 30 | 0.87 | 0.89 | 0.91 |
| 15 | 1.02 | 1.01 | 1.01 |
| 0 | 1 | 1 | 1 |

302 on the topographic profile as larger vertical fault displacements. However, for
 303 associated models ($\phi = \psi$) the build up of volumetric strain within the shear
 304 bands introduces a vertical motion in the hanging wall block that limits vertical
 305 fault displacement (Fig.4c; Table 5.1).

306 Another effect of dilatant plasticity is that the additional inelastic strain
 307 rate results in higher inelastic work dissipation, which translates to higher tem-
 308 peratures due to additional shear heating. This could have consequences in, for
 309 example, numerical models of rifted margins, as it might further enhance strain
 310 localisation when using a temperature-dependent viscosity.

311 6. Case study: rifting of continental crust

312 In the previous section we examined the effects of dilatant plasticity on shear
 313 band initiation in small scale models. However, it is not obvious what *a priori*
 314 effects this behaviour will have on large scale geodynamic problems, or whether
 315 it even has any noticeable effects at all. We therefore further explore the effect
 316 of dilatant plasticity on strain localisation, fault geometry, thermal structure,

317 and structural evolution of an evolving rift.

318 *6.1. Model set-up*

319 The domain consists of a 500 by 500 km box (Fig.5) divided into distinct
320 rheological layers: wet quartzitic upper crust (UC), dry olivine upper mantle
321 (UM) and wet olivine lower mantle (LM). We further assess the impact of dila-
322 tion on lower crusts (LC) with different strengths: i) a weak end member (wet
323 quartzite), and ii) a strong end member (mafic granulite). The corresponding
324 rheological parameters are shown in Table 6.1. Pure shear far-field boundary
325 conditions are prescribed at the boundaries of the model (i.e. half and full exten-
326 sion rate at the lateral and bottom boundaries of the domain, respectively), and
327 a tangential free slip condition is employed at the lateral and bottom boundaries;
328 the top boundary of the domain is treated as a free surface (Andrés-Martínez
329 et al., 2015). Temperature is kept fixed at 0°C at the surface and at 1200° C
330 below 120 km depth. In order to localise deformation at the centre of the model
331 at the onset of extension and to avoid artefacts arising from boundary effects,
332 we introduce a thermal Gaussian-shape perturbation in the lower crust at 35
333 km depth. The half extension rate prescribed at the edges of the domain is 32.5
334 mm/yr.

335 We note that the set-up of the models is representative of typical numerical
336 experiments aiming at modelling different rifting scenarios; however, the aim of
337 these models is to assess solely the effects of associative plasticity on lithospheric
338 stretching. Therefore, we will not explore the detailed evolution of rifting in
339 these models.

340 [Figure 5 about here.]

341 *6.2. Strong lower crust: fault geometry and model evolution*

342 Models with a mafic granulite lower crust result in a pair of asymmetric con-
343 jugate rifted margins (Fig.6). After an initial phase of widely distributed fault-
344 ing strain localises in a pair of main conjugate normal faults that cut throughout

345 the whole crust, coupling crust and upper mantle, and deformation is accom-
346 modated by pure shear. As the crust stretches, strain further localises in the
347 E-dipping conjugate normal fault. The W-dipping normal fault becomes pro-
348 gressively inactive and the deformation mode switches to simple shear. In this
349 stage, many secondary shear zones develop in the rifting region, along with a
350 high strain listric fault that root into the main one. Further extension results in
351 the rotation of the main normal fault and progressive crustal thinning of one of
352 the margins, whereas the crust of the other conjugate margin is sharply thinned.
353 Due to the rapid extension rate, the upwelling of mantle material is quite fast
354 and crustal break-up and mantle exhumation are reached at about 2.5-2.7 Ma.

355 The effect of dilatant plasticity on the fault distribution is most noticeable in
356 the 2 and 2.2 Ma snapshots (Fig.6a). Localisation of plastic strain is enhanced
357 and dilatancy angles of $\psi > 0^\circ$ yield a higher number of well-defined shear
358 zones at crustal depths. Structural hardening is produced by the blocks of
359 unyielded material being *locked* by the expansion of the shear zones, becoming
360 more noticeable after the rifting develops (2.0 Ma and onwards). This effect is
361 ultimately responsible for a delay in the evolution of the rifting. Consequently,
362 mantle exhumation occurs earlier (at c. 2.5 Ma) for $\psi = 0^\circ$, whereas mantle
363 is not exhumed until c. 2.6 and 2.7 Ma after onset of extension for $\psi = 15^\circ$
364 and $\psi = 30^\circ$, respectively. Plastic strain is less effectively localised in the non-
365 dilatant model which results in smoother relief (Fig.7a,b,c). In contrast, dilatant

Table 3: Rheological parameters. The upper and lower crust are weak wet quartzite (Gleason and Tullis, 1995), the upper and lower mantle are dry olivine and wet olivine (Hirth and Kohlstedt, 2003), respectively.

| Parameter | Description (Units) | Wet olivine | Dry olivine | Mafic granulite | Wet quartzite |
|----------------------|--|------------------|------------------|------------------|------------------|
| ψ_0/ψ_∞ | $^\circ$ | 30/0 | 30/0 | 30/0 | 30/0 |
| ϕ_0/ϕ_∞ | $^\circ$ | 30/15 | 30/15 | 30/15 | 30/15 |
| c | Cohesion (Mpa) | 20 | 20 | 20 | 20 |
| A | Pre-exponential factor ($\text{Pa}^{-n_s s^{-1}}$) | $10^{-15.56}$ | $10^{-15.56}$ | $10^{-21.05}$ | 10^{-28} |
| E | Activacion energy (KJmol^{-3}) | $480 \cdot 10^3$ | $530 \cdot 10^3$ | $445 \cdot 10^3$ | $223 \cdot 10^3$ |
| V_o | Activation volume($\text{m}^3 \text{mol}^{-3}$) | 10^{-4} | 10^{-6} | 0 | 0 |
| n_{dis} | Power-law exponent (dislocation creep) | 3.5 | 3.5 | 4.2 | 4 |
| n_{dif} | Power-law exponent (diffusion creep) | 1 | 1 | 0 | 0 |

366 models present a rougher surface due to more effective faulting at shallow depths;
367 however, the amplitude of the relief is slightly smaller as a result of the structural
368 hardening.

369 [Figure 6 about here.]

370 [Figure 7 about here.]

371 *6.3. Weak lower crust: fault geometry and model evolution*

372 [Figure 8 about here.]

373 As is well known from previous numerical studies of rifted margins (e.g. Huis-
374 mans and Beaumont, 2011, 2014; Ros et al., 2017; Tetreault and Buitert, 2017),
375 a weak lower crust results in different margin evolution with respect to models
376 with a strong end member. In early stages, strain localises into a set of widely
377 distributed shear zones. After c. 1.5 Ma, strain localises into two high strain
378 rate regions located at the flanks of the weak seed which results in the develop-
379 ment of two basins. These high strain regions are characterised by several sets
380 of conjugate normal faults. Due to its weak mechanical behaviour, ductile flow
381 predominates in the lower crust and focussed shear zones do not continue into
382 the mantle lithosphere. Hence the conjugate normal faults die out in the upper
383 section of the lower crust and the crust and mantle lithosphere remain mechan-
384 ically decoupled within the rifting region. Further stretching is accompanied by
385 ductile flow of the lower crust that leads to a thickening of the upper crust and
386 thinning of the lower crust at the centre of the model, with two domes of lower
387 crustal material forming at ~ 45 km away from the centre.

388 Extension is accompanied by ascent of lower mantle material and break-up
389 of the mantle lithosphere occurs prior to crustal break-up at c. 1.5 Ma. As
390 stretching continues, the deeper mantle asthenosphere flows laterally towards
391 regions where the crust is weaker (i.e. where a considerable amount of upper
392 crust has been removed and replaced by weak and hot lower crust), eventually
393 leading to crustal break-up and mantle exhumation, and resulting in two highly
394 asymmetric rifted margins.

395 Wet quartzite is mechanically weaker than mafic granulite, hence it more
396 readily deforms by ductile flow than by brittle failure. Therefore, the effects
397 of associative plasticity are less evident here than for the strong end member
398 case in terms of rift evolution. Nonetheless, strain localises better in associated
399 models when plastic deformation is efficient enough (Fig.8, 2.2 Ma). The over-
400 all ductile behaviour of these models results in smooth topographies with two
401 major basins and very little difference between dilatant and non-dilatant models
402 (Fig.7d,e,f).

403 *6.4. Thermal evolution*

404 Plastic dilation may not only affect the deformation history of conjugate
405 rifted margins, but has consequences for the thermal evolution due to two main
406 mechanisms: 1) dilatant hardening leads to a locking of non-yielded material in
407 between shear zones; as a consequence, the locking strengthens the crust and
408 slows down the upwelling of mantle material, thereby delaying mantle exhumation
409 and crustal break-up with respect to non-dilatant models; and, 2) dilatant
410 plastic strain leads to a higher inelastic strain rate and therefore higher heating
411 due to inelastic work. The first mechanism promotes the cooling of the model
412 as vertical motion is reduced with increasing dilation. The second mechanism
413 enhances thermal localisation, as the heating from inelastic work would increase
414 with increasing dilatancy angle. The resulting thermal evolution is determined
415 by the competition between both mechanisms.

416 For a strong lower crust, plastic deformation is very effective and shear
417 heating becomes the dominant mechanism for intermediate degrees of dilation
418 ($\psi = 15^\circ$), resulting in a higher geotherm at lithospheric depths in comparison
419 with a non-associated model. However, hardening cooling dominates for a highly
420 dilatant ($\psi = 30^\circ$) crust, thus leading to cooler thermal structures (Fig.9a).
421 For a weak lower crust, the plastic contribution to shear heating is lower than
422 viscous dissipation and volumetric locking becomes a more efficient as a cooling
423 mechanism with increasing dilatancy angle (Fig.9c). Nonetheless, independent
424 of the strength of the lower crust, the surface heat flow is significantly reduced as

425 a consequence of increasing dilation within the faults at shallow crustal depths
426 (Fig.9b,d) as the result of cooling derived from the structural hardening being
427 more effective than shear heating.

428 [Figure 9 about here.]

429 6.5. Depth dependant dilatancy angle

430 As argued by Detournay (1986), dilation is not only a function of plastic
431 strain, but also of confining stress. The effect of confining stress was further
432 noted by Medhurst (1997), from triaxial tests on multi-scale coal samples, and
433 Ribacchi (2000) from standard triaxial tests on limestone samples with variable
434 fractures. We investigate this argument by introducing a depth-dependent di-
435 latancy angle. As in Alejano and Alonso (2005), at null confining pressure we
436 consider that the dilation and friction angle are the same, i.e. $\psi_0(z = 0) = \phi_0$.
437 The dilatancy angle is then linearly reduced as a function of depth until a cer-
438 tain depth limit, i.e. $\psi_0(z = z_{limit}) = 0^\circ$. We perform a sensitivity test of the
439 depth-dependency of the dilatancy angle in the model with a strong lower crust
440 taking values of $z_{limit} = 60, 40$ and 20 km (Fig.10).

441 A similar initial phase of distributed faulting at the initial stages of deforma-
442 tion (≤ 1.7 Ma) is observed in all the models with different depths of dilatation
443 efficiency. For $z_{lim} \leq 40$ km, strain localises in two main conjugate shear zones
444 that cut through the whole crust, coupling its deformation with the mantle
445 lithosphere (Fig.10a, 2.0 Ma). With further extension, the W dipping shear
446 zone becomes inactive and strain is localised in the remaining main shear zone.
447 The latter rotates to lower dipping angles and listric faults that root on this
448 main detachment begin to develop (Fig.10a, 2.2 Ma), thereby forming a single
449 basin (Fig.11b,c). At this stage, deformation shifts from pure shear to simple
450 shear. At c. 2.7 the lithosphere breaks-up and mantle is exhumed, resulting in
451 two asymmetric conjugate margins (Fig.10b).

452 In contrast, models with a depth limit deeper than the base of the crust for
453 dilatant plasticity show a different history of brittle deformation after the initial

454 phase of distributed faulting (≤ 1.7 Ma). At c. 2 Ma, strain has localised in two
455 shear zones that start at the centre of the surface with opposite orientations.
456 These faults also root into the base of the lower crust, coupling deformation
457 between the crust and mantle lithosphere. This different set of main shear
458 zones results in two mini basins with an uplifted crustal block in between them
459 (Fig.10a, 2.2 Ma; Fig.11a). Due to the stronger hardening effects, lithospheric
460 break-up and mantle exhumation is delayed until c. 3 Ma. However, this model
461 also results in a pair of asymmetric conjugate margins (Fig.10b).

462 The depth at which dilation ceases is also reflected in the thermal signature
463 in the models (Fig.10c): shallower depth limits for dilation yield higher peaks
464 of surface heat flow, whereas for deeper depth limits the peak of surface heat
465 flow at the same time is approximately three times smaller.

466 [Figure 10 about here.]

467 [Figure 11 about here.]

468 **7. Compaction of shear bands**

469 Opposite to the porosity increase due to the brittle-faulting-linked dilatancy,
470 other mechanisms lead to mass loss or compaction are known to occur within
471 shear zones. Some of these processes include pressure solution (e.g. Renard et al.,
472 2000; Bos et al., 2000; Gratier et al., 2011) and reactive fluid or melt migration
473 along faults and fractures. While the physical and mathematical description of
474 these mechanisms is not represented by the equations developed for an associ-
475 ated Drucker-Prager flow rule, we can simulate the effect of compaction within
476 shear bands by changing the sign of the right-hand-side of eq. (23). We must
477 emphasize this is a more a "thought experiment" than a quantitative simulation
478 of the effects of fault-slip-linked mass removal processes.

479 In this case, the compaction of yielded material leads to a structural softening
480 of the crust, enhancing strain localisation and leading to focussed shear bands
481 with higher strain rates (Fig. .12). The faulting history of the crust follows the

482 same pattern as described in Sections 6.2 and 6.3. However, the enhancement
483 of strain localisation during the early stages of extension (0.9-1.4 Ma) in the
484 model with a weak lower crust results in two narrower basins at the flanks of
485 an non-deformed crustal block (Fig. .12b).

486 As described in the earlier sections, an associated (dilatant) plastic model
487 causes a structural hardening that slows down the deformation history of the
488 lithosphere. In contrast, compaction within shear zones will leads to structural
489 softening that speeds up the evolution of deformation. This results in break-up
490 of the lithosphere occurring around 30% (c. 1.75 Ma) and 42% (c. 2.15 Ma)
491 earlier for the strong and weak lower crust cases, respectively, with respect to
492 previous models with $\psi = 0^\circ$.

493 [Figure 12 about here.]

494 8. Discussion

495 The admissible shear band orientation lies within the range of Roscoe-
496 Arthur-Coulomb angles, and orientations outside this range are not expected.
497 Using a Drucker-Prager yield surface, Lemiale et al. (2008) inferred that high
498 spatial resolutions are required in order to obtain shear bands at Coulomb an-
499 gles. This was later affirmed by Kaus (2010), who attested that the mechani-
500 cal heterogeneity has to be well resolved with at least 5-10 elements. In con-
501 trast, Choi and Petersen (2015) obtained shear bands spread around Arthur
502 and Roscoe angles for non-associated $\phi_0 \neq \psi_0$ flow rules using a Mohr-Coulomb
503 failure criterion. Our analysis based on the formulation presented in this work
504 using a Drucker-Prager criterion is in agreement with the results from Choi and
505 Petersen (2015) for $\phi_0 = \psi_0$, with shear band orientations that yield Coulomb
506 angles; however, our results show shear bands tightly bounded to the analytical
507 solution derived by Rudnicki and Olsson (1998). Coulomb angles also form with
508 $\psi = 15^\circ$ and the orientations are scattered around the Coulomb-Arthur-Roscoe
509 spectrum only in the non-dilatant case.

510 The stress-displacement curves obtained from the shear band initiation il-
511 lustrate the structural hardening inherent from associated models and the post-
512 peak softening, as described by Vermeer and De Borst (1984) and also in agree-
513 ment with hardening models addressing, for example, earthquake instabilities
514 (Rudnicki (1988) and references therein).

515 As discussed in Spiegelman et al. (2016), visco-plastic Stokes problems may
516 become ill-posed for pressure-dependant rheologies (e.g. Drucker-Prager) and
517 Picard iterations might stall at large residuals. In spite of the deep implications
518 of this issue for geodynamical models of lithospheric deformation, this diffi-
519 culty did not arise during this work. However, following the conclusions drawn
520 by Spiegelman et al. (2016), we must emphasize the importance of cautious-
521 ness when interpreting results that may not reflect fully converged numerical
522 experiments. To enhance the convergence rate, the extension rate applied to
523 the model is typically linearly increased during the first ten time-steps until it
524 reaches 1 cm/yr. Strong convergence may not be fully reached in the initial
525 time steps; however, convergence improves significantly after a few time steps
526 and consistent patterns of shear bands develop.

527 The solution of the associative Drucker-Prager flow rule still exhibits a con-
528 siderable mesh-dependent effects where shear band orientations in models with
529 a fine and coarse mesh can have up to 5° of difference. One proposed remedy
530 for this algorithmic issue consists adjusting the softening modulus as a func-
531 tion of element size. This was first proposed Pietruszczak and Mroz (1981) for
532 shear softening in plasticity and by Bažant and Oh (1983) for tensile softening
533 caused by smeared cracking. These techniques lead to improved element-size
534 insensitivity; however, the solution can still depend on the assumed element
535 shape and orientation. This computational issue has been overcome by apply-
536 ing so-called non-local continuum plasticity models (Bažant et al., 1988). In
537 these models, stress is no longer determined uniquely by the strain history and
538 temperature at that given point alone but also depends on the strain history of
539 surrounding material points, with interactions exceeding a certain length-scale
540 being neglected. Although effective computationally, this approach still involves

541 ad-hoc, non-physical assumptions. Mesh-dependent effects on strain localisation
542 problems remain unresolved and are clearly a key phenomenon that still needs
543 to be properly addressed in geodynamic models using (visco-)plastic rheologies.

544 Numerical examples of rifting of continental crust suggest a non-trivial ef-
545 fect of plastic dilation on large scale geodynamic problems. Even though the
546 geometry of numerical fault evolution during continental rifting does not seem
547 to be greatly altered by associated plasticity, the resulting hardening creates a
548 strengthening of the crust that leads to a delay in the evolution of the margins,
549 consequently delaying crustal break-up and mantle exhumation. The surface
550 heat flow and thermal structure of the models is also altered as a result of the
551 competition between shear heating and cooling resulting from the structural
552 hardening.

553 As pointed out from experimental observations, the dilation of rocks de-
554 pends on the confining pressure (e.g. Medhurst, 1997; Ribacchi, 2000), or in
555 other words, depth. Our experiments with linear depth-dependent dilatancy
556 angle suggest that there is not much difference between models where dilation
557 becomes negligible at upper ($z_{lim} = 20$ km) and lower crustal depths ($z_{lim} = 40$
558 km); the only apparent difference between these models is a small delay in the
559 evolution of deformation, and a slightly lower surface heat flow peak. However,
560 a different faulting history is observed if brittle fault dilation can remain effec-
561 tive at lithospheric depths ($z_{lim} = 60$ km); for these the heat signature at the
562 surface is considerably reduced.

563 In spite of the mathematical and physical incompleteness of the formulation
564 presented in this paper to properly describe compaction effects in shear bands
565 due to processes such as pressure solution or reactive fluid migration through
566 fractures, a small modification of the continuity equation allows us to explore
567 the geodynamic consequences of relative mass loss within a shear zone at a
568 lithospheric scale. Our results show that this effect leads to a strong structural
569 softening of the lithosphere that speeds up the deformation of the crust and
570 asthenospheric upwelling. From this, we suggest that volume losses to reactive
571 fluid migration along shear zones could play a significant role driving the defor-

572 mation of the lithosphere during tectonic events, and that this process deserves
573 a more complete thermo-mechanical treatment and exploration.

574 Depth-dependent associated plasticity may also be important in other tec-
575 tonic settings such as in subduction zones, as it could enhance plate-bending
576 and increase the porosity of fracture zones, thus facilitating fluid circulation. It
577 could also influence in the seismic cycle, as failure/earthquake occurrence can
578 be delayed by the resultant structural hardening.

579 **9. Summary**

- 580 • Plastic models with a Drucker-Prager yield function and an associative
581 flow rule yield shear bands that dip at the angles predicted by Rudnicki
582 and Olsson (1998), whereas a non-associative deviatoric Prandtl-Reus flow
583 rule yields shear band orientations scattered within the wider range of
584 Coulomb-Arthur-Roscoe angles.
- 585 • Volumetric expansion within the shear bands results in a structural hard-
586 ening of the faulting domain.
- 587 • The evolution and final geometry of rifted margin experiment does not
588 show significant structural differences between dilatant and non-dilatant
589 models.
- 590 • The slower vertical motion of mantle upwelling induced by the struc-
591 tural hardening of the lithosphere leads to relative cooling in the models,
592 whereas volumetric plastic strain introduces an extra local source of heat
593 derived from inelastic work. The thermal structure and surface heat flow
594 in dilatant models is therefore altered with respect to non-dilatant models.
595 For a strong lower crust, brittle deformation is highly effective and dilatant
596 models yield similar temperature fields for the range of dilatancy angles
597 considered in this paper. However, plasticity is less effective in models with
598 a weak lower crust and shear heating dominates for $\psi_0 = 15^\circ$, whereas the

599 model with $\psi_0 = 30^\circ$ is cooler, due to the slower upwelling rates of mantle
600 material.

- 601 • The evolution of the faults and shear zones of a model with a depth-
602 dependent dilatancy angle with $\psi_0 = \phi_0$ at null confining pressure repre-
603 sents an intermediate state between using constant $\psi_0 = 30^\circ$ and constant
604 $\psi_0 = 15^\circ$. During the early stages of extension, this model is very similar
605 to a model with constant $\psi_0 = 30^\circ$, while it evolves towards a similar
606 strain rate state as the model with $\psi_0 = 15^\circ$ as strain softening kicks in.

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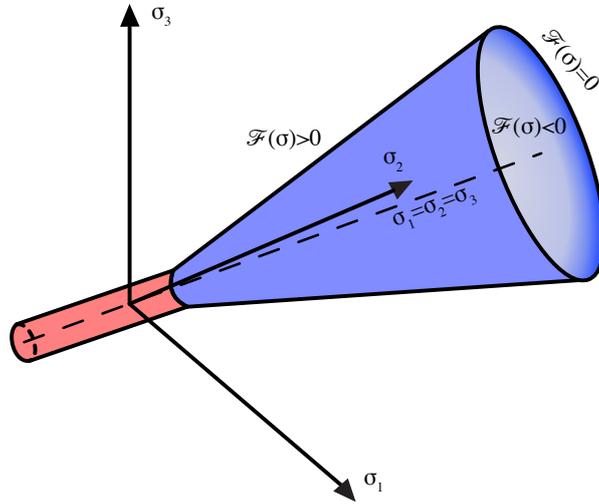


Figure .1: Sketch of the combined Drucker-Prager (in blue) and von Mises (in red) yield surface.

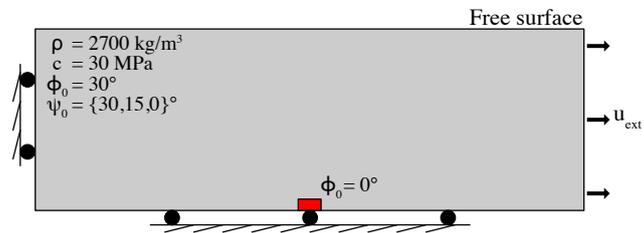


Figure .2: Set-up of the shear band initiation model. The region in red represents a mechanical heterogeneity with $\phi = 0^\circ$.

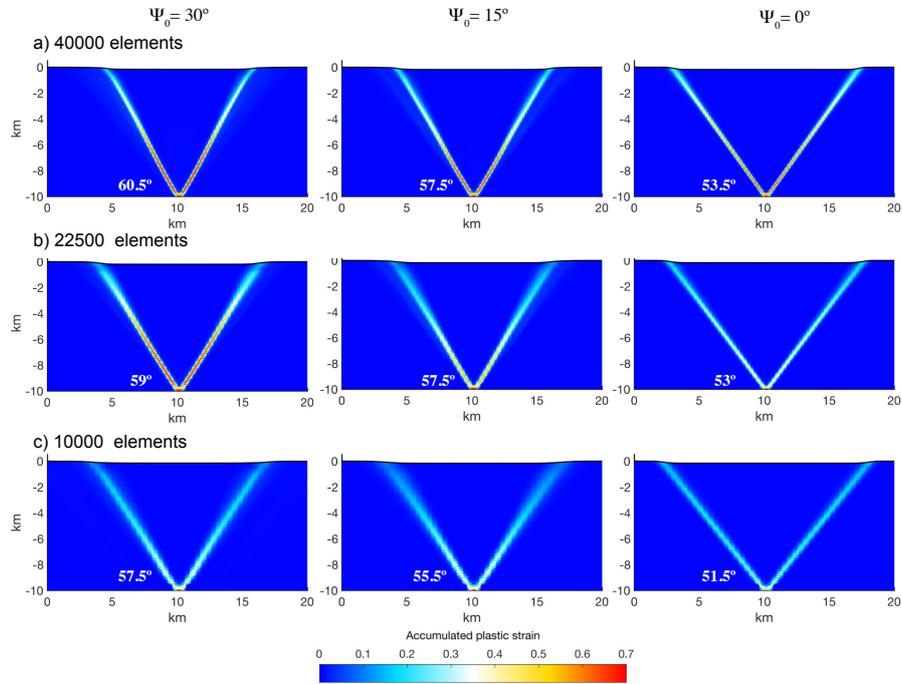


Figure .3: Shear band results for associative and non-associated flow rules with $\psi = 30^\circ$, $\psi = 15^\circ$ and 0° . Models have a resolution of **a)** 40000, **b)** 22500 and **c)** 10000 triangular elements with an aspect ratio (height/base) of 0.25. Strain softening is applied to ϕ and ψ in all models. The colour maps represent the rate square root of the second invariant of the accumulated plastic strain rate \mathbf{E}^{pl} .

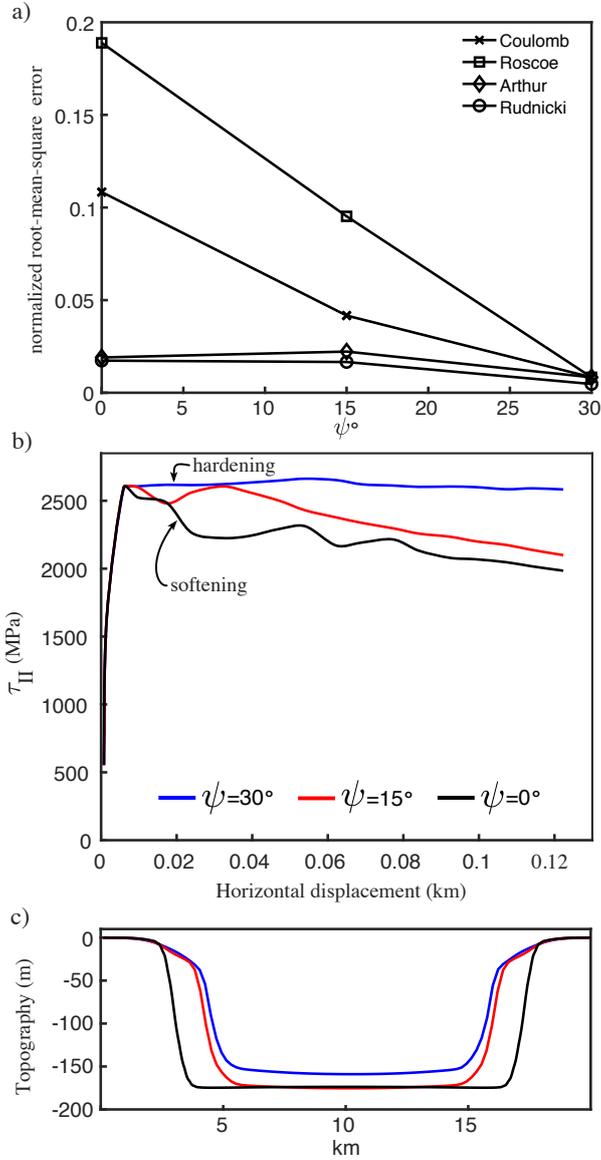


Figure 4: **a)** Normalised root-mean-square (rms) error between the numerical orientations of the shear bands and the analytical Coulomb, Roscoe, Arthur and Rudnicki angles. **b)** Stress-horizontal displacement curve. The vertical axis corresponds to the stress integrated along a vertical profile at the right boundary of the domain. **c)** Topographic expression for all models; the vertical axis of the plot is vertically exaggerated. The results in a), b) and c) correspond to the high-resolution models (40000 elements).

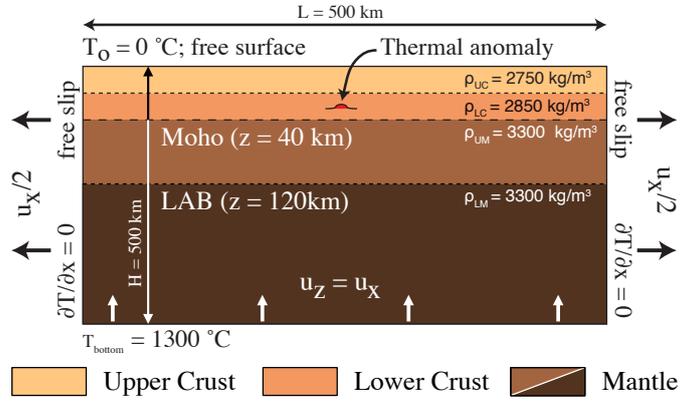


Figure .5: Model set-up. The model is stretched using pure shear boundary conditions. To localise the deformation in the centre of the model, a thermal anomaly is introduced in the middle of the lower crust at $x = 0$ km. The rheological parameters are given in Table 6.1.

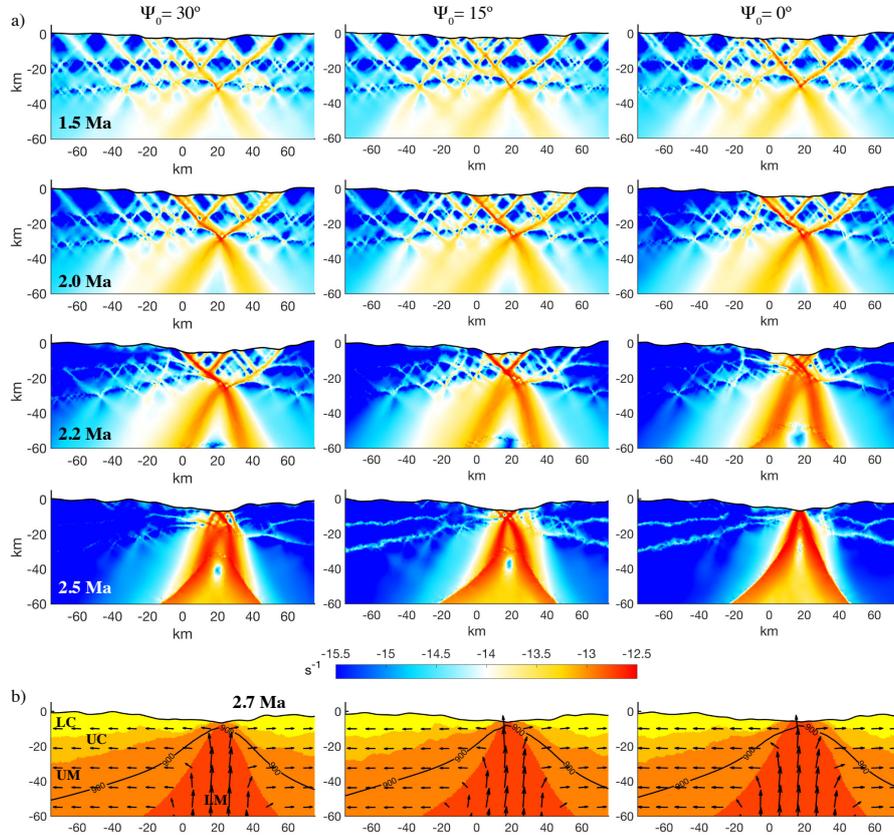


Figure 6: Models with a mafic granulite lower crust (strong end member). **a)** Evolution of the square root of the second invariant of strain rate ($\dot{\epsilon}_{II}$) for $\psi_0 = 30^\circ$, 15° (associative plasticity) and 0° (non-associative plasticity). **b)** Comparison of the geometry of the model at 2.5 Ma. Colour maps indicate the different rheological layers and arrows indicate velocity.

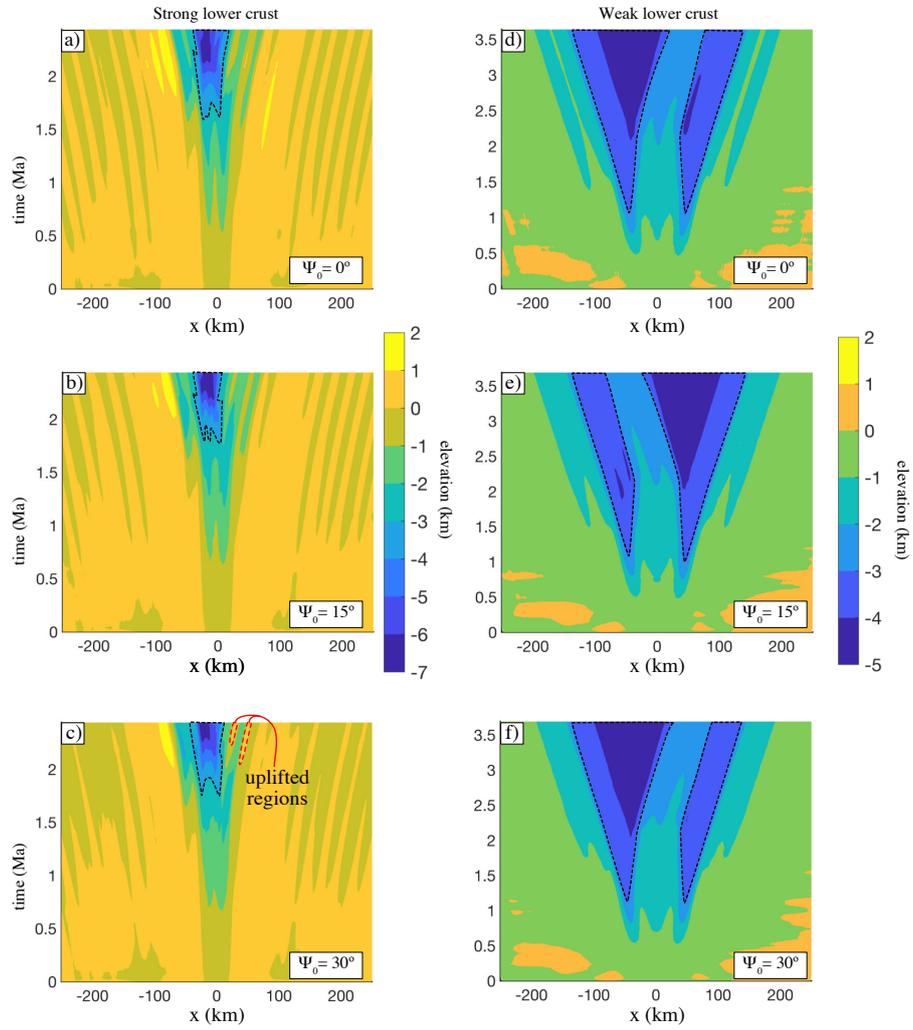


Figure .7: Surface evolution with time for a strong **a)** lower crust and **b)** weak lower crust for different values of dilatancy angle. Areas within black and red dashed polygons indicate deep water basins and uplifted regions, respectively.

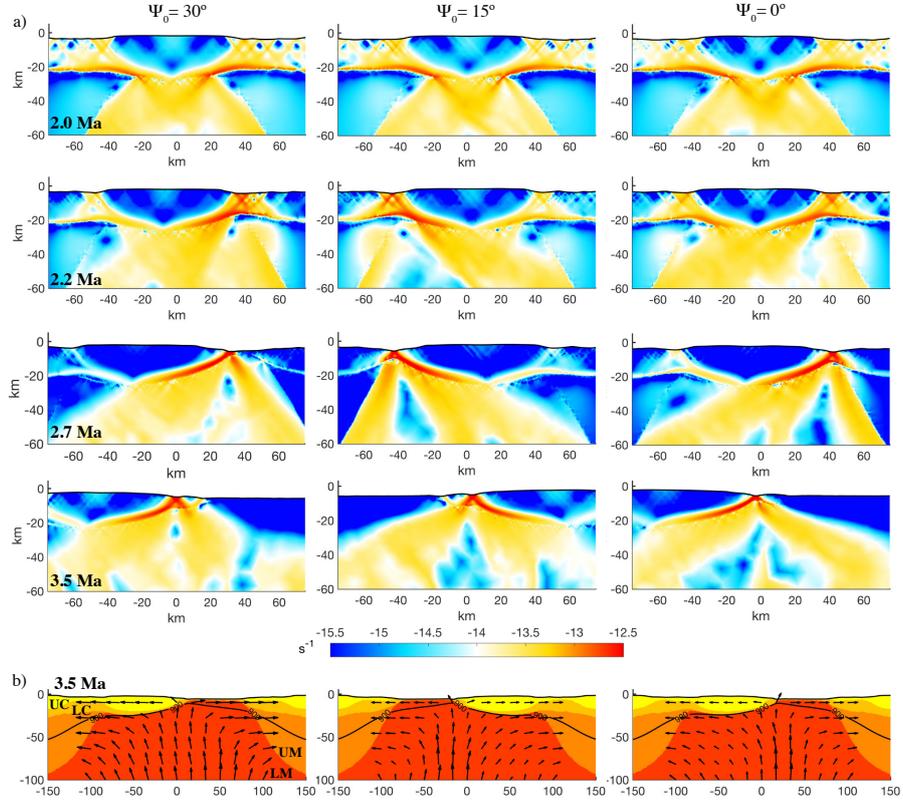


Figure 8: Models with a wet quartzitic lower crust (weak end member). **a)** Evolution of the square root of the second invariant of strain rate ($\dot{\epsilon}_{II}$) for $\psi_0 = 30^\circ$, 15° (associative plasticity) and 0° (non-associative plasticity). **b)** Comparison of the geometry of the model at 2.5 Ma. Colour maps indicate the different rheological layers and arrows indicate velocity. **c)** Values of the dilatancy angle at 2.5 Ma.

5 | Rapid cooling and exhumation of lower crust. Insights from numerical models and application to SE Asia.

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Authors contribution

AdM, JM and RH defined the research. AdM designed the numerical experiments and analysed the results in discussion with JH, JM and RH. AdM wrote the manuscript in collaboration with JH, JM and RH.

Rapid cooling and exhumation of lower crust. Insights from numerical models and application to SE Asia

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Abstract

Recent episodes of extension in SE Asia have been associated with rapid sedimentary basin growth, and phases of crustal melting, uplift and extremely rapid exhumation of young (Early-Late Pliocene) metamorphic complexes. We combine geochronological and geothermobarometric data with two-dimensional numerical models to investigate exhumation of metamorphic core complexes in Sulawesi, Indonesia. The lithospheric thermal conditions and extension rates at which these developed and later exhumed are poorly known. Therefore, we explore a wide range of potential initial conditions with different permutations of extension rate and initial Moho temperature. The numerical models show that high Moho temperatures are key to shaping the architecture of the stretched lithosphere. Hot and weak lower crust fails to transmit stress and brittle deformation to deeper regions, resulting in a strong decoupling between crust and lithospheric mantle. In this case, deformation is dominated by ductile flow, yielding the exhumation of one-to-several partially molten lower crustal bodies. Continental break-up is often inhibited by the ductile behaviour of the lower crust, and is only achieved after considerable cooling of the lithosphere. Further comparison between the observed and synthetic model T-t paths confirms that extremely rapid exhumation of lower crustal bodies should be linked to very fast extension rates (~ 75 mm/yr) and Moho temperatures (>740 °C) higher than those in more commonly studied rift settings (e.g. Atlantic opening,

East African Rift, Australia-Antarctica opening).

Keywords: SE Asia, Sulawesi, Palu Metamorphic Complex, Fast exhumation of lower crust, Slow, rapid and ultra-rapid metamorphic complexes

1. Introduction

SE Asia represents one of the most tectonically active regions of the Earth, as the result of the convergence of the Eurasian, Indo-Australian and Philippine Sea plates. Recent thermobarometric studies have revealed SEA as the host of arguably the youngest continental Metamorphic Core Complexes (MCCs) in Sulawesi (Hennig et al., 2014, 2016, 2017; van Leeuwen et al., 2016), Seram (Pownall, 2015; Pownall et al., 2017) and D'Entrecasteaux Islands (Little et al., 2011; Baldwin et al., 2004) ever reported, as the consequence of a series of extensional episodes during the Neogene (Spakman and Hall, 2010; Hall, 2011, 2012; Pownall et al., 2013; Hennig et al., 2014).

A two-phase thermal history of the lower crustal rocks, consisting of an initial phase of slow cooling, followed by a brief (few million years) period of rapid cooling, at rates of 60-280 °C/Ma, have been reported in extension-driven Cordilleran-type MCCs in the Basin and Range province (e.g. John and Howard, 1995), Cyclades (e.g. Scott et al., 1998) and North China Craton (e.g. Yang et al., 2007), as well as in MCCs in the Canadian Cordillera thought to have been developed by gravitational collapse (e.g. Vanderhaeghe et al., 2003). Although this two-phase thermal history has been attributed to variations in denudation rate, localised thermal perturbations or changes in the geometry of the detachment faults, the nature of the pulses of rapid cooling remain poorly understood.

Recent thermobarometric and geochronological data from rocks of the Palu Metamorphic Complex (PMC) in Sulawesi (Hennig et al., 2017), indicate that cooling rates in MCCs can be higher than those previously reported (≥ 540 °C/Ma). Unfortunately, and despite the increasing number of geological and geophysical studies carried out in the last decades, the initial thermal structure,

27 and extension rates, responsible for the development of the PMC and other
28 MCCs in SEA remains poorly constrained.

29 A numerical approach has become a powerful tool to investigate different
30 aspects that control the formation of the MCCs, such as partial melting and
31 extension rates Rey et al. (2009), inherited crustal layering Huet et al. (2011a)
32 and wedge structures Huet et al. (2011b), origin of the heat source necessary
33 for migmatization Schenker et al. (2012), and role of the density and viscosity
34 of the deep crust Korchinski et al. (2018). The thermobarometric history of
35 MCCs has been also explored by numerical models, and synthetic p-T paths
36 have yielded reasonable comparisons with natural counterparts (Huet et al.,
37 2011a,b; Schenker et al., 2012). However, the cooling history of the MCCs have
38 received little attention from a numerical point of view.

39 In this paper, we use two-dimensional thermo-mechanical numerical models
40 to study the thermal history of rocks comprising MCCs and what extension
41 rates are necessary to reproduce the cooling rates observed in the geological
42 record. In particular, we compare synthetic T-t paths against available ther-
43 mochronological data from the PMC (Hennig et al., 2017). To find the best fit
44 with the natural cooling paths, we run a suite of numerical models with differ-
45 ent permutations of initial conditions, namely extension rate and initial Moho
46 temperature.

47 We further investigate whether the volume and distribution of crustal melt-
48 ing plays a crucial role in the thermal history of continental MCCs. For this
49 reason, we consider two solidi with that lead to different volumes of partially
50 molten crust. An initial volume of partial melting is not prescribed anywhere
51 in the model, thus partial melting is self-consistently produced according to
52 pressure and temperature conditions.

53 The different combinations of Moho temperature and extension rate result
54 in the formation of continental core complexes with distinct thermal histories
55 showing that: i) slow extension rates lead to cooling rates of 60-300 °C/Ma,
56 and ii) high cooling rates as in the PMC (>300-540 °C/Ma) are reproduced in
57 models under rapid and ultra-rapid extension (≥ 35 mm/yr).

58 **2. Extension of Central Sulawesi: rapid exhumation of the Palu Meta-**
59 **morphic Complex**

60 [Figure 1 about here.]

61 The western continental margin of Sundaland in Sulawesi is formed by Aus-
62 tralian crust that rifted from the Australian margin during the Late Jurassic
63 and was accreted to Sundaland in the Late Cretaceous (Smyth et al., 2007; Hall
64 et al., 2009; Hall, 2011). Rifting from Borneo occurred in the Eocene (Hamilton,
65 1979; Weissel, 1980), related to widespread extension in SE Asia which opened
66 the Celebes Sea followed by formation of the North Sulawesi volcanic arc (Hall,
67 2012).

68 At c. 45 Ma, Australia started to move northward. The collision of the
69 northern promontory of the Australian margin, called the Sula Spur (Klompe,
70 1954), with the forearc of North Sulawesi volcanic arc started in the Early
71 Miocene at c. 23 Ma (Spakman and Hall, 2010; Hall, 2011). Northward move-
72 ment of Australia continued after the collision until c. 17 Ma, when eastward
73 propagation of a tear in the slab from the Java trench was accompanied by
74 rapid subsidence and rollback of the subduction zone into the Banda embay-
75 ment (Spakman and Hall, 2010; Hall, 2018). Banda rollback led to widespread
76 extension above the subducting slab, resulting in the formation of oceanic crust
77 and opening of the North Banda Sea between 12.5 and 7 Ma (Hinschberger
78 et al., 2000), and caused extension-related magmatism in West Sulawesi (Polvé
79 et al., 1997; Hennig et al., 2016). Rocks associated with this magmatism are
80 not well known and were described as a high- potassium suite, which includes
81 nepheline gabbros, quartz-syenites, and monzodiorites of presumed Middle to
82 Late Miocene age (Priadi et al., 1994; Polvé et al., 1997; Elburg et al., 2003).
83 Extension led to crustal thinning and weakening in the study area, which de-
84 fines the starting conditions for our numerical model that is based on a hot and
85 relatively thin lithosphere.

86 The PMC is located in the neck of Sulawesi (Fig. D.1a,b). The metamorphic
87 rocks of the PMC were initially interpreted to be Permo-Triassic basement rocks

88 recording a major rift phase on the margin of New Guinea (Sukanto, 1973; van
89 Leeuwen and Muhardjo, 2005). More recent studies from U-Pb zircon dating of
90 schists of the PMC revealed that some of these rocks have Eocene protoliths,
91 and therefore must be younger (van Leeuwen et al., 2016; Hennig et al., 2016).
92 The PMC is typically divided into a metapelite unit in the west and a gneiss
93 unit in the east Sukanto (1973); van Leeuwen and Muhardjo (2005). The meta-
94 morphic grade increases to the east, and the gneiss unit is comprised mainly of
95 high-grade metamorphic rocks: biotite granite-gneisses and biotite-amphibole
96 granite-gneisses, and subordinate pyroxene gneisses, marbles and migmatites.

97 Radiometric dating of magmatic and metamorphic rocks of the PMC showed
98 evidence of contemporaneous Pliocene magmatism, metamorphism and exhuma-
99 tion, suggesting the PMC must have developed in an extensional setting that
100 includes significant stretching of the upper plate and rapid exhumation of deep
101 crust; features that resemble a metamorphic core complex in an extensional
102 setting (van Leeuwen and Muhardjo, 2005; Hennig et al., 2014, 2017). The de-
103 tachment fault has not been observed and its inferred position is in mountains
104 in the neck of Sulawesi where there is dense rainforest vegetation, and accessi-
105 bility is limited. Mylonitic shear zones observed in biotite hornfels and slates in
106 the northern PMC are interpreted as related to subordinate detachment faults
107 (Hennig et al., 2017). The metapelites of the PMC are strongly deformed and
108 the youngest metamorphism has pervasively overprinted older fabrics. Some
109 textural evidence for near-isothermal decompression was observed by staurolite,
110 andalusite, and cordierite or pinitite porphyroblasts which are sometimes man-
111 tled by white mica coronas, indicating disequilibrium to secondary low-pressure
112 conditions.

113 T-t paths of S-type magmatic and metamorphic rocks (Fig. D.1d) from
114 the PMC have been estimated by combining U-Pb zircon rim ages from biotite-
115 amphibole gneiss Hennig et al. (2016) with $^{40}\text{Ar}/^{39}\text{Ar}$ cooling ages of amphibole
116 and biotite from amphibolite and biotite schist (Hennig et al., 2017), yielding
117 remarkably high cooling rates (Fig. D.1d) that suggest an unusually rapid ex-
118 humation during the Late Pliocene to Pleistocene (Hennig et al., 2017). At the

119 same time there was the youngest extension phase affecting northern Sulawesi
120 due to northward subduction rollback of the North Sulawesi trench. Previously,
121 this rollback was explained as a response to clockwise rotation of the North
122 Arm (Hamilton, 1979; Silver et al., 1983; Surmont et al., 1994) and left-lateral
123 movement along the Palu-Koro fault at the western end of the trench (Walpers-
124 dorf et al., 1998; Socquet et al., 2006). However, new studies reveal a more
125 complex history related to development of subduction, and the deepest part of
126 the subducted slab is located in the centre of the subduction zone (Hall, 2018).
127 Our numerical experiments are aimed to model the period of the last 3-5 Ma of
128 exhumation of the PMC, during which the subducted slab was deep and dense
129 enough to drive subduction rollback.

130 **3. Numerical modelling of stretched crust**

131 *3.1. Numerical code*

132 We use a modified version of the viscous flow solver MILAMIN (Dabrowski
133 et al., 2008), based on a Lagrangian formulation of the Finite Element Method,
134 to solve the coupled equations of conservation of momentum, conservation of
135 mass and conservation of energy. Incompressibility is incorporated by using the
136 Boussinesq approximation and the mechanical behaviour of the rocks is modelled
137 as a Maxwell body with a visco-elasto-plastic rheology: viscous deformation
138 is described by a power-law constitutive equation, and plastic deformation is
139 computed employing a pressure dependent Drucker-Prager yield surface. Strain
140 localisation is enhanced by the reduction of the angle of friction as a linear
141 function of the accumulated plastic strain (e.g. de Souza Neto et al., 2011). A
142 brief description of the numerical formulation is provided in Appendix A.

143 *3.2. Model set-up and initial conditions*

144 [Figure 2 about here.]

145 The domain of the models consists in 500 by 400 km rectangular box (Fig.
146 D.2a) that is divided in four laterally homogeneous rheological layers: a 35

147 km thick crust comprised of a wet quartzitic upper crust (UC) and lower crust
148 (LC), a 85 km thick dry olivine lithospheric mantle (LM) and wet olivine as-
149 thenospheric mantle (AM). The rheological and mechanical parameters used to
150 describe the thermo-mechanical behaviour of these layers are shown in Table 3.2.
151 Pure shear far-field boundary conditions (BCs) are prescribed on the boundaries
152 of the model (half extension rate prescribed at the side boundaries, and full ex-
153 tension rate at the bottom), with a shear stress-free BC employed on the lateral
154 and bottom boundaries. The top boundary of the domain is treated as a free
155 surface. We introduce a Gaussian-shape thermal perturbation to promote strain
156 localisation after onset of extension. This thermal weak seed is emplaced at the
157 centre of the model to avoid boundary effects that might corrupt the results.
158 At the onset of the extensional event that led to the formation of the PMC the
159 crust had already undergone a previous stretching event. For this reason, we
160 define a relatively thin crust of 35 km.

161 The thermal gradient in the Banda Sea region is poorly constrained. How-
162 ever, the lithosphere must have been very hot during the extensional event that
163 triggered the development and exhumation of the PMC (Hall, 2018). To provide
164 better constraints on over the initial thermal gradient, we explore a set of differ-
165 ent permutations of initial Moho temperature: 1) cold: 710 °C; 2) intermediate:
166 844 °C ; 3) warm: 911 °C ; and 4) hot: 1040 °C.

167 We also consider a range of different extension rates: 1) slow: 10 mm/yr;
168 2) rapid: 35 mm/yr; and, 3) ultra-rapid: 75 mm/yr. The rapid and ultra-rapid
169 rates are faster than the ones used in previous numerical studies of the formation
170 of MCCs (e.g. Huet et al., 2011a; Rey et al., 2011; Schenker et al., 2012), and
171 are chosen to investigate the plausible range of extension rates responsible for
172 the evolution of the Banda Sea region during the last 5-10 Ma.

173 *3.3. Partial melting*

174 Partial melting is computed using the approach described by Morgan (2001)
175 when pressure and temperature conditions of any parcel of the model exceed the
176 solidus temperature (Fig.D.2b). Partial melting occurring at the LC is calcu-

Table 1: Rheological parameters. The upper and lower crust are weak wet quartzite (Gleason and Tullis, 1995), the upper and lower mantle are dry olivine and wet olivine (Hirth and Kohlstedt, 2003), respectively. Density values within brackets indicate the density of a fully molten rock.

| Parameter | Description | Lower Mantle | Upper Mantle | Lower Crust | Upper Crust |
|---------------|--|-------------------|-------------------|---------------------|---------------------|
| c | Cohesion (Mpa) | 20 | 20 | 20 | 20 |
| ϕ_o | Peak friction angle ($^\circ$) | 30 | 30 | 30 | 30 |
| ϕ_∞ | Minimum friction angle ($^\circ$) | 15 | 15 | 15 | 15 |
| ρ | Density (kgm^{-3}) | 3300 (2900) | 3300 (2900) | 2850 (2400) | 2700 (2400) |
| G | Shear modulus (GPa) | 74 | 74 | 40 | 36 |
| α | Thermal expansivity | $3 \cdot 10^{-5}$ | $3 \cdot 10^{-5}$ | $2.4 \cdot 10^{-5}$ | $2.4 \cdot 10^{-5}$ |
| H_Q | Radioactive heating (Wm^{-3}) | 0 | 0 | $0.2 \cdot 10^6$ | $1.3 \cdot 10^6$ |
| K | Thermal conductivity ($\text{Wm}^{-3}\text{K}^{-3}$) | 3.3 | 3.3 | 2.5 | 2.1 |
| A | Pre-exponential factor ($\text{Pa}^{-n_s s^{-1}}$) | $10^{-15.56}$ | $10^{-15.56}$ | 10^{-28} | 10^{-28} |
| E | Activacion energy (KJmol^{-3}) | 480 | 530 | 223 | 223 |
| V_o | Activation volume($\text{m}^3\text{mol}^{-3}$) | 10^{-4} | 10^{-6} | 0 | 0 |
| n_{dis} | Power-law exponent (dislocation creep) | 3.5 | 3.5 | 4 | 4 |
| n_{dif} | Power-law exponent (diffusion creep) | 1 | 1 | 0 | 0 |

177 lated using the solidus (Solidus A) of a hydrated granite (Boettcher and Wyllie,
178 1968), and decompression melting of the mantle is calculated according to the
179 solidus of a fertile peridotite (Morgan, 2001). Additionally, we investigate the
180 evolution of MCCs in reduced crustal partial melting conditions by considering
181 the solidus (Solidus B) of a fluid-absent MORB-derived amphibolite (López and
182 Castro, 2001). We note that this solidus does not correspond to the LC rheol-
183 ogy used in the models and has been chosen merely as an end member of low
184 partial melting productivity. Segregation of the melt from its source is not con-
185 sidered and we assume that it moves *en masse* (Teyssier and Whitney, 2002).
186 Rosenberg and Handy (2005) pointed out that the viscosity of crustal rocks is
187 significantly reduced for $F > 7\%$, independent of the melting or crystallizing
188 state of the rock. Following this conclusion, the viscosity of crustal rocks in the
189 models is linearly reduced as a function of melt fraction. Density is also linearly
190 reduced with increasing melt fraction.

191 **4. Dynamics of lower crustal exhumation**

192 A total of 24 model calculations have been performed for all the possible
193 combination of initial Moho temperature, extension rate and LC solidus stated
194 in Section 3.2 . All these models develop a LC dome that is exhumed along a
195 detachment fault. The dynamics of the formation of these domes share some
196 common features that are observed in all the models:

197 [Figure 3 about here.]

- 198 1. An initial stage (total extension: $\leq 4\%$) dominated by either pure shear
199 (under slow extension; Fig. D.3.1a) or simple shear (under rapid and
200 ultra-rapid extension; Fig. D.3.1b), resulting in two conjugate normal
201 faults that root in the LC, producing a single graben.
- 202 2. Between 4-6 % of total extension (Fig. D.3.2), the LC dome starts to
203 develop, and the conjugate normal faults are abandoned. During this
204 stage, ductile flow of basal crustal material feeds the dome, resulting in
205 a gradual thinning of the LC at the flanks of the dome. Strain localises
206 in a detachment fault that starts at the centre of the graben and roots at
207 the top of the crustal dome. At this stage, some secondary shear bands
208 may develop, but crustal deformation remains mainly accommodated by
209 the detachment fault.
- 210 3. As extension continues (total extension: 6-8%), the detachment fault ro-
211 tates, reaching low angles (15-20°) close to the surface, whereas the dip
212 remains within 50-60° at mid-crustal depths. Ductile flow of the LC keeps
213 filling the gap left by the stretched UC and the dome is progressively
214 exhumed along the detachment (Fig. D.3.3).
- 215 4. Exhumation of the LC dome occurs typically after 8-10% of extension and
216 it is followed by a period of lateral spreading (Fig. D.3.4).

217 After 10-15% of extension, the evolution of the models differs depending
218 on the initial conditions. Moreover, the topography of the Moho, formation
219 of secondary domes, partial melting production, and thermal history also vary

220 according to the initial conditions. The results are further summarised in the
221 following sections.

222 *4.1. Cold models: $T_{Moho} = 710\text{ }^{\circ}\text{C}$*

223 [Figure 4 about here.]

224 The dynamics of extension and crustal dome formation during the early stages
225 (extension $< 10\%$) are as described above. Synchronous with the growth of
226 the dome, the upwelling of mantle material leads to a positive topography of
227 the Moho underneath the crustal dome, and decompression melting takes place
228 in the ascended mantle at about 50 km depth. In the late stages of exten-
229 sion ($> 10\%$), shear zones penetrate into the Moho, coupling the deformation
230 between lithosphere and asthenosphere and leading to exhumation of mantle
231 material and, consequently, mantle exhumation and formation of oceanic crust
232 (Fig. D.4a). Melting of the LC is not predicted under these thermal conditions
233 for models employing the Solidus B and the final state of the model corresponds
234 to two conjugate margins with a gradually thinned crust.

235 *4.2. Intermediate models : $T_{Moho} = 844\text{ }^{\circ}\text{C}$*

236 Under this initial thermal structure (Fig. D.4b), the detachment remains
237 active throughout the period of domal growth and it is only abandoned after
238 the upper crust is broken apart and lateral spreading of the dome begins. In
239 contrast to models with a cold initial geotherm the elevated temperature of
240 the LC inhibits the transmission of stress to the Moho. As a consequence,
241 and even though the Moho bends upwards due to upwelling of asthenospheric
242 material, the crust never ruptures. The base of LC remains under partial melting
243 conditions, and molten material is advected towards the surface as the dome
244 grows, resulting in a mushroom-shaped region of partially molten crust.

245 During the first stages of the development of the LC dome in rapid and ultra-
246 rapid models ($\geq 35\text{ mm/yr}$), the asthenosphere flows upwards and localises right
247 underneath the dome. With further extension, the rise of asthenospheric ma-
248 terial concentrates below one of the flanks of the dome, and as a consequence,

249 the LC experiences conductive heating at the contact with the lower mantle,
250 leading to additional partial melting production and Ultra High Temperature
251 (UHT) conditions. These heated rocks form a secondary dome that is pro-
252 gressively exhumed along a new detachment fault that roots in the LC. These
253 secondary domes are not predicted by slow extension rates. Further extension
254 of the model produces extreme thinning of the LC and decompressional melting
255 of the mantle, located underneath the dome or the (if there is any) secondary
256 dome. Mantle exhumation is only reached at very late stages of deformation
257 (extension $> 45\%$), after significant crustal thinning and cooling.

258 In slow and intermediate temperature models, ductile flow of the deep crust
259 inhibits crustal thinning and the Moho remains flat during the growth and
260 exhumation of the LC dome. Under these conditions, doming of the Moho and
261 decompression melting of the mantle are only observed after c. 18 Ma, triggered
262 by the ascent of asthenospheric material.

263 4.3. Warm and hot models : $T_{Moho} \geq 911 \text{ }^\circ\text{C}$

264 Strain localises in the LC as a detachment fault, leading to the formation
265 of a LC dome. After the dome has been exhumed to the surface along the de-
266 tachment and the latter becomes inactive, strain localises in new detachments
267 along the model. These detachment faults are associated with the development
268 of LC domes. As a consequence of the high temperatures, stress is not trans-
269 mitted through the LC, the crust remains uncoupled from the mantle, and the
270 topography of the Moho remains almost flat (Fig. D.5a,b). Exceptionally, the
271 topography of the Moho of warm models under fast extension ($\geq 35 \text{ mm/yr}$) is
272 slightly bent upwards.

273 During the growth and exhumation of the LC, there is no melt productivity
274 in the mantle, and UHT conditions at the base of the crust are reached only
275 for $u_{ext} \geq 35 \text{ mm/yr}$. In this set of experiments, decompression melting of the
276 asthenosphere and mantle exhumation is observed only at c. 15 Ma and c. 8
277 Ma for rapid and ultra-rapid extension, respectively, after hyperextension of the
278 crust and a considerable cooling.

279 *4.4. p-T paths of lower crustal rocks*

280 Synthetic p-T paths (supplementary Figs. 1 and 2) are obtained by tracking
281 the evolution through time of pressure and temperature of a set of passive mark-
282 ers located within the LC. We refer to the markers initially located at 19, 27 and
283 34 km depth as shallow, mid and deep markers respectively (see Fig.D.8a,b for
284 their location). The shape of the p-T paths shows a high sensitivity to extension
285 rate as fast extension rates promote near-isothermal decompression. In contrast,
286 the initial Moho temperature does not have a significant influence in the shape
287 of the p-T paths and only increases (or reduces) the average temperature of the
288 curves.

289 Similar p-T paths are observed in rapid and ultra-rapid models with cold-
290 to-warm ($710 \leq T_{Moho} \leq 911$ °C) initial Moho temperatures, where a bimodal
291 thermobarometric history of the LC rocks record is recorded. While deep and
292 mid markers record an initial phase of near-isothermal decompression (down
293 to 0.2 GPa), shallow markers experience an initial phase of decompressional
294 heating as the result of heat advection, reaching peak conditions of 550-600 °C
295 and c. 0.3 GPa. The markers cross from the kyanite to the sillimanite to the
296 andalusite stability field during isothermal decompression, with the exception of
297 shallow rock domes that cross from kyanite directly into the andalusite stability
298 field. In contrast, slow models show significant differences in the p-T records
299 between cold ($T_{Moho} = 710$ °C) and intermediate-to-warm ($844 \leq T_{Moho} \leq 911$
300 °C) initial Moho temperatures. In the cold case, near-isothermal decompression
301 of deep markers occurs from about 0.85 GPa to 0.45 GPa, and mid and shallow
302 markers record cooling following a path subparallel to the kyanite-andalusite
303 transition. In warmer models ($T_{Moho} \geq 844$ °C), deep markers record decom-
304 pressing heating with peak conditions of 0.6-0.75 GPa and 750-800 °C, followed
305 by cooling along a geothermal gradient of c. 10 °C/km that switches to c. 35
306 °C/km at c. 0.3 GPa, until final exhumation. Despite not being exhumed to the
307 surface, shallow markers located at the flanks of the LC dome record isobaric
308 heating during the early stages of extension.

309 In all models, we observe that deep markers are always under partial melting

310 conditions during near-isothermal decompression and recrystallize at 0.15-0.2
311 GPa for rapid and ultra-rapid extension, and between 0.3-0.4 GPa for slow
312 models.

313 p-T paths do not show a significant dependence on the solidus of the LC. The
314 most significant difference is that for models with Solidus B, recrystallization of
315 rock domes that underwent partial melting occurs at a lower pressure of c. 0.1
316 GPa, in the stability field of andalusite.

317 4.5. Cooling rates of lower crustal rocks

318 The shape of synthetic T-t paths (Fig. D.7Supplementary Fig. 3 and 4)
319 shows a positive (and non-linear) correlation with extension rate and Moho
320 temperature, and to a lesser extent, volume of partial melting of the crust.

321 Deep dome rocks record a brief period (c. 0.5 Ma) of heating (from +50 °C
322 in the cold models, and up to +150 °C for hotter Moho temperature) before
323 experiencing rapid cooling along average cooling rates of about 2000 °C/Ma,
324 350 °C/Ma and 40-50 °C/Ma for ultra-rapid (75 mm/yr), rapid (35 mm/yr)
325 and slow (10 mm/yr) extension rates, respectively. On the other hand, mid and
326 shallow dome rocks do not experience any initial heating, and average cooling
327 occurs at 1000-1500 °C/Ma, 200-250 °C/Ma and 70-100 °C/Ma for the same
328 range of extension rates. This period of rapid cooling takes place at c. 0.7 Ma
329 (ultra-rapid), c. 1.5 Ma (rapid) and c. 4 Ma (slow) after the onset of extension
330 for deep dome rocks and after c. 0.3 (ultra-rapid), c. 0.5 Ma (rapid) and 1-2
331 Ma (slow) for shallow dome rocks. We further note that the initial period of
332 heating of deep dome rocks is not observed in the slow models with initial Moho
333 temperatures lower than 911 °C. This range of cooling rates is consistent for all
334 cold-to-warm models ($T_{Moho} \leq 911$ °C); the only significant difference between
335 them is the peak temperature during decompressional heating of the deep dome
336 rocks, which increases with the Moho temperature.

337 A significant difference in the cooling paths is observed for the hottest models
338 ($T_{Moho} = 1040$ °C). Deep crustal rocks of models employing the Solidus A
339 undergo an initial period of 3-4 Ma of considerable heating (> 150 °C) after

340 onset of extension, followed by rapid cooling at 350-400 °C/Ma. Mid and shallow
341 dome rocks record residual heating and are rapidly cooled as they are advected
342 towards shallow depths, with cooling rates of 200-475 °C/Ma (ultra-rapid), 200-
343 250 °C/Ma (rapid) and 60-100 °C/Ma (slow). On the other hand, mid and
344 shallow markers at the flanks of the LC dome barely experience any heating
345 or cooling throughout the duration of the model. Cooling paths of hot models
346 using Solidus B show a similar trend to the ones using Solidus A; however,
347 cooling of mid and shallow dome rocks occurs within a much shorter period of
348 time (near-instantaneously for rapid and ultra-rapid models).

349 *4.6. Influence of extension rate on the evolution of metamorphic complexes*

350 [Figure 5 about here.]

351 For a set of models with the same initial thermal structure, different aspects
352 such as geometry, timing of crustal break-up, volume of partial melting, p-T
353 and T-t paths of the LC and UHT conditions are affected by the extension rate
354 to different extents.

355 First, the evolution and final architecture of models with a cold initial Moho
356 temperature differ under slow, rapid and ultra-rapid extension (Fig. D.4a). In
357 slow models, the growth and exhumation of the LC dome is accompanied by
358 the ascent of the asthenosphere, followed by the break-up of the crust after sig-
359 nificant extension (> 40%) and consequent seafloor spreading, resulting in two
360 asymmetric conjugate margins. Under rapid extension, a secondary dome is
361 formed due to the asymmetric ascent of the asthenosphere, and the detachment
362 faults associated to both primary and secondary remain simultaneously active
363 until the crust is completely broken apart by the exhumation of the mantle at
364 c. 3.9 Ma, resulting in two symmetric conjugate margins. Ultra-rapid models
365 display a similar evolution to rapid models; however, these models yield two
366 asymmetric conjugate margins after crustal break-up and sea-flow spreading
367 take place (c. 3 Ma). The architecture of the crust and the dynamics of exten-
368 sion of intermediate, warm and hot models show little variability with different

369 extension rates. However, fast extension rates promote doming of the topog-
370 raphy of the Moho synchronous with the growth of the LC dome; the age of
371 break-up of the upper crust shows a negative correlation with extension rate
372 (Fig. D.7a).

373 Second, fast extension rates speed up the vertical velocity of the lower man-
374 tle, thus enhancing the amount and onset of decompression melt of astheno-
375 spheric material. Similarly, with Solidus B, crustal partial melting is enhanced
376 by fast extension. On the other hand, if Solidus A is considered, the base of
377 the crust is under partial melting conditions from the onset of extension, and
378 similar amounts of melting are predicted for all the extension rates considered
379 here.

380 Third, p-T and T-t paths exhibit a considerable sensitivity to extension rate.
381 Rocks at the core of the LC dome record near-isothermal decompression from
382 about to 0.85 GPa to 0.1-0.2 GPa during fast extension (≥ 35 mm/yr), whereas
383 they experience decompressional heating followed by cooling after reaching peak
384 conditions of c. 0.6 GPa and 750-825 °C under slow extension. Average cooling
385 rates are remarkably higher during fast extension, with maximum rates for deep
386 dome rocks of 40-50 °C/Ma under slow extension, and 1500-2000 °C/Ma for
387 ultra-rapid extension.

388 Finally, UHT conditions are only predicted for $u_{ext} \geq 35$ mm/yr and initial
389 Moho temperatures of $T_{Moho} \geq 940$ °C.

390 4.7. Influence of lower crustal melting

391 While Solidus A and B predict similar volumes of crustal partial melting
392 ($F_{max} = 25 - 35$ %) for warm and hot models, major volumetric differences are
393 observed for $T_{Moho} \leq 844$ °C. Since similar volumes of partial melting of the
394 crust are predicted for models with Solidus A and B and initial $T_{Moho} \geq 911$ °C,
395 the crust undergoes similar levels of weakening due to the presence of partial
396 melting (i.e. the viscosity fields are very similar), and thus the choice of the
397 solidus has little effect on potential geometrical discrepancies between these
398 models. In the colder models, Solidus A predicts maximum partial melting

399 fractions of at least 20%, whereas for the Solidus B, partial melting of the crust
400 only takes places at advanced stages of deformation for intermediate models
401 when hot asthenospheric material reaches the bottom of the crust. Partial
402 melting of the crust is not observed for the coldest models for Solidus B.

403 Geometrical differences linked to partial melting of the crust are only evi-
404 dent amongst models with very different volumes of partially molten crust (i.e.
405 $T_{Moho} \leq 844$ °C). A weakened base of the LC due to partial melting leads to
406 the growth and consequent exhumation of a secondary dome of LC rocks and
407 UHT conditions at the base of the crust at the stages of crustal break-up with
408 an asymmetric flow of the asthenosphere (Fig.D.6a). None of these features are
409 observed in the near-absence of partially molten crust (Fig.D.6b).

410 [Figure 6 about here.]

411 **5. Comparison of the numerical models with the Palu Metamorphic** 412 **Complex**

413 [Figure 7 about here.]

414 The geometry of the PMC is well reproduced by the exhumation of a LC dome
415 as predicted by the numerical models. The formation of several extension-driven
416 LC domes and secondary domes on the flank of the primary dome is compatible
417 with the presence of other metamorphic complexes in the region, including the
418 Gumbasa, Wana and Karossa Metamorphic Complexes, that are exposed within
419 less than 80 km south-west from the southernmost tip of the PMC (Fig. D.1b).

420 To constrain the extension rates and Moho temperatures at the onset of the
421 extensional event that drove the later exhumation of the PMC, we compare the
422 observed cooling paths from samples of magmatic and metamorphic rocks in
423 Hennig et al. (2017) with synthetic cooling paths obtained from our numerical
424 models. Analysis of the error (Appendix D) between these paths (Fig. D.8)
425 reveals that extension of the crust in this region must have taken place at very
426 fast extension rates of c. 75 mm/yr and initial Moho temperatures of c. 710
427 °C.

428 The p-T evolution in the ultra-rapid model with initial $T_{Moho} = 710$ °C
429 and Solidus A is consistent with the high temperature/low pressure mineral
430 assemblages found in the PMC: deep dome rocks are under partial melting
431 conditions and recrystallize at low pressure (c. 0.2 GPa) shortly after onset of
432 extension (c. 0.8 Ma). An emplacement depth of c. 13 km has been suggested for
433 S-type granitoids assuming a linear geothermal gradient of 30 °C/km (Hennig
434 et al., 2017). However, synthetic p-T paths for rapid and ultra-rapid extension
435 predict partial melting conditions at depths below 4-7 km for rocks comprising
436 the dome. This discrepancy can be explained by the choice of the geothermal
437 gradient used to estimate the emplacement depth: a linear geotherm of 30
438 °C/km is compatible with the crustal geotherm predicted by the numerical
439 experiments far away from the MCC, whereas the geotherm within the MCC
440 is significantly steeper and non-linear at crustal depths, thus yielding higher
441 temperatures at shallower depths.

442 Shallow LC markers ($z_{marker}(t = 0) = 19$ km) located initially at the flanks
443 of the dome experience a brief period of isothermal decompression followed by
444 isobaric- and decompression- heating. Some of the mid crustal rocks are in-
445 corporated into a secondary dome. Ascent of asthenosphere underneath the
446 secondary dome brings rocks at the Moho under UHT and partial melting con-
447 ditions for a brief period of time. Although exhumed granulite facies rocks are
448 found along the Palu-Koro Fault Zone, they pre-date the formation and ex-
449 humation of the PMC and may be associated with the rollback in the Banda
450 Sea region (Spakman and Hall, 2010).

451 Magmatic and metamorphic rocks of the PMC have been estimated to have
452 been exhumed at rates of 1-4 mm/yr (Hennig et al., 2017). However, these
453 values are small, similar to other MCC that presumably formed and exhumed
454 at slower extension rates, such as the Naxos MCC, which has an estimated
455 exhumation rate of the order of 1-10 mm/yr (Duchene et al., 2006)) and has
456 been numerically modelled using extension rates of 10 mm/yr (Huet et al.,
457 2011b). Maximum exhumation rates obtained from our models (Fig. D.8b)
458 suggest maximum values of 25-50 mm/yr for fast extension (≥ 35 mm /yr). It

459 is possible that the estimated exhumation rates for the PMC are capturing only
460 the last stages of exhumation, thus yielding lower rates.

461 **6. Discussion**

462 Even though our numerical experiments yield different types of crustal de-
463 formation, the development of LC domes is predicted for the whole range of
464 Moho temperatures considered here. These results reinforce the idea that MCC
465 are able to develop in relatively cold crust (Schenker et al., 2012). MCC in
466 relatively cold crustal conditions have been also predicted by models with an
467 inherited reversed lithological layering (Huet et al., 2011b). Rey et al. (2009)
468 pointed out that a point-like heterogeneity resulted in symmetric extension,
469 whereas a fault-like weak zone yielded asymmetric extension. Interestingly, our
470 results illustrate how a Gaussian-shaped thermal anomaly can lead to different
471 degrees of asymmetry. We argue that this discrepancy may be mainly related
472 to the use of a different rheological law (elasticity is not considered in Rey et al.
473 (2009)) and implementation of strain softening.

474 All the models described in this paper predict similar kinematics at early
475 stages of deformation, dominated by conjugate normal faults rooting in the LC.
476 Our results suggest a substantial difference between the evolution of models
477 with initial $T_{Moho} = 710$ °C and models with hotter initial conditions. In the
478 first case, extension is accompanied by the upwelling of the asthenosphere and
479 decompression melting of the mantle underneath the LC dome. Subsequently,
480 shear zones penetrate the Moho due to the embrittlement of the LC caused
481 by lower temperatures, resulting in break-up of the LC dome and mantle ex-
482 humation. In contrast, hotter crustal temperatures ($T_{Moho} \geq 844$ °C) and fast
483 extension (≥ 35 mm/yr) inhibit the transmission of stress to the Moho a more
484 vigorous ductile flow of the lower crust, maintaining a relatively flat Moho topog-
485 raphy. In these latter cases, crustal break-up is only reached after considerable
486 cooling and extreme thinning of the crust.

487 Depending on its origin, we can distinguish two kinds of MCCs: 1) a primary

488 LC dome (Fig. D.3) and 2) a secondary asthenospheric-heat induced MCC (Fig.
489 D.9). The first are common to all our numerical experiments. As concluded by
490 Huet et al. (2011b), they are driven by strain localization in the LC due to a ther-
491 mal or mechanical heterogeneity and far-field extensional forces, rather than by
492 buoyancy forces. On the other hand, fast extension rates (≤ 35 mm/yr) promote
493 an asymmetric upwelling of the asthenosphere. As the asthenosphere reaches
494 the base of the LC at the flanks of the positive topography of the Moho (offset
495 by 30-50 km with respect to the centre of the primary dome), it produces a shift
496 of the ductile flow at the base of the LC, redirecting the crustal flow towards
497 the flank of the of primary dome and ending the phase of lateral spreading of
498 the dome. Conductive heating of the bottom of the LC induced by contact
499 with hot asthenosphere material leads to UHT conditions, additional produc-
500 tion of partially molten crust, development of a secondary asthenospheric-heat
501 induced dome, and localization of strain in a new detachment fault. This kind of
502 secondary dome is only predicted for intermediate temperatures, as colder con-
503 ditions lead to crustal break-up, and hotter crustal temperatures favour ductile
504 flow and lateral migration of the LC. Asthenospheric-heat induced MCCs have
505 been previously described and compared to the Rhodope Metamorphic Complex
506 by Schenker et al. (2012).

507 p-T path diagrams of primary and secondary domes reveal different ther-
508 mal histories. As characteristic of many migmatitic MCC (Rey et al., 2011;
509 Huet et al., 2011a; Schenker et al., 2012), rocks of the primary dome are rapidly
510 advected near-vertically towards the surface and record near-isothermal decom-
511 pression to shallow depths, followed by recrystallization at low pressure and
512 rapid cooling (a small amount of heating might be possible for deep dome rocks;
513 see red markers and red p-T paths in Fig. D.8a). On the other hand, the sec-
514 ondary dome is comprised of LC rocks located beneath the upper-lower crust
515 boundary that migrate laterally and are incorporated into the secondary dome
516 (black markers in Fig. D.8a). These later rocks experience near-isobaric heat-
517 ing, induced by heat advection, at pressures of 0.4-0.5 GPa and from ~ 400 to
518 ~ 600 °C. This is followed by near-isothermal decompression to 0.1-0.2 GPa, and

519 a last phase of isobaric cooling.

520 [Figure 8 about here.]

521 Computed maximum cooling rates (Fig. D.7c) for slow models are consistent
522 with many thermochronological data from different MCC all over the world
523 (e.g. John and Howard (1995); Scott et al. (1998); Vanderhaeghe et al. (2003);
524 Yang et al. (2007)), yielding cooling rates of 70-300 °C/Ma. An increase of the
525 extension rate from slow to rapid and ultra-rapid extension results in cooling
526 rates of 700-4000 °C/Ma, an increase of almost one order of magnitude, yielding
527 cooling rates closer to the ones observed in the metamorphic complexes exhumed
528 under very rapid exhumation, as inferred for the PMC Hennig et al. (2017).
529 Exceptionally, the hottest models using Solidus B yield maximum cooling rates
530 of >5000 °C/Ma for rapid and ultra-rapid extension.

531 As obvious as it might seem, we further note that rocks located a few km
532 away with respect to the centre of the metamorphic complex (e.g. as the markers
533 located at ± 15 km) experience a slower and slightly longer period of cooling.
534 Furthermore, these rocks can experience reheating if they are reincorporated
535 into secondary domes.

536 [Figure 9 about here.]

537 7. Conclusions

- 538 • Our numerical results on extension of thinned crust with different permu-
539 tations of initial thermal structure and extension rates considered in this
540 work yield formation of lower crustal domes, suggesting that anomalously
541 elevated thermal conditions are not a pre-requisite for the formation of
542 MCCs in thinned crust.
- 543 • Three different final modes of model architecture are observed: 1) lo-
544 calised doming of the lower crust with synchronous upwelling of the as-
545 thenosphere, followed by crustal break up and resulting in two conjugate

546 margins; 2) localised doming of the lower crust accompanied by doming
547 of the Moho, occasionally followed by secondary asthenospheric-heat in-
548 duced lower crustal dome; and 3) doming of the lower crust with a flat
549 Moho.

- 550 • Two different kinds of lower crustal domes are identified: 1) a primary
551 lower crustal dome driven by far-field forces and lower crustal flow; and
552 2) a secondary asthenospheric-heat induced MCC. Rocks comprising the
553 latter domes record different p-T-t histories: primary lower crustal dome
554 rocks show very rapid vertical ascend towards the surface reflected by near-
555 isothermal p-T paths followed by rapid cooling at shallow depths, whereas
556 asthenospheric-heat induced MCC rocks experience isobaric heating at
557 intermediate pressure and 400-600 °C followed by decompression heating
558 to peak conditions (c. 0.2 GPa and 700-750 °C). These secondary domes
559 are only observed for $T_{Moho} \geq 844$ °C and rapid and ultra-rapid extension.
- 560 • Extension rate plays a crucial role in shaping the thermobarometric history
561 of MCCs, as increasing extension promotes near-isothermal decompression
562 and shorter periods of rapid cooling.
- 563 • Crystallization of the partially molten core of fast MCC in a thinned crust
564 occurs at low pressure (0.15-0.2 GPa) and intermediate pressure (0.3-0.4
565 GPa) for slow MCC.
- 566 • Rocks with a solidus corresponding to a granite are weakened by partial
567 melting even for the lowest Moho temperature considered here. Rocks
568 with a lower water content require hotter conditions for partial melting to
569 occur, and they remain stronger. The geometry is affected by the volume
570 of crustal melting; however, MCCs are still predicted in absence of partial
571 melting. This suggests that advection of lower crust towards the surface
572 is driven by the ductile flow of hot lower crust that fills the space left by
573 the stretching and break-up of the lower crust, rather than by buoyancy
574 forces alone.

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579 **Appendix A. Numerical formulation**

580 The thermo-mechanical behaviour of Earth's interior is described the equa-
 581 tions of conservation of momentum, conservation of mass, and conservation of
 582 energy, respectively:

$$\nabla \sigma = \rho \mathbf{g} \quad (\text{A.1})$$

$$\nabla \cdot \mathbf{u} = 0 \quad (\text{A.2})$$

$$\rho C_p \dot{T} = k \nabla^2 T + H \quad (\text{A.3})$$

583 where σ is the Cauchy stress tensor, ρ is the density, \mathbf{g} is the gravitational
 584 acceleration, \mathbf{u} is the velocity field, C_p is specific heat, T is temperature, κ is
 585 thermal conductivity, H is a source term that includes radioactive and shear
 586 heating. The set of equations A.2-A.3 is solved on a deformation Lagrangian
 587 mesh using a version of MILAMIN (Dabrowski et al., 2008) that has been mod-
 588 ified to include elastic deformation and non-linear rheologies, namely diffusion
 589 creep, dislocation creep and plastic deformation.

590 Viscous deformation is calculated assuming a strain rate- and temperature-
 591 dependent power law rheology for diffusion creep and dislocation creep (Karato
 592 et al., 2001):

$$\eta = \frac{1}{2} (A)^{-\frac{1}{n}} (\dot{\epsilon}_{II})^{\frac{1}{n}-1} \exp\left(\frac{E + pV_a}{nRT}\right) \quad (\text{A.4})$$

593 where the deviatoric strain rate is $\dot{\epsilon} = [1/2(\nabla \mathbf{u} + (\nabla \mathbf{u})^T) - 1/3(\nabla \cdot \mathbf{u})]$ and the
 594 sub index II denotes the square root of the J_2 invariant (i.e. $\mathbf{C}_{II} = \sqrt{(1/2)\mathbf{C} : \mathbf{C}}$
 595 and \mathbf{C} is any given tensor), A is a pre-exponential parameter, n is a power-law
 596 exponent, E is activation energy, p is pressure, V_a is activation volume and
 597 R is the universal gas constant. The second invariant of the deviatoric strain

598 rate in eq.(A.4) corresponds to the deformation of either diffusion or dislocation
 599 deviatoric strain rate tensor. A resultant composite viscosity is obtained:

$$\eta = \frac{1}{\frac{1}{\eta_{dif}} + \frac{1}{\eta_{dis}}} \quad (\text{A.5})$$

600 By adopting this composite viscous rheology, deformation is dominated by the
 601 mechanism that has the smallest activation stress.

602 Stress produced by elastic deformation is a large contributor to the global
 603 deformation-budget in the uppermost layers of the Earth, thus the necessity to
 604 add elastic deformation to the constitutive law if we aim to study the kinematics
 605 of the lithosphere. The visco-elastic constitutive law is given by:

$$\tau = 2\eta_{eff}\dot{\epsilon} + \chi\hat{\tau} \quad (\text{A.6})$$

606 with

$$\eta_{eff} = \frac{1}{\frac{1}{\eta} + \frac{1}{G\Delta t}} \quad (\text{A.7})$$

$$\chi = \frac{1}{1 + \frac{G\Delta t}{\eta}} \quad (\text{A.8})$$

$$\hat{\tau} = \tau^o + (\omega^o\tau^o - \tau^o\omega^o)\Delta t \quad (\text{A.9})$$

607 where η_{eff} is the effective visco-elastic viscosity, the super-script o refers to
 608 values at the previous time step, Δt is the time step, G is the shear modulus
 609 and $\omega = 1/2(\nabla\mathbf{u} - (\nabla\mathbf{u})^T)$ is the skew symmetric part of the velocity gradient
 610 tensor, commonly known as spin tensor.

611 Plastic deformation (i.e. fault-like behaviour of the rocks) is computed
 612 adopting the pressure dependent Drucker-Prager yield surface \mathbb{F} :

$$\mathbb{F} = \tau_{II} - p \sin(\phi) - c \cos(\phi) \quad (\text{A.10})$$

613 where τ_y is the yield stress. If the stress is such that $\mathbb{F} > 0$, the stress is brought
 614 back to the yield surface ($\mathbb{F} = 0$) using the deviatoric, corner-free and non-
 615 associative Prandtl-Reus flow rule. For points in the domain where $\mathbb{F} = 0$, the
 616 effective plastic viscosity η_{pl} is then given by:

$$\eta_{pl} = \frac{\tau_y}{2\dot{\epsilon}_{II} + \chi\hat{\tau}_{II}} \quad (\text{A.11})$$

617 The incompressible Boussinesq approximation is assumed, thus volumetric
618 strains and/or buoyancy forces due to pressure effects are not included in the
619 numerical formulation. However, density changes and buoyancy forces derived
620 from thermal expansion (and partial melting, see Appendix B) are permitted in
621 the incompressible approximation. Therefore, we use a temperature dependent
622 equation of state:

$$\rho = \rho_o (1 - \alpha(T - T_o)) \quad (\text{A.12})$$

623 where ρ_o is the reference density and α is the coefficient of thermal expansion.

624 Appendix B. Partial melting

625 The production of partial melting is calculated following Morgan (2001).
626 The mantle solidus temperature T^s is defined as:

$$T^s = T_o^s + \left(\frac{\partial T^s}{\partial p} \right)_F p + \left(\frac{\partial T^s}{\partial F} \right)_P F \quad (\text{B.1})$$

627 where T_o^s is the solidus temperature at the surface, $\partial T^s / \partial P$ is the solidus-
628 pressure gradient, $\partial T^s / \partial F$ is the solidus-depletion gradient and F is melt frac-
629 tion. Melting is produced in a parcel of the model if $T > T^s$, and two mech-
630 anisms (or their combination) are responsible for melt productivity: 1) an in-
631 crease in the material due to thermal diffusion; and 2) variations in the solidus
632 curve due to changes in pressure or melt fraction. The decompression melt
633 productivity for a lithology i within a lithology j is given by (Morgan, 2001) :

$$-\frac{\partial F_i}{\partial p} = \frac{\frac{\partial T_i^s}{\partial p} - \frac{\alpha T}{\rho C_p} + \frac{T}{C_p} \phi_j \Delta S_j \left(\frac{\partial T_i^s}{\partial p} - \frac{\partial T_j^s}{\partial p} \right)}{\frac{T}{C_p} \phi_i \Delta S_j \left(\frac{\partial T_i^s / \partial F_i}{\partial T_j^s / \partial F_j} \right) + \frac{\partial T_i^s}{\partial F_i}} \quad (\text{B.2})$$

634 where ΔS is the entropy of the solid-melt phase change, which can be related
635 to the latent heat of melting ΔH , for a pure substance, as $\Delta H = T \Delta S$. We
636 consider only a single-component melting, thus the amount of decompression
637 melting is:

$$dF_p = dP \left(-\frac{\frac{\partial T^s}{\partial p}}{\frac{\Delta H}{C_p} + \frac{\partial T^s}{\partial F}} \right) \quad (\text{B.3})$$

638 where the adiabatic term is missing because the temperatures are potential
639 temperature. The temperature change during decompression melting is given
640 by:

$$\frac{dT}{dp} = \frac{\partial T_s^i}{\partial p} + \frac{\partial T_s^i}{\partial F_i} \frac{dF_i}{dp} \quad (\text{B.4})$$

641 The amount of melt under isobaric conditions is given by:

$$dF_T = \frac{T^m - T^s}{\frac{\Delta H}{C_p} + \frac{\partial T^s}{\partial F}} \quad (\text{B.5})$$

642 where T^m is the mantle temperature and the total amount of instantaneous
643 melt is $dF = dF_p + dF_T$. The total amount of melting produced in a parcel is
644 then the summation of dF over time:

$$F = \sum_{t=1}^n dF^t \quad (\text{B.6})$$

645 where the superscript t is the time step and n is the total number of time steps.
646 For undepleted mantle, the wet solidus is used initially, and the dry solidus
647 is used after 2% melting (Braun et al., 2000). Buoyancy forces due to melt
648 production are included in the following temperature and depletion dependant
649 equation of state (EOS):

$$\rho(T, P) = \rho_o(1 + \alpha(T - T_o) - \beta F) \quad (\text{B.7})$$

650 where F is the melt fraction and β is defined as:

$$\beta = 1 - \frac{\rho_{molten}}{\rho_{solid}} \quad (\text{B.8})$$

651 where ρ_{solid} and ρ_{molten} are the reference densities of the rock in its solid and
652 molten states. We consider the density of molten crust to be $\rho = 2400 \text{ kg/m}^3$
653 and $\rho = 2900 \text{ kg/m}^3$ for molten mantle material (values taken from (Gerya and
654 Meilick, 2011)).

655 **Appendix C. Strain softening**

656 We define the accumulated plastic strain as :

$$\mathbf{E}^{pl} = \int_t \dot{\lambda} \frac{\partial \mathbf{G}}{\partial \tau_{ij}} dt \quad (\text{C.1})$$

657 where $\dot{\lambda} \geq 0$ is the so-called plastic multiplier, and $\mathbb{G} = \tau_{II}$ is the plastic
658 potential. Strain softening is then applied to the brittle domain by reducing the
659 friction angle ϕ as a linear function of the finite plastic strain, with $\phi(\mathbf{E}^{pl} =$
660 $0) = 30^\circ$ and $\phi(\mathbf{E}^{pl} \geq 1) = 15^\circ$.

661 **Appendix D. Error between synthetic and natural cooling paths**

662 To compare the synthetic and observed cooling path, we define the following
663 error

$$e = \min \left(\frac{N^{nat} - N_i^{syn}}{N^{nat}} \right) 100 \quad (\text{D.1})$$

664 where the error e is given in percentage, N^{nat} is the natural cooling rate and
665 N_i^{syn} is a vector that contains the synthetic cooling paths at the i -times.

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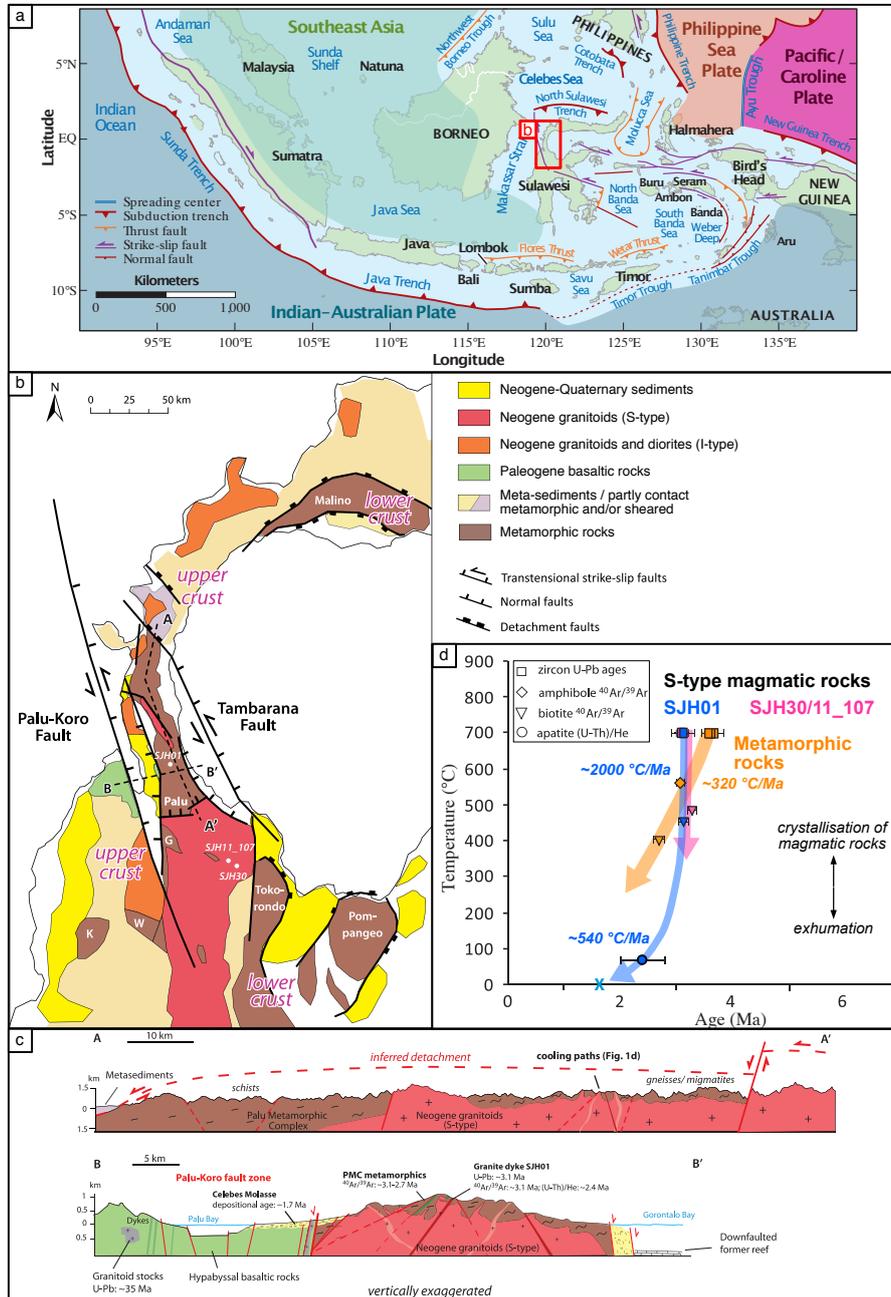


Figure D.1: **(a)** Geological map of SE Asia. The strongly coloured areas correspond to rigid plates, whereas the pale blue shaded region should be treated as micro-plates. After Hall (2017). **(b)** Simplified geological map of west Central Sulawesi showing the location of the Palu, Malino, Tokorondo, Pompangeo, Gumbasa (G), Wana (W) and Karossa (K) metamorphic complexes. **(c)** Cross-section along the A-A' and B-B' transects in panel (b). From Hennig et al. (2017). **(d)** Cooling paths of S-type granites (blue and pink) and metamorphic rocks (orange) of the Neck and mid Central Sulawesi (modified from Hennig et al. (2017)). The blue cross marks the depositional age of 1.7 Ma reported for the Celebes Molasse by van Leeuwen and Muhandjo (2005).

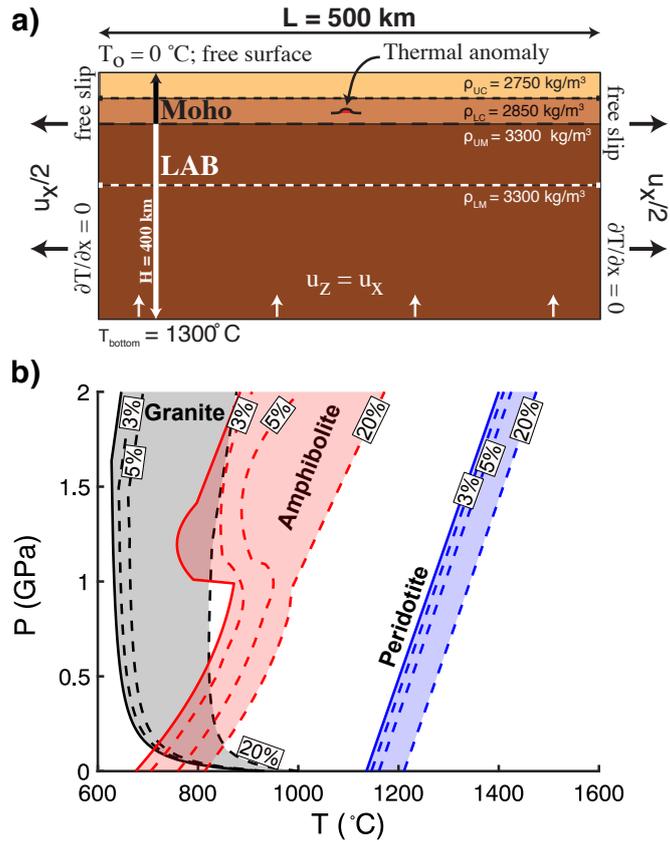


Figure D.2: **a** Model set-up. The model is stretched using pure shear boundary conditions and the temperature at the surface and bottom of the domain are fixed at a constant value. To localise the deformation in the center of the model, a thermal anomaly is introduced in the middle of the lower crust at $x = 0$ km. The domain is comprised of four laterally homogeneous rheological layers: a 17.5 km thick upper crust (UC), a 17.5 km thick lower crust (LC), a 85 km thick lithospheric mantle (LM) and a 280 km thick asthenospheric mantle (AM). The rheological parameters are given in Table 3.2. **b** Solidus of a granite with excess water (Boettcher and Wyllie, 1968) (in black), fluid-absent MORB-derived amphibolite (López and Castro, 2001) (in red) and fertile peridotite (Morgan, 2001) (in blue). Dashed lines represent the solidus at different degrees of melt fraction.

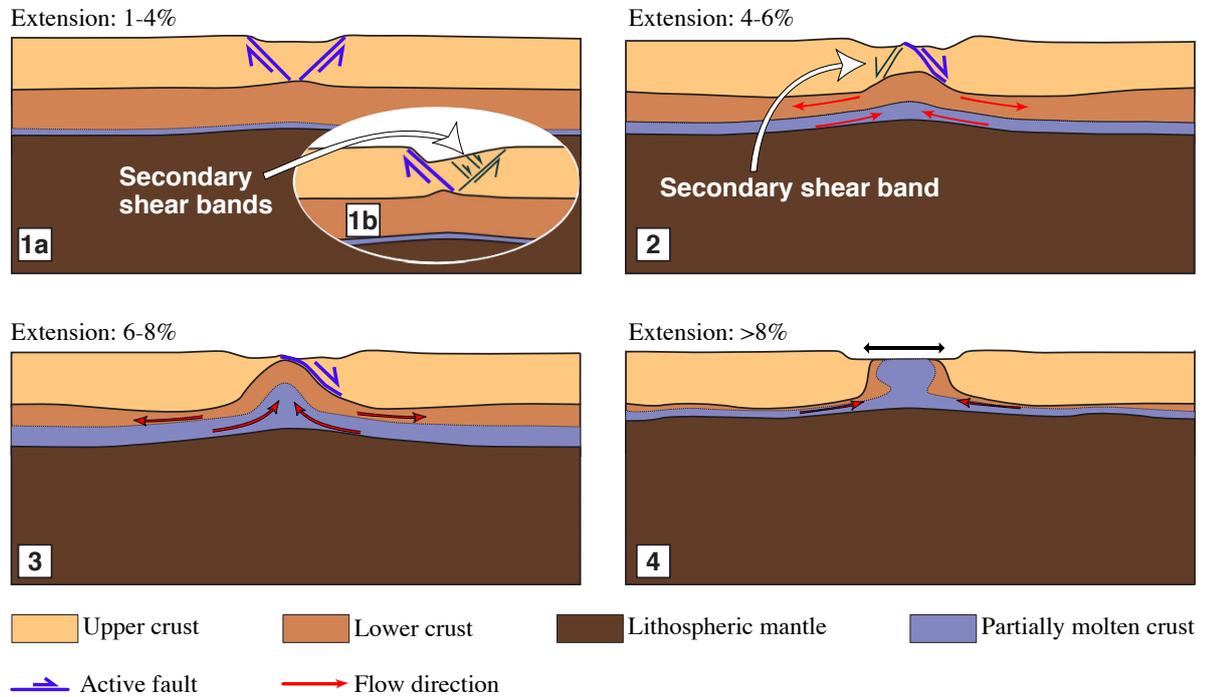


Figure D.3: Sketch illustrating the general dynamics of the formation and exhumation of lower crustal domes. **(1)** The initial stage (total extension: $\leq 4\%$) is dominated by either pure shear (under slow extension; 1a) or simple shear (under rapid and ultra-rapid extension; 1b), resulting in two conjugate normal faults that root in the LC, producing a single graben. **(2)** The conjugate normal faults are abandoned and the LC dome starts to develop. Ductile flow of basal crustal material feeds the dome, resulting in a gradual thinning of the LC at the flanks of the dome. Strain localises in a detachment fault that starts at the centre of the graben and roots at the top of the crustal dome. **(3)** The detachment fault rotates, reaching low angles close to the surface, but maintaining dips between $50\text{-}60^\circ$ at mid crustal depths. Ductile flow of the lower crust progressively fills the gap left by the stretched upper crust, and the dome is gradually exhumed along the detachment fault. **(4)** Exhumation of the lower crustal dome occurs typically after $8 - 10\%$ of extension and it is followed by a period of lateral spreading. This sketch is an idealised representation of the formation of LC domes, some features (such as the topography of the Moho) might differ depending on the initial conditions of the model.

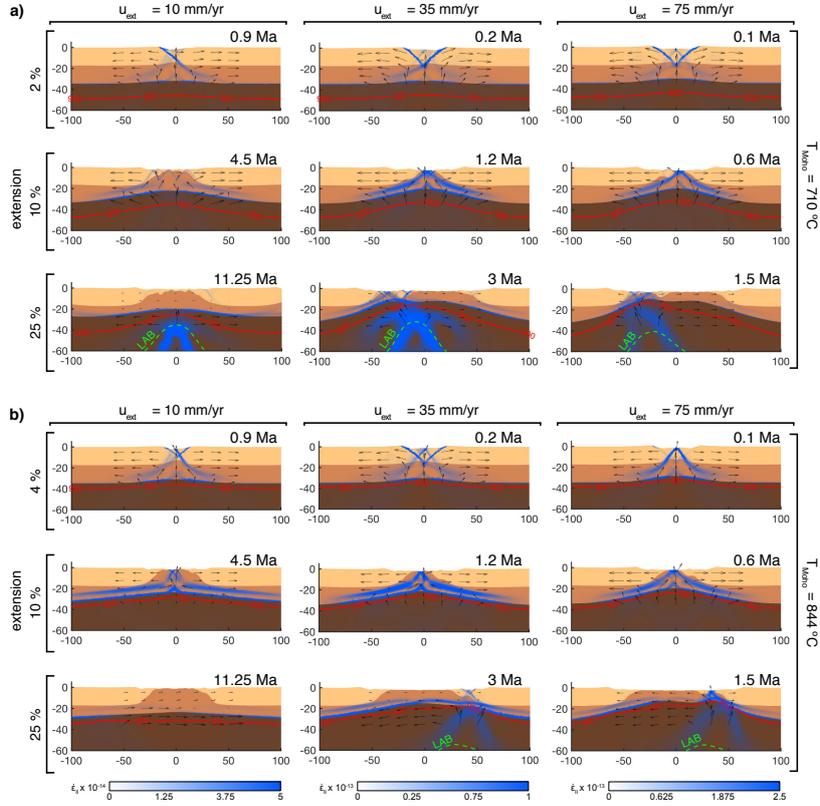


Figure D.4: Snapshots of illustrating the evolution of the (a) cold ($T_{Moho} = 710 \text{ }^\circ\text{C}$) and (b) intermediate ($T_{Moho} = 844 \text{ }^\circ\text{C}$) models for slow (10 mm/yr), rapid (35 mm/yr) and ultra-rapid (75 mm/yr) extension rates. Note that the colour map corresponds to different lithologies, and the strain rate is shown as a blue shading. The thick red line corresponds to the $900 \text{ }^\circ\text{C}$ isotherm and the green dashed line marks to the Lithosphere Asthenosphere Boundary (LAB). The black vectors represent the velocity field. These models have been computed using the Solidus A. The results of models with Solidus B are found in the supplementary material.

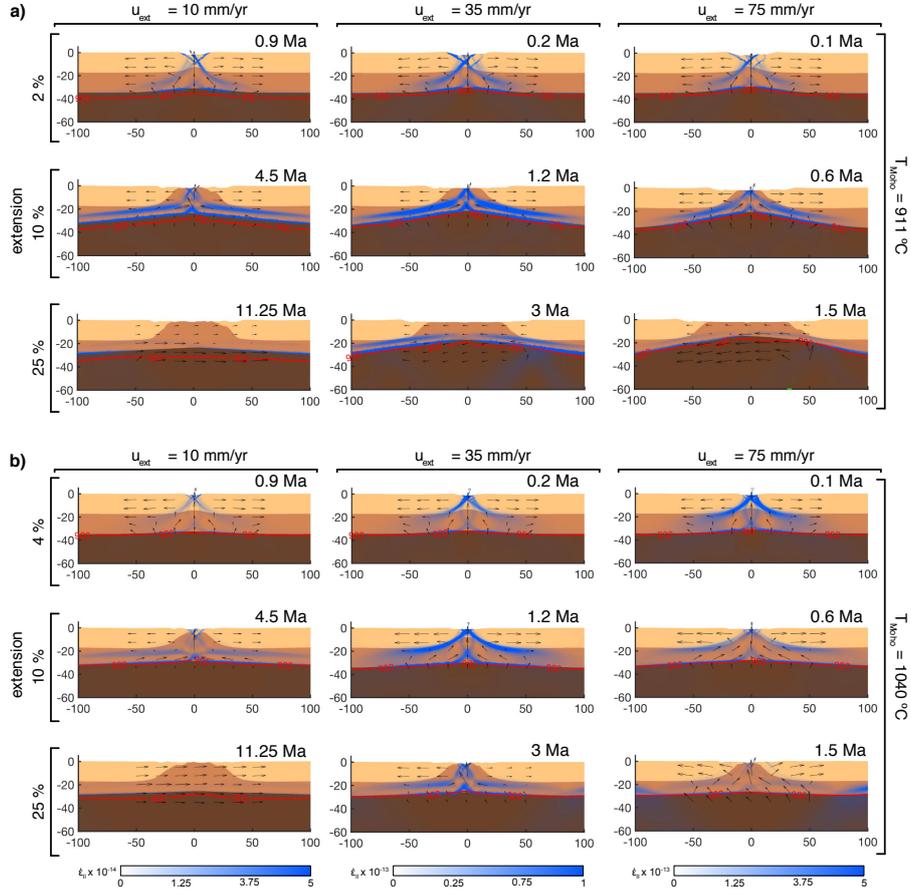


Figure D.5: Snapshots of illustrating the evolution of the (a) cold ($T_{Moho} = 911 \text{ }^{\circ}\text{C}$) and (b) intermediate ($T_{Moho} = 1040 \text{ }^{\circ}\text{C}$) models for slow (10 mm/yr), rapid (35 mm/yr) and ultra-rapid (75 mm/yr) extension rates. Note that the colour map corresponds to different lithologies, and the strain rate is shown as a blue shading. The thick red line corresponds to the $900 \text{ }^{\circ}\text{C}$ isotherm and the green dashed line marks to the Lithosphere Asthenosphere Boundary (LAB). The black vectors represent the velocity field. These models have been computed using the Solidus A. The results of models with Solidus B are found in the supplementary material.

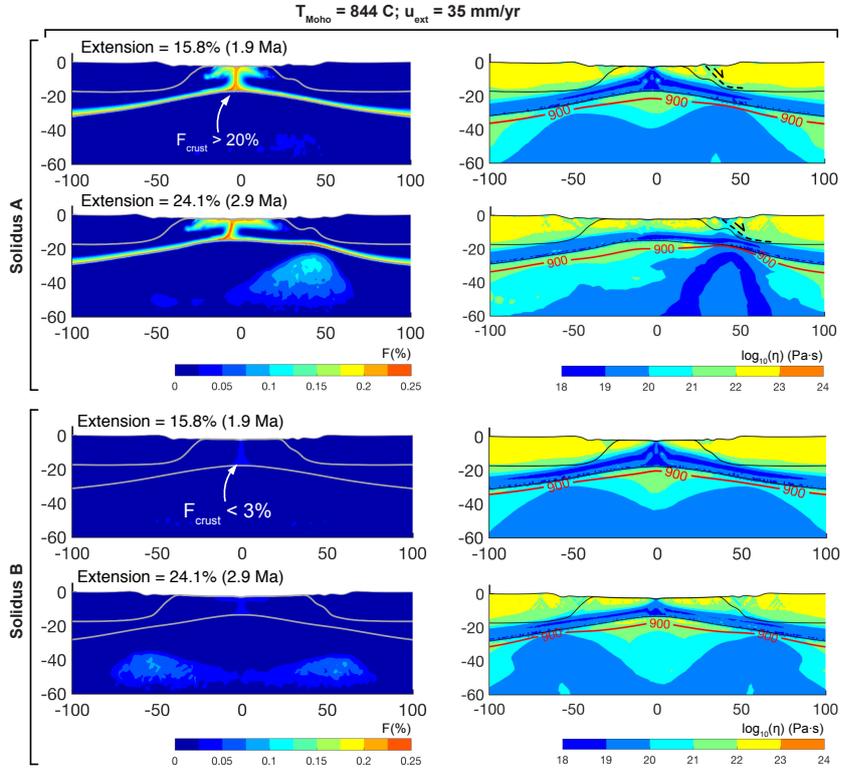


Figure D.6: Snapshots of the evolution the model with an initial intermediate temperature $T_{Moho} = 844 \text{ }^{\circ}\text{C}$ under rapid (35 mm/yr) stretching boundary conditions for models employing the Solidus A and Solidus B . The colour maps of the figures at the left-hand-side correspond to the percentage of melt fraction and the figures at the right-hand-side show the logarithmic viscosity field.

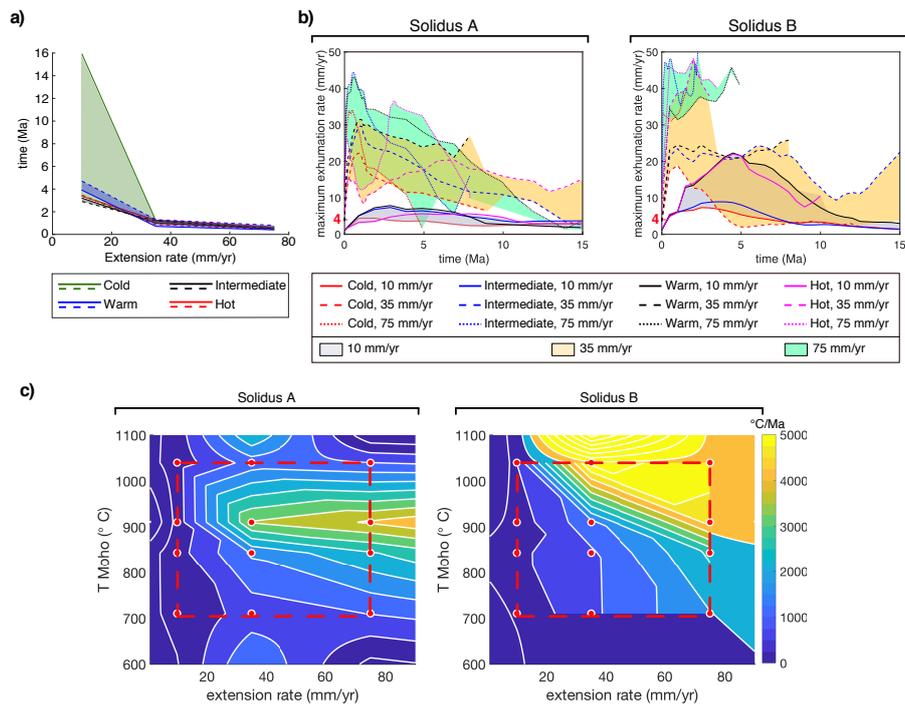


Figure D.7: **(a)** Exhumation times as a functions of extension rate and initial Moho temperature. Dashed lines and thick lines correspond to models employing Solidus A and B, respectively. **(b)** Maximum exhumation rates of the lower crust for the different combinations of extension rate, initial Moho temperature and crustal solidi. Shaded regions represent the envelope of maximum extension rates for models with equal extension rate. **(c)** Contours of the maximum cooling rates obtained from the synthetic T-t paths for models with the Solidus A and B. Red circles indicate the combination of Moho temperature and velocity chosen for the numerical models. Isolines within the red dashed rectangle are result from interpolation of data from our numerical models and isolines outside the rectangle are projected maximum cooling rate values

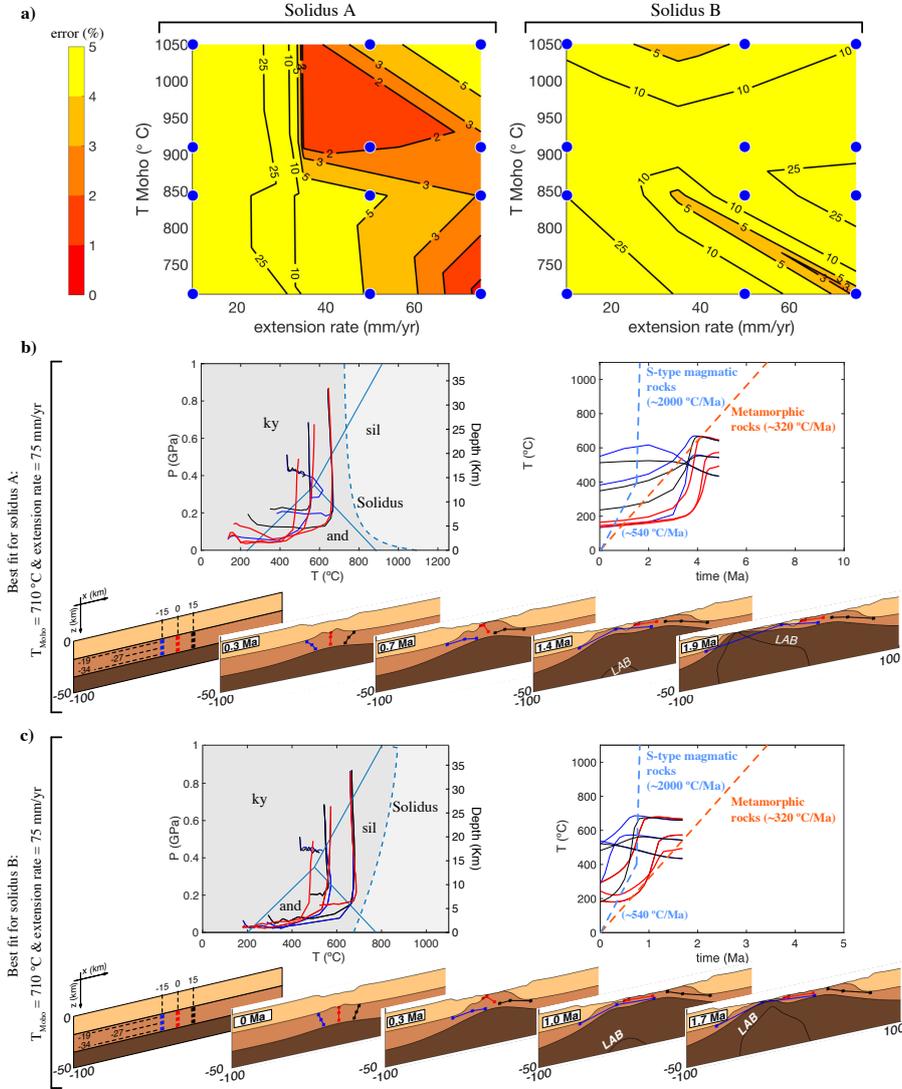


Figure D.8: (a) Isolines of the combined error of the magmatic and metamorphic natural paths of samples of the Palu MMC (Hennig et al., 2017) with respect to the synthetic T-t paths. Blue dots indicate the combinations of initial Moho temperature and extension rate used in the numerical experiments. p-T and T-t paths corresponding to the best fit of the synthetic cooling paths with the natural cooling paths for models using (b) the Solidus A and (c) Solidus B. The colour plots in (b) and (c) show the different rheological layers, and the blue, black and red dots are the markers where the pressure and temperature are tracked through time. These markers are located initially forming a grid with $x = -15, 0, 15$ km and $z = -19, -27$ and 39 km. Tracers are located at both sides of the centre of the model ($x = \pm 15$ km) to capture any asymmetry during exhumation.

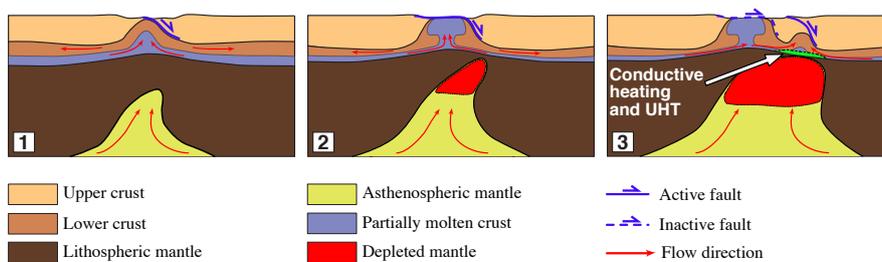


Figure D.9: Sketch showing the development of a secondary asthenospheric-heat induced core complex found for initial Moho temperatures of 710-844° and rapid and ultra-rapid extension (≥ 35 mm/yr). **1** Exhumation of the lower crust is accompanied by an asymmetric ascent of asthenospheric material. **2** The lower crustal MCC is exhumed along the detachment normal fault; the asthenospheric material continues to raise and undergoes decompression melting. **3** The asthenospheric material reaches the lower crust and conductive heating leads to UHT conditions and a peak of partial melting in the lower crust. The material in this region becomes more buoyant and is eventually exhumed along a new detachment fault.

Supplementary material

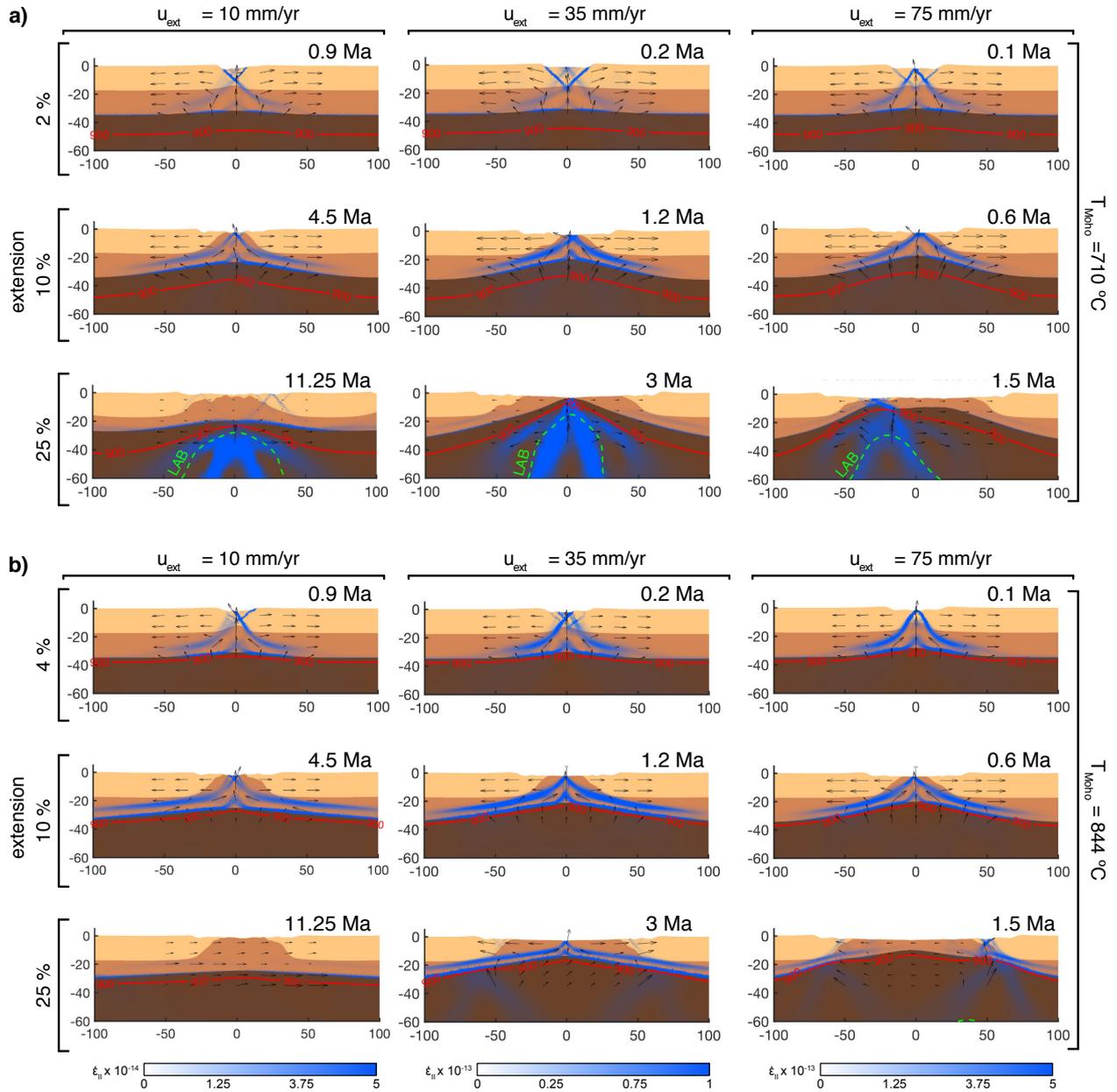


Figure 1: Snapshots of illustrating the evolution of the (a) cold ($T_{Moho} = 710 \text{ }^{\circ}\text{C}$) and (b) intermediate ($T_{Moho} = 844 \text{ }^{\circ}\text{C}$) models for slow (10 mm/yr), rapid (35 mm/yr) and ultra-rapid (75 mm/yr) extension rates. Note that the colour map corresponds to different lithologies, and the strain rate is shown as a blue shading. The thick red line corresponds to the 900 °C isotherm and the green dashed line marks to the Lithosphere Asthenosphere Boundary (LAB). The black vectors represent the velocity field. These models have been computed using the Solidus B.

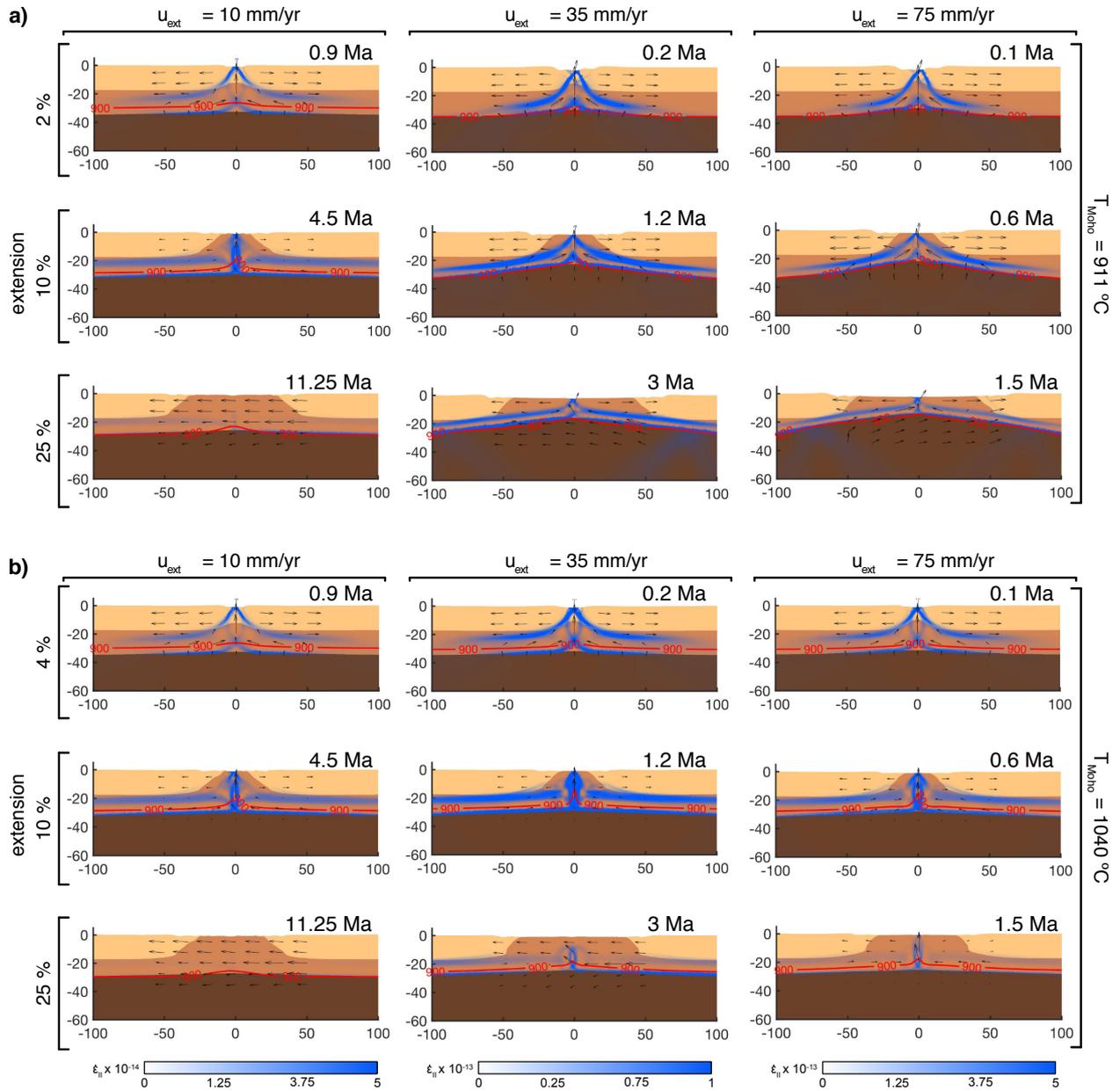


Figure 2: Snapshots of illustrating the evolution of the (a) cold ($T_{Moho} = 911 \text{ }^{\circ}\text{C}$) and (b) intermediate ($T_{Moho} = 1040 \text{ }^{\circ}\text{C}$) models for slow (10 mm/yr), rapid (35 mm/yr) and ultra-rapid (75 mm/yr) extension rates. Note that the colour map corresponds to different lithologies, and the strain rate is shown as a blue shading. The thick red line corresponds to the 900 $^{\circ}\text{C}$ isotherm and the green dashed line marks to the Lithosphere Asthenosphere Boundary (LAB). The black vectors represent the velocity field. These models have been computed using the Solidus B.

Solidus A

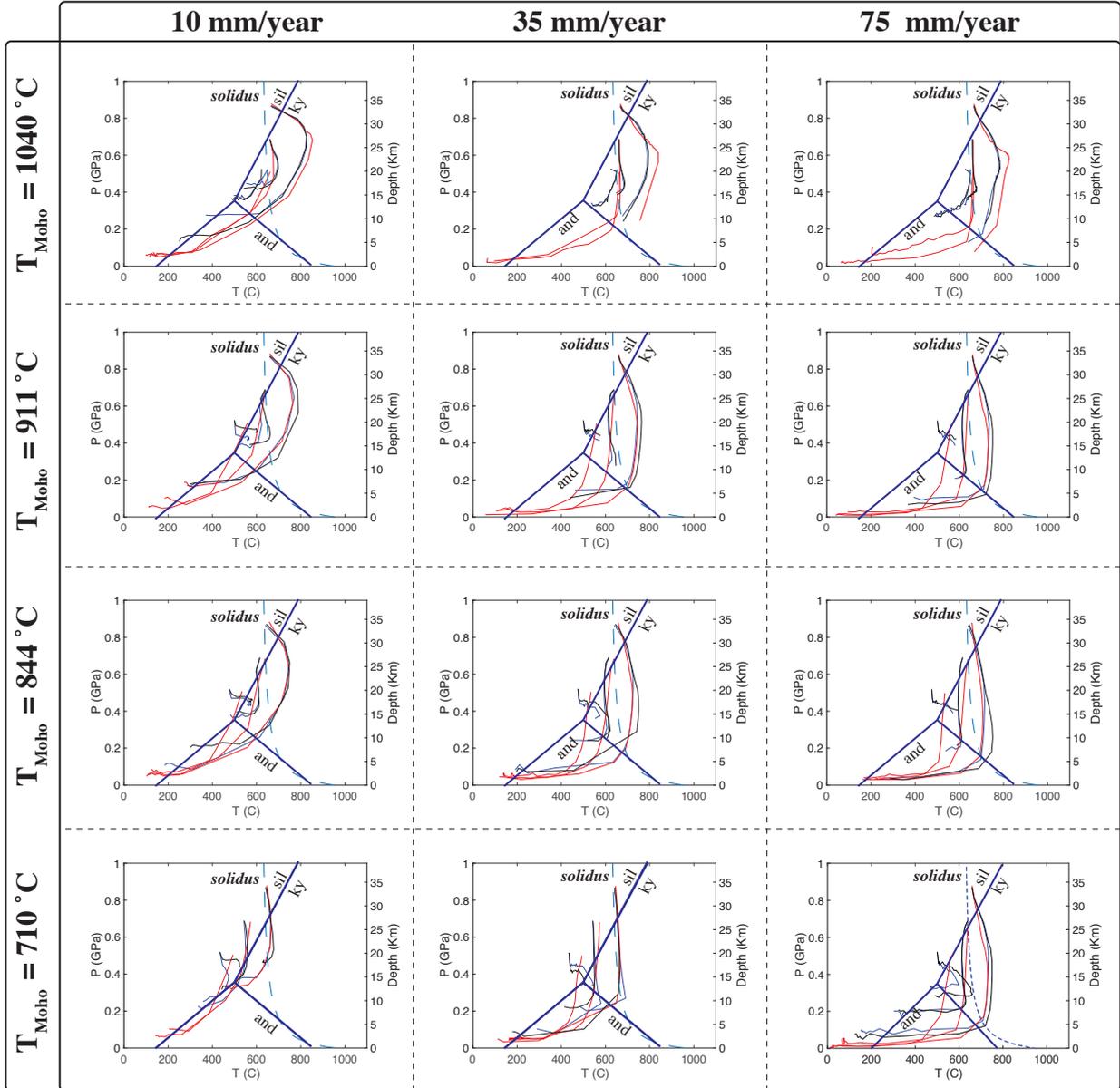


Figure 3: p-T paths corresponding to Solidus A.

Solidus B

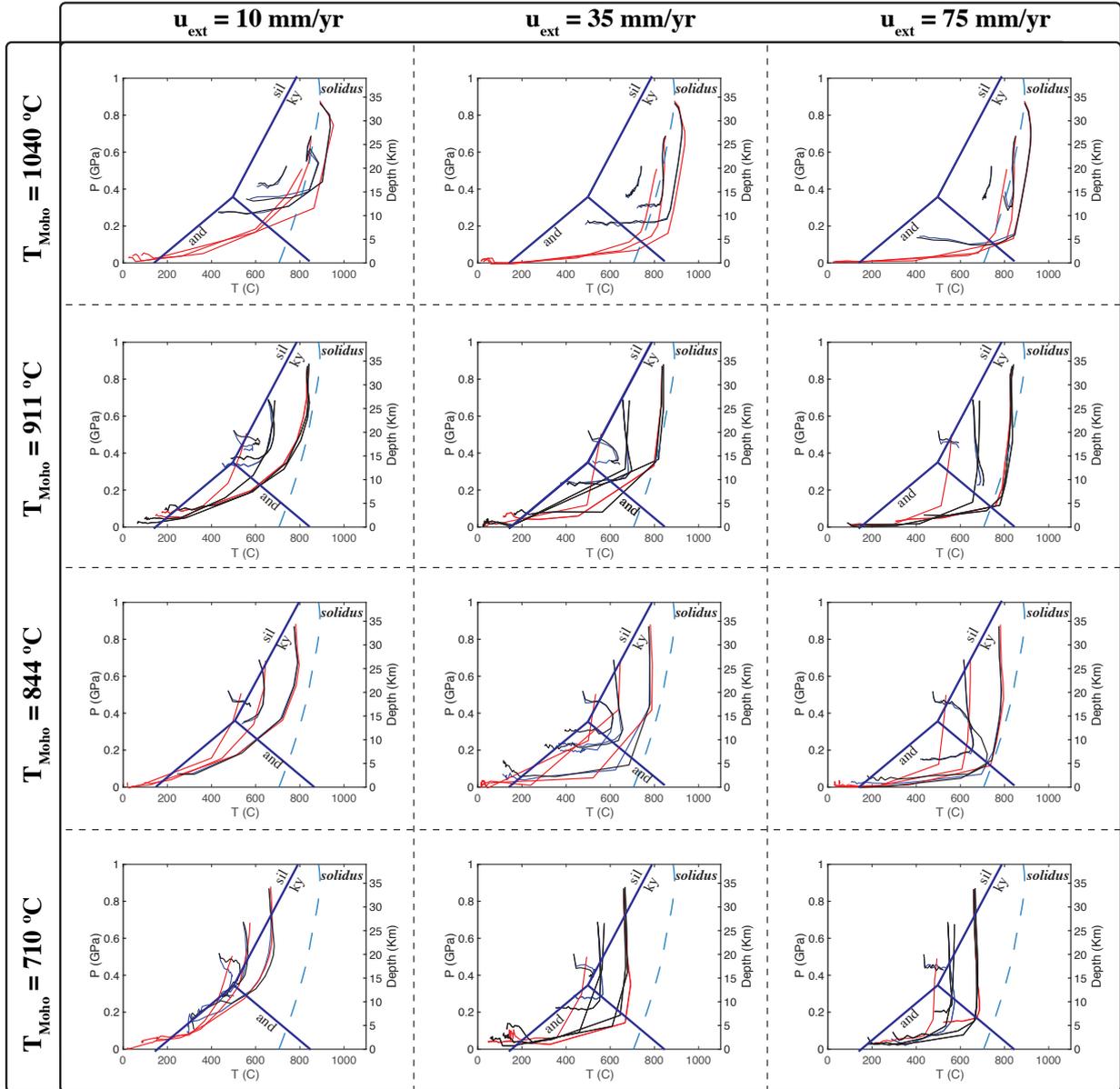


Figure 4: p-T paths corresponding to Solidus B

Solidus A

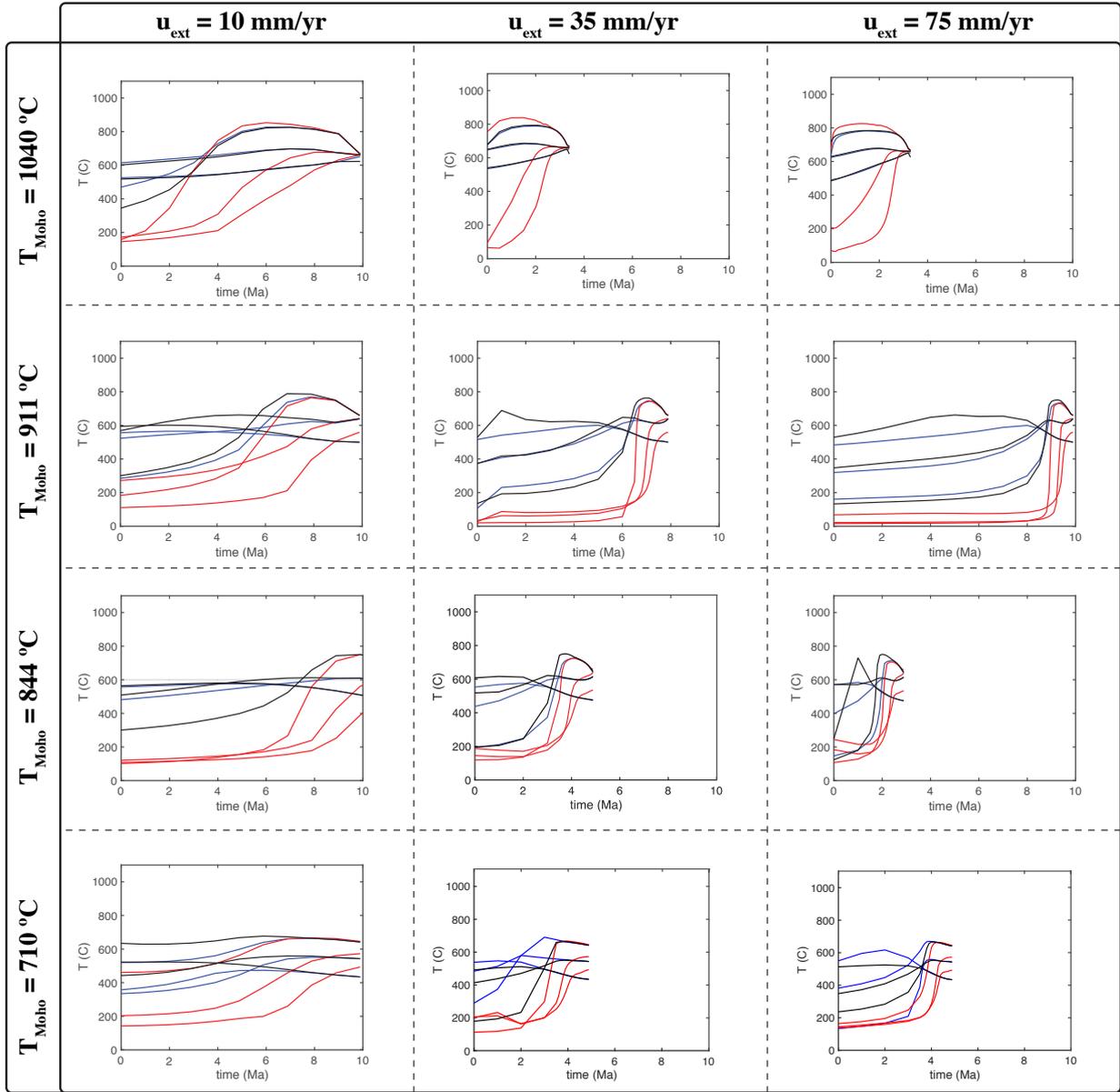


Figure 5: T-t paths corresponding to Solidus A.

Solidus B

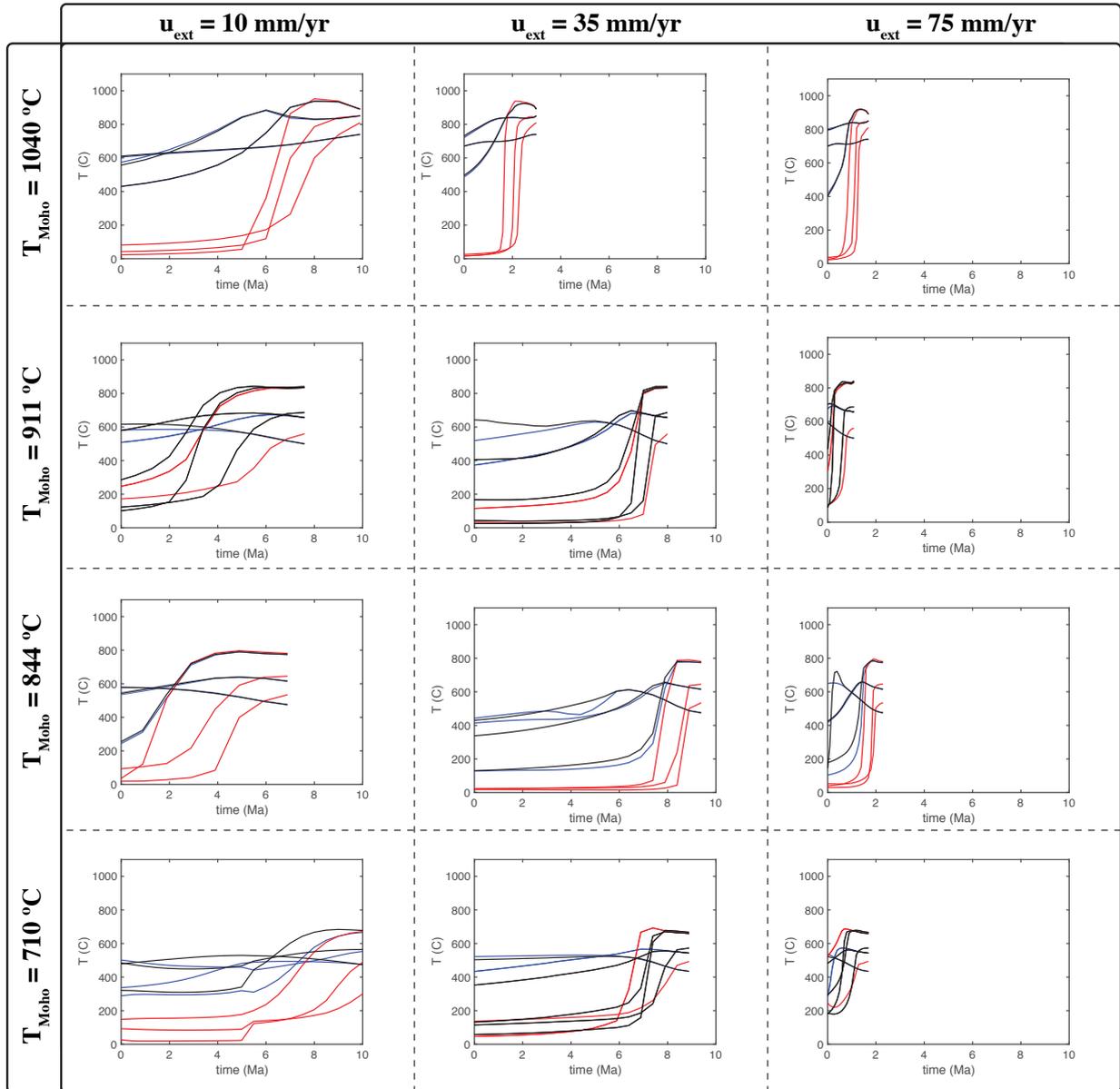


Figure 6: T-t paths corresponding to Solidus B

Solidus A

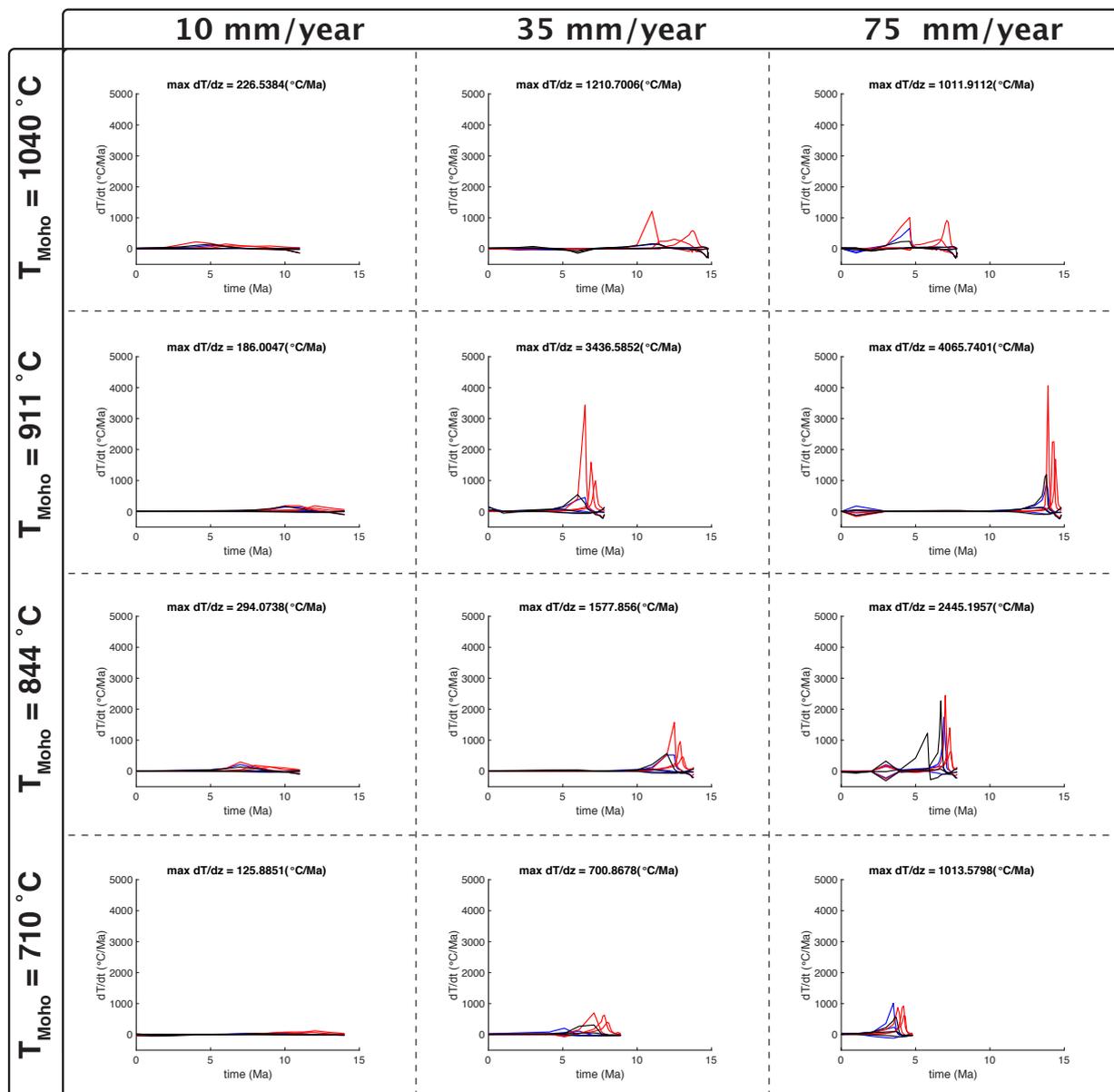


Figure 7: dT/dt paths corresponding to Solidus A.

Solidus B

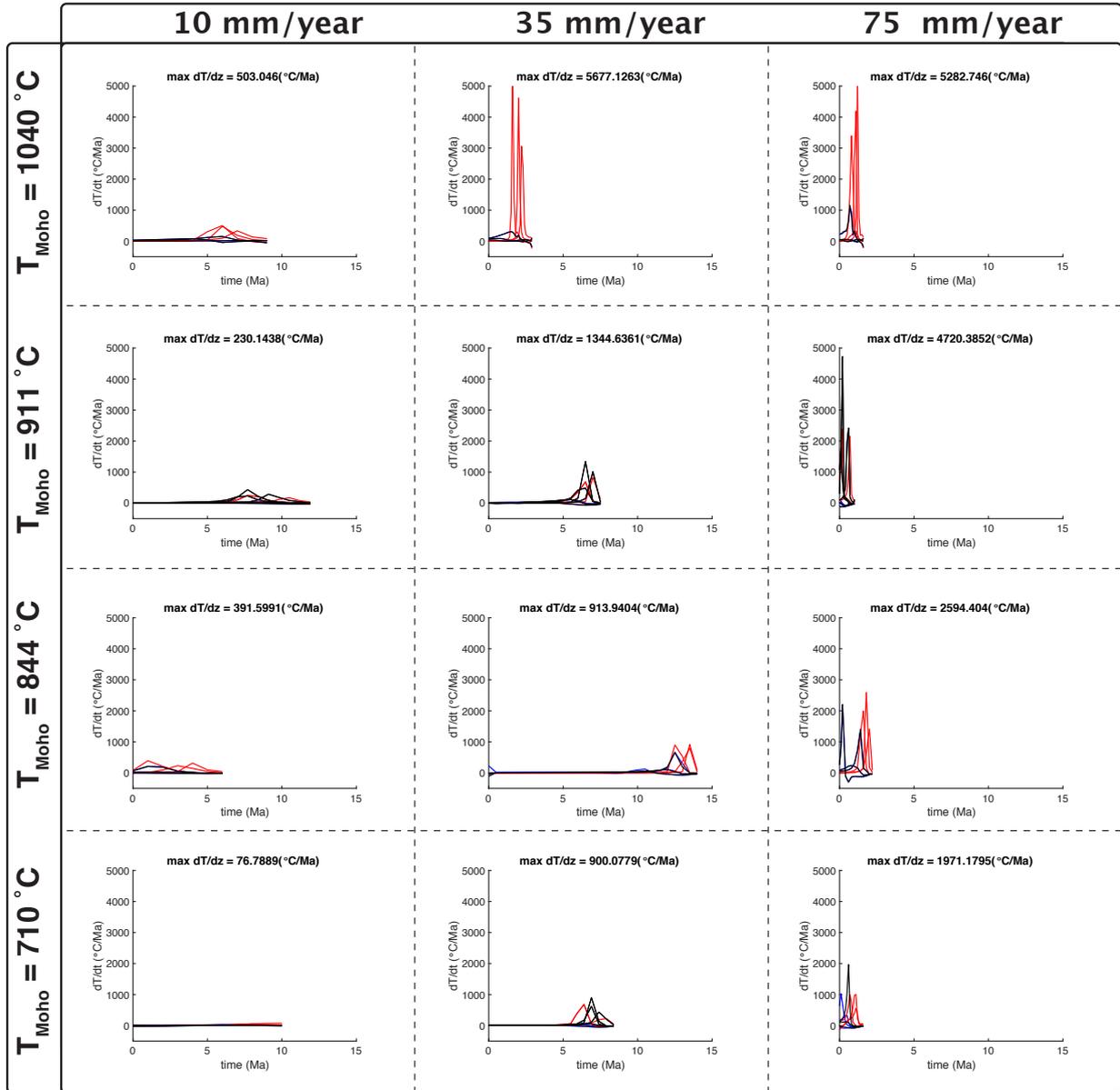


Figure 8: dT/dt paths corresponding to Solidus B

6 | Discussion

6.1 Numerical modelling of tectonic processes: critical evaluation

In Chapter 3, the formulation and benchmarks for the fully compressible code LaCoDe are presented. The compressible Stokes equations are solved in an implicit manner using the Augmented Lagrangian Method (ALM) [Rockafellar, 1974]. This approach has proven to be numerically stable and passed accurately all the benchmarks and numerical experiments in Chapters 3 and 4. In the current state of the code, all the non-linearities are treated using Picard iterations. Heister et al. [2017] showed that, using a Generalized Minimal Residual Method (GMRES) scheme, a 'more implicit' formulation requires fewer non-linear iterations, whereas a 'more explicit' formulation required fewer linear iterations because better preconditioners can be used. These authors concluded that, for stationary computations, the implicit approach is computationally cheaper and more stable; however, for time-dependent problems the explicit approach may become cheaper as, after few time steps, a good approximation of the solution from the previous time step will be available. This conclusion reaffirms the implicit formulation of the continuity equation used in LaCoDe as a valid and efficient approach to solve time-dependent problems in geodynamics. However, some considerations could be considered in the future to improve the robustness, efficiency and versatility of the code.

In the next sections I discuss some aspects of LaCoDe where there is still room for improvement, and some potential implementations that would make this code completer and more versatile.

6.1.1 Solution scheme

As mentioned above, the governing equations of a compressible visco-elastic flow are solved using the ALM, and the non-linearities that appear in the continuity equation due are treated using Picard iterations. If the rheology is non-linear (i.e. non-Newtonian and/or plastic), an outer level of Picard iterations is added to the solution scheme. This split of the non-linearities in two different Picard levels has proven to be a stable scheme for highly non-linear problems.

In order to improve the performance of the code solving non-linear problems, two options might be worth of consideration. First, as illustrated in Popov and Sobolev [2008], and discussed in further details in Spiegelman et al. [2016], a Newton-Raphson solution scheme accelerates and improves the convergence for non-linear problems that include plastic deformation. Both studies show how the Newton-Raphson scheme offers a quadratic convergence for visco-plastic problems, whereas Picard iterations show linear convergence and might stall around a non-converged solution. However, the implementation of such a method would require the calculation of the derivative of the stiffness matrix (either analytically or numerically), which would lead to a reformulation of the governing equations.

The second option consists in parallelising additional parts of the code. The most computationally expensive operations are the assembly and inversion of the block matrices that comprise the global stiffness matrix, and the Cholesky factorization of the latter. LaCoDe uses the Mutils package [Krotkiewski and Dabrowski, 2013] to perform the Cholesky factorisation, which already works in parallel. This done, a parallelisation of the matrices assembly would considerably reduce the computational cost per time step. This technique would not require a reformulation of the governing equations, but the matrix assembly algorithm should be rewritten.

6.1.2 Spatial discretisation

The spatial discretisation of the domain is carried out using the package Triangle [Shewchuk, 1996] to create a mesh of triangular elements. During the mesh generation, a set of regions with different numerical resolutions are defined, according to the degree of interest in each region. For example, in models aiming at studying continental rifting at a lithospheric scale, a fine spatial resolution is of higher necessity in the upper crust than in the deeper mantle. During the evolution of the model, the interface between different rheological layers is tracked so that the resolution after remeshing is as constant as possible.

However, a robust adaptive mesh would be a very interesting and convenient feature to be implemented in LaCoDe. This mesh generation tool consists in performing a mesh refinement in specific regions of interest, given a certain constraint or set of constraints. For example, this method can be used to create highly refined areas that follow the upwelling of a mantle plume by forcing the mesh to have a higher resolution where temperature gradient is the highest. Furthermore, it may help to reduce the mesh-dependency of localisation problems in plasticity [Zienkiewicz et al., 1995]. An adaptive mesh has already been implemented in some geodynamic codes, such as ASPECT [Kronbichler et al., 2012] or I2ELVIS [Gerya et al., 2013].

6.1.3 Rheological laws

The results of any mechanical model are ultimately defined by a common denominator: the rheology of the material. Hence the development and employability of the appropriate rheological law is of paramount importance for numerical modelling of any thermo-mechanical process. LaCoDe includes a very complete set of rheological laws aiming at mimicking the physical behaviour of rocks: 1) linear elasticity, 2) dislocation creep, 3) diffusion creep, 4) non-dilatant plasticity using a Prandtl-Reus flow rule, and 5) dilatant plasticity using an associated flow rule.

Linear elasticity is presumably the simplest rheological law, and its mathematical description and applicability to all sort of mechanical processes is far beyond any doubts. Furthermore, the elastic parameters of Earth's interior are fairly well constrained. Non-Newtonian power-laws such as dislocation and diffusion creep have been extensively studied, and the physical parameters defining them have also been fairly well constrained by triaxial laboratory experiments [Wilks and Carter, 1990; Gleason and Tullis, 1995; Hirth and Kohlstedt, 1996; Rybacki and Dresen, 2000; Hirth and Kohlstedt, 2003]. Nonetheless, one should be cautious as these experiments were performed at relatively low deviatoric stresses (the order of 100 MPa) and high strain rates (10^{-1} to $10^{-5} s^{-1}$), and yet they are used in models where deviatoric stresses can be almost an order of magnitude higher and strain rates may be several order of magnitude lower with respect to laboratory experiments. Thus, uncertainties resulting from the upscaling to geological processes should be expected.

Plasticity limits the maximum stress that a material can withstand, and it is employed to simulate brittle failure at lithospheric depths, i.e. shear zones and faults [e.g. Buck, 1991; Poliakov et al., 1994; Fullsack, 1995; Buitter et al., 2006; Gerya and Yuen, 2007; Huisman and Beaumont, 2007; Braun et al., 2008; Choi

et al., 2013; Ros et al., 2017]. As in elasticity and non-Newtonian flows, the physical plastic parameters, namely friction angle and cohesion, are also well constrained from laboratory experiments. During the calculation of the plastic strain, the effective viscosity is corrected so that the stress field is modified, and the yield stress is never exceeded. More over, plastic strain in geodynamics is often defined by a pressure-sensitive yield surface. Because the effective viscosity is updated via non-linear iterations, the pressure in each one of these iterations will not be the same. Hence plasticity introduces two non-linearities into the system: a first one due to the correction of the effective viscosity and a second one derived from the dependency of the yield criterion on the pressure. The convergence of such a highly non-linear problem is still under debate [Spiegelman et al., 2016] and further research needs to be conducted in this topic.

It is further known that plastic deformation of geomaterials is usually accompanied by a volumetric increase. Yet this feature is typically ignored in geodynamic models and has only been briefly discussed in Choi and Petersen [2015]; however, these authors focused only on the orientation of shear bands, and a deeper study of the effects of dilatant plasticity on tectonic processes is still lacking. In Chapter 4 I propose and discuss a new visco-elastic-plastic constitutive law for geodynamic models that includes volumetric plastic strains via an associative Drucker-Prager flow rule. The robustness of this model is demonstrated by its capability to yield shear bands at orientations predicted by the bifurcation study of Rudnicki and Olsson [1998], where he derived an analytical expression for shear band orientation using the Drucker-Prager's failure criterion. To test the impact of plastic dilation on tectonic processes, I designed a rift-like scenario with different values of dilatancy angle, and two end-members of lower crustal strength: 1) a weak wet quartzite lower crust, and 2) a strong mafic granulitic lower crust. The results show that even though plastic dilation is not likely to play a key role in shaping the faulting history of continental crust, it leads to a hardening of the crust that slows down the evolution of the rifted margins, and thus promotes a faster cooling of the model and yields lower values of surface heat flow. For simplicity, partial melt production was not computed in these models; nonetheless, plastic dilation might potentially have a negative feedback with partial melt occurring in the mantle, as it may enhance the cooling of the raising asthenospheric material. Whether it affects the onset and/or the amount of partial melt, needs to be further investigated.

Despite the fairly complex set of constitutive laws included in LaCoDe, I believe there is a long way to go for the geodynamic community to develop advanced rheological models to better describe the tectonic history of the Earth, and the geodynamic community would largely benefit from models and techniques that are used more frequently in engineering. Some of those that might be useful for modelling geological processes are damage models [e.g. Kachanov, 1958; Simo and Ju, 1987; Oller et al., 1995], Bingham or Perzyna visco-plastic models, viscous and elastic anisotropy/orthotropy, and anisotropic yield surfaces [e.g. Moresi and Mühlhaus, 2006].

6.1.4 Strain hardening and softening

After reaching yield conditions, a perfect plastic body will maintain a constant stress as strain increases. This behaviour is characteristic of some metals; however, after onset of inelastic deformation, geomaterials are characterised by a brief period where stress still grows with increasing strain, the so-called *strain hardening*, followed by a stress drop, referred as *strain weakening* or *strain softening*. While strain hardening does not always occur (serpentine is one example [Escartin et al., 1997]) strain weakening is a crucial mechanism to better understand the post-failure behaviour of rocks and other granular materials.

As discussed in Chapter 4, strain hardening is not implemented as an intrinsic material property; instead, the increase of volume produced by dilatant plasticity leads to a structural hardening. On the other hand, strain softening is accounted for by introducing a dependence of the plastic parameters (cohesion, friction angle

and dilatancy angle) on the softening parameter h . The plastic parameters are the friction angle, cohesion and the dilatancy angle. In order to reduce the complexity of the problem, strain softening is only applied to the friction ($\phi = \phi(h)$) and dilatancy angles ($\psi = \psi(h)$); nevertheless, it is common to find strain softening applied also to the cohesion [e.g. Huismans and Beaumont, 2007; Choi and Petersen, 2015; Tetreault and Buiter, 2017].

A good parametrisation of the strain softening behaviour is important to reproduce the post-peak behaviour of different materials, and while there is an extensive amount of studies addressing the topic [Sibson, 1990; Rice, 1992; Ridley, 1993; Streit, 1997; Ingebritsen and Manning, 1999; Bos and Spiers, 2002; Handy and Stünitz, 2002], more laboratory experiments are needed in order to further constrain the physical parameters defining the strain softening curve.

The implementation of strain softening also introduces interesting numerical complications. Let us consider a rectangular domain subject to tensile stress at its edges and with a small mechanically weaker region located at its centre. When stresses reach the weaker region, softening begins, stress drops, and the material around this region must unload elastically. As a consequence, the softening region is defined by the size of the region with the minimum strength. This leads to two major drawbacks: 1) to solve the Stokes equations, the domain is subdivided in cells or elements with a finite size; consequently the size of these elements will define the strain softening regions and influence the solution; and, 2) strain softening leads to the loss of ellipticity of the governing equations and the boundary problem becomes ill-posed, i.e. an infinitesimal change in the stiffness matrix leads to a large change of the solution.

To my knowledge, the only solution applied to geodynamics was given by Lavier et al. [2000], who proposed the addition of a mesh-independent length scale to define the area of influence of strain softening. In Chapter 4 I have further discussed alternative solutions to the mesh dependency such as non-local continuum techniques. As remarked in Section 6.1.3, modelling of geological processes will largely benefit from importing some of these solutions common in engineering.

6.1.5 Dilatancy angle

During tectonic processes, the strain of the cold lithospheric rocks can go far beyond peak conditions. Therefore, having a model that captures the whole stress-strain post-peak curve will help to produce more realistic results. The post-failure strength behaviour is controlled by the strain softening, discussed in Section 6.1.4, and the dilatancy angle. Since, in rock mechanics, many problems are solved by avoiding failure, the dilatancy angle has not received a great deal of attention and, when used, it has been frequently considered in a simple way either as an associated flow rule ($\psi = \phi$) or fully non-associated ($\psi = 0^\circ$) [Alejano and Alonso, 2005]. However, Vermeer and De Borst [1984] pointed out that an associated flow does not represent well the post-peak behaviour for soils, rocks and concrete and concluded that, for rocks, the dilatancy angle had to be $\psi \leq \phi - 20^\circ$. Later on, Detournay [1986] noted the inappropriateness of this approach and proposed a dilatancy angle that depended on the shear plastic strain. Additionally, Medhurst [1997] and Ribacchi [2000] found out that the dilatancy angle further depends on the confining stress.

One could attempt to reproduce features of post-peak curves mentioned above by writing complex expressions for the dilatancy angle [e.g. Alejano and Alonso, 2005], so that the results would fit the stress-strain curves obtained from laboratory experiments. However, the level of uncertainty surrounding the modelling of tectonic processes is already very elevated and attempting to upscale every single detail featured in results derived from laboratory experiments might be unnecessary. In Chapter 4 I attempted to capture the main features previously mention in a simple manner by introducing a linear dependency of the dilatancy angle on the softening parameter h and on the depth to approximate the pressure-dependency. Interestingly,

results with different degrees of the depth dependency yielded different faulting histories. Thus, further investigations are required in order to better constrain the depth dependency of the dilatancy angle.

6.1.6 Serpentinisation and magmatism

The rationale behind the implementation of a fully consistent compressible formulation for viscous flows was not only to include volumetric strain derived from pressure and temperature changes with respect to the reference state (which has obvious implications in, for example, mantle convection and subducting slabs), but also to account for volumetric changes linked to phase changes (e.g. expansion caused by serpentinisation reactions during mantle exhumation in magma-poor conjugate margins, or during slab fault bending; or contraction caused by deserpentinisation reactions or removal of mass in the form of fluid or melt).

Serpentinisation

Serpentinisation reaction of peridotites is an important process involved in both rifting of continental crust and slab subduction that produces an increase of volume [Gresens, 1967; O'Hanley, 1992]. The volumetric strain is known to reduce the crustal strength [Escartin et al., 1997, 2001] and plays an important role during continental break-up by controlling part of the formation of continent-ocean transition and oceanisation. It has also been hypothesised that the volume increase linked to serpentinisation of the subducting slab enhances fault-bending [Phipps Morgan, 2001a], hence aiding the process of subduction initiation.

As described in Section 2.4.2, I have implemented parametrisation of serpentinisation and deserpentinisation reactions based on the work of Malvoisin et al. [2012]. For completeness, this should be coupled with the Darcy equations to properly quantify the amount of water circulating within the lithosphere and also the spatial distribution with time. Unfortunately, during my Ph.D. I did not further investigate the implications of volumetric changes in geodynamic models. Development and design of models aiming at studying these processes, along with the coupling the equation of flow in a porous media, remains one of the major topics to be further studied with LaCoDe.

Magmatism

The buoyancy of migmatites is believed to have an active role in exhuming mid-crustal rocks and in supporting vertical motion in metamorphic core complexes [e.g. Gautier et al., 1990; Burg and Vanderhaeghe, 1993; Ledru et al., 2001]. Furthermore, partial melting of the mid-lower crust is likely to promote strain localisation as the viscosity lowers (if one assumes that the partially molten crust moves *en mass* along with the non-molten material). Hence, there has been the argument around whether magmatism promotes and facilitates the formation of core complexes, or whether these are a result of extension [Gans, 1989; Lee Armstrong and Ward, 1991; Lister and Baldwin, 1993; Spencer et al., 1995; Foster et al., 2001; Tucholke et al., 2008; Olive et al., 2010].

In Chapter 5 I have tested the influence of melting in the formation of metamorphic complexes by using solidus curves corresponding to 1) an hydrated granite [Boettcher and Wyllie, 1968] and 2) a fluid-absent MORB-derived amphibolite [López and Castro, 2001]; see Fig. 2.5 for a representation of the correspondent solidus curves. Models using the first solidus curve to define crustal melting produce a higher amount of partial melt at mid-lower crustal depths as melting occurs at 600-700 °C at 1 GPa, whereas the fluid-absent amphibolite requires temperatures of around 900 °C at similar pressures. It is therefore easier to produce partial melting in the crust if the source is a hydrated MORB and it will lead to higher amounts of partial

melt volume. However, the results shown in Chapter 5 suggest that the amount of partial melting does not play a crucial role in the development of the metamorphic complex. The appreciable effect of the amount of melting is observed in the p-T paths, where at very high geotherms, the amphibolite leads to near-isothermal decompression of lower crustal rocks of the metamorphic complex. These results are in agreement with the observations of Rey et al. [2009b], who concluded that metamorphic complexes develop both in the presence and absence of crustal melting.

6.1.7 Weak Seeds

Tectonic models are almost always designed so that the rheological layers are laterally homogeneous. Therefore, upon prescription of the boundary conditions, strain is likely to localise at a non-desired region. For example, in models of stretching of continental lithosphere where the domain is laterally homogeneous, strain might localise at geometrical singularities, such as the corners of the model. A simple solution is to introduce the so-called *weak seed*, which consists of a thermo-mechanical heterogeneity. This artificial weakness is commonly used in tectonic models to enforce strain localisation within a desired region. In rifting models, these heterogeneities are typically small regions where the temperature is anomalously higher [e.g. Brune et al., 2012; Ros et al., 2017] or plastic parameters are lower than the surrounding material [e.g. Buck, 1991; Huisman and Beaumont, 2007; Tetreault and Buitert, 2017]. On the other hand, in models aiming at studying subduction zones, it is common to define a weak zone between the subducting slab and the overriding plate [e.g. Gerya and Yuen, 2003; Warren et al., 2008; Taramón et al., 2015]. I must note that one must be careful when introducing mechanical heterogeneities in any model, as the nature and location of the weak seed might have a large influence on the results [Dyksterhuis et al., 2007].

Of particular interest are the weak seeds used to reproduce the development of metamorphic core complexes. A frequent way to force the models to produce metamorphic core complexes is by imposing a fault-like region of low viscosity or plastic parameters that cuts through the crust [e.g. Lavier and Buck, 2002; Rey et al., 2009b,a; Whitney et al., 2013]. Contrary to this trend, other authors choose not to prescribe a weak seed [Schenker et al., 2012]. While the first approach assumes the pre-existence of a normal fault cutting through the crust is a condition for the development of metamorphic complexes, it is not clear the role of boundary effects in the absence of a weak seed. Due to the uncertainties around the hypotheses adopted by these two approaches, in Chapter 5 I decided to use a Gaussian-shaped perturbation of the thermal field located at mid-crustal depths and found that, when combined with a weak and ductile lower crust, it results in the formation and exhumation of lower crustal domes that resembles the metamorphic complexes. In this way, the metamorphic complex develops without the necessity of a pre-existent fault and boundary effects are avoided.

6.2 Future work

Several questions arise from the work presented in this Ph.D. thesis, some related to numerical and/or theoretical concepts, and some others related geological processes that have not been addressed throughout this thesis.

Many of the technical aspects to be considered have already been discussed in Section 6.1, such as the implementation of a Newton-like solution scheme and an adaptive mesh, or importing engineering solutions to overcome the issue of a mesh-dependent solution in strain localisation problems. However, the implementation of the fully compressible Stokes equations opens the doors to investigate a wide spectrum

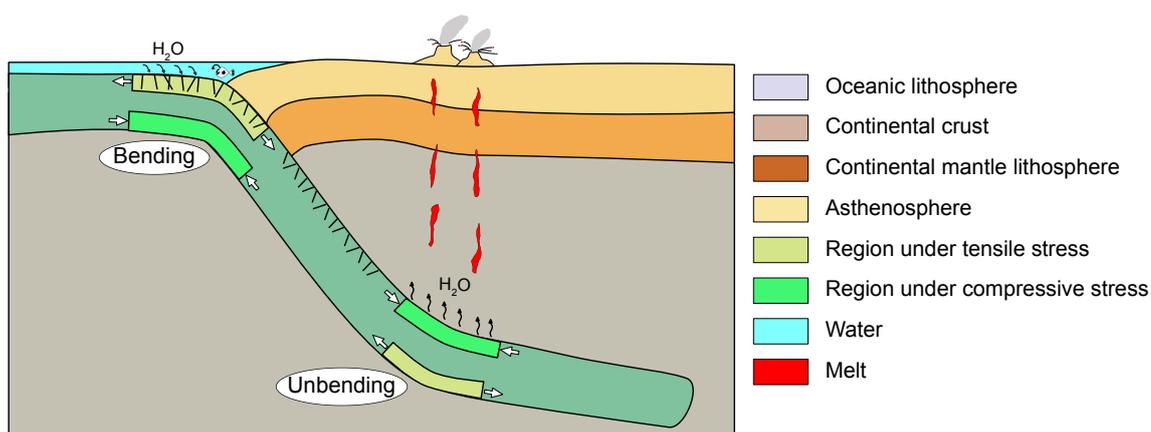


Fig. 6.1 Schematic representation of bending (and consequent faulting) and unbending of the subducted slab due to hydration and dehydration reactions.

of geological processes that, to my knowledge, have not been studied from a numerical point of view, or received very little attention: the study of the mechanical effects of volume-changing processes.

One of these processes are the consequences of serpentinisation and deserpentinisation reactions on different tectonic processes. In Section 6.1.6 I have briefly discussed its potential effects on rifted margins and aiding subduction initiation. However, serpentinisation may play an important role in other geological scenarios. For example, dehydration of the subducted slab would lead to deserpentinisation reactions, reverting the increase of volume produced at shallow depths. This would produce a contraction at the upper *face* of the subducted slab, potentially forcing it to unbend.

Another scenario related to this process are transform faults in rifted margins, i.e. East Pacific and Atlantic Ocean. Cross-sections of transform faults typically display a narrow area at the centre that is topographically depressed (up to a few hundred meters) with respect to the surrounding region and surrounded by two elevated bulges at its sides. Furthermore, the blocks at each side of the valley have different depths. The differential depth of the blocks can easily be explained by the offset of their thermal age. On the other hand, the presence of the valley and the bulges supports the idea of the existence of some tectonic activity, as one of the blocks is likely to have rotated due to flexure caused by extension.

A possible driver mechanism for extension is the presence of a mechanically weaker region between the blocks of the transform fault, caused by serpentinisation reactions. Cooling down of the lithosphere will lead to a differential thermal subsidence at both sides of the transform fault, and the weak region will promote strain localisation and extension in the surroundings of the fault. This hypothesis could easily be tested using the current state of LaCoDe and will be addressed in the near future. Furthermore, mid-ocean-ridges are characterised by several transform faults, making this a highly three-dimensional process. The interaction between adjacent transform faults and the degree of obliquity of extension remains to be further investigated using a three-dimensional approach.

These geological scenarios are only some of examples where volume changes can become relevant and should not be neglected if one aims at studying geological processes with the maximum possible degree of complexity. Unfortunately, this has not been possible during my Ph.D. and will be the subject of future research using the numerical code developed here.

In Chapter 5 I have shown how the high cooling and exhumation rates of metamorphic complexes found in Sulawesi can be explained by extension rates that are higher with respect to the ones that lead to the development of other extensively studied metamorphic complexes. Furthermore, these anomalously rapid

exhumation processes are not endemic of Sulawesi, but they are likely to be widespread along the Banda Sea region. Recent studies suggests similar or even higher rates [Pownall et al., 2013, 2014; Pownall, 2015] for metamorphic complexes found Seram (located at the Banda Sea, between Sulawesi and West Papua New Guinea) which are age-comparable to the events in Sulawesi. Therefore, more numerical (and field) studies are required to constrain the cause of such rapid process and verify whether they are all related to the same tectonic event (e.g. roll-back of the Banda trench and consequent opening of the Banda Sea) or they are independent.

7 | Conclusions

7.1 Conclusions

- The solvability and accuracy of a general expression of the compressible Stokes equations using an iterative Augmented Lagrangian Method using an implicit approach has been demonstrated. This approach has passed several benchmarks for elastic and viscous deformation with errors comparable to other available codes.
- The density dependence of the continuity equation introduces a non-linearity into the problem that is lacking in the incompressible approximations. Hence the total number of iterations per time step will be substantially more elevated. I find that for non-linear rheologies, one could treat all the non-linearities in one single level of Picard iterations. However, as the complexity of the model increases, it is more convenient to split the non-linearities arising from the rheology and the continuity equation into two levels of Picard iterations, leading to faster convergence rates and better resolved solutions.
- While the Boussinesq approximation is a valid hypothesis for simple modelling of crustal deformation, more complex models that aim to study processes such as phase changes or partial melting will require a modification of the Boussinesq approximation to accommodate the effects of volumetric strains and volume-change-linked stresses.
- The inclusion of a self-consistent volume change source term is a powerful tool that opens an opportunity to study the effects of overpressure caused by the inflow and outflow of mass into geological features (e.g. serpentinisation and melt extraction).
- Plastic models with a Drucker-Prager yield function and an associative flow rule yield shear bands dipping at Coulomb angles. On the other hand, a non-associative deviatoric Prandtl-Reus flow rule yields shear band orientations scattered within the range of Coulomb-Arthur-Roscoe angles.
- Volumetric expansion within the shear bands results in a structural hardening of the domain. However, the evolution and final geometry of the rift do not present significant differences between dilatant and non-dilatant models.
- The slower vertical motion of mantle upwelling induced by the structural hardening of the lithosphere promotes a cooling of the models, whereas volumetric plastic strain introduces an extra source of heat derived from inelastic work. The thermal structure and surface heat flow in dilatant models is therefore altered with respect non-dilatant models. For a strong lower crust, brittle deformation is highly effective and dilatant models yield similar temperature fields for the range of dilatancy angles considered in this paper. However, plasticity is less effective in models with a weak lower crust and shear heating dominates for $\psi_0 = 15^\circ$, whereas the model with $\psi_0 = 30^\circ$ is cooler, due to the slower upwelling of mantle material.

- The evolution of the faults and shear zones of a model with a depth-dependent dilatancy angle with $\psi_0 = \phi_0$ at null confining pressure represents an intermediate state between using constant $\psi_0 = 30^\circ$ and $\psi_0 = 15^\circ$. During the early stages of extension, this model is very similar to a model with constant $\psi_0 = 30^\circ$ and it evolves towards a similar strain rate state as the model with $\psi_0 = 15^\circ$ as strain softening kicks in.
- Our numerical results on extension of thinned crust with different permutations of initial thermal structure and extension rates considered in this work yield formation of lower crustal domes, suggesting that anomalously elevated thermal conditions are not a pre-requisite for the formation of metamorphic core complexes (MCCs) in thinned crust.
- Three different final modes of model architecture are observed: 1) localised doming of the lower crust with synchronous upwelling of the asthenosphere, followed by crustal break up and resulting in two conjugate margins; 2) localised doming of the lower crust accompanied by doming of the Moho, occasionally followed by secondary asthenospheric-heat induced lower crustal dome; and 3) doming of the lower crust and a flat Moho.
- Two different kinds of lower crustal domes are identified in intermediate temperature ($T_{Moho} = 844^\circ\text{C}$) models under fast extension (≥ 35 mm/yr): 1) a primary lower crustal dome driven by far-field forces and lower crustal flow; and 2) a secondary asthenospheric-heat induced MCC. Rocks comprising these lower crustal domes record different p-T-t histories: primary lower crustal dome rocks show very rapid vertical ascent towards the surface reflected by near-isothermal p-T paths followed by rapid cooling at shallow depths, whereas asthenospheric-heat induced MCC rocks experience isobaric heating at intermediate pressure and 400-600 °C followed by decompression heating to peak conditions (c. 0.2 GPa and 700-750 °C).
- Extension rate defines the shape of the p-T and T-t paths of lower crustal rocks of MCC, promoting near-isothermal decompression and shorter periods of rapid cooling. The final architecture of the models and volume of partial melt is also affected by the extension rate. This dependence on extension rate is reduced with hotter initial conditions.
- Rocks with a solidus similar to a hydrated MORB are weakened by partial melt even for the lowest Moho temperature considered here. Rocks with a lower water content require hotter conditions in order to be partially molten, and they remain stronger. The geometry is affected by the volume of crustal melting, however, MCCs are still predicted when there is no partial melting of the crust. This suggests that advection of lower crust towards the surface is driven by the ductile flow of hot lower crust that fills the space left by the stretching and break-up of the lower crust, rather than by buoyancy forces alone.

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A | Appendix

A.1 The Finite Element Method

The description of the FEM presented here is based on Hughes [1987]. I will adopt Einstein index notation (i.e. repeated indices indicates the sum: $u_{ii} = u_{11} + u_{22} + u_{33}$), and the following notation for partial derivatives:

$$u_{i,j} = \frac{\partial u_i}{\partial x_j} \quad (\text{A.1})$$

so that $u_{i,j}$ denotes the derivative of the i -component of u respect to j . When appropriate, I will also use a compact matrix notation. Let us consider a simple 1D boundary problem described by the following PDE in the domain $\Omega = [0, 1]$:

$$u_{,xx} + f(x) = 0 \quad (\text{A.2})$$

with the following boundary conditions on the boundary Γ :

$$u(1) = g \quad \text{on} \quad \Gamma_d \quad (\text{A.3})$$

and

$$-u_{,x}(0) = h \quad \text{on} \quad \Gamma_n \quad (\text{A.4})$$

Eqs. (A.3) (A.4) are commonly known as *Dirichlet* and *Neumann boundary conditions*, respectively. Eq.(A.2) is known as the *strong form* of the problem. The FEM is based on solving the integral form of the given PDE, the so-called *weak form*. To build the weak form of eq. (A.2) we need to define a set of trial solutions \mathcal{U} that satisfy eq. (A.2), and a set of weighting functions \mathcal{W} that satisfy $w(1) = 0$:

$$\mathcal{U} = \{u | u \in H^1, u(1) = g\} \quad (\text{A.5})$$

$$\mathcal{W} = \{w | w \in H^1, w(1) = 0\} \quad (\text{A.6})$$

where H^1 means that the derivatives must be square-integrable. To obtain the weak form, the strong form is first pre-multiplied by the weighting functions and integrated over the domain:

$$\int_{\Omega} w u_{,xx} d\Omega + \int_{\Omega} w f d\Omega = 0 \quad (\text{A.7})$$

Eq.(A.7) is then reduced by performing an integration by parts:

$$\int_{\Omega} w_{,x} u_{,x} d\Omega = \int_{\Omega} w f d\Omega + w(0)u_{,x}(0) - w(1)u_{,x}(1) \quad (\text{A.8})$$

It shall be noted that Neumann boundary conditions (the new terms in the right-hand-side) appear in a natural way after integrating the weak form by parts. In this particular case we have:

$$w(0)u_{,x}(0) = w(0)h \quad (\text{A.9})$$

$$w(1)u_{,x}(1) = 0 \quad (\text{A.10})$$

Hence the weak form of eq.(A.2) yields:

$$\int_{\Omega} w_{,x}u_{,x}d\Omega = \int_{\Omega} wf d\Omega + w(0)h \quad (\text{A.11})$$

The strong and weak forms are equivalent and have identical solutions. The latter is solved by approximating the trial solutions and the weighting functions as a set of functions multiplied by unknown parameters a_n and b_n :

$$u(x) \approx \tilde{u}(x) = \sum_{n=1}^d \phi_n a_n + g\phi_{n+1} \quad (\text{A.12})$$

$$w(x) \approx \tilde{w}(x) = \sum_{m=1}^d \psi_m b_m \quad (\text{A.13})$$

where $n = 1, 2, \dots, d$, d is the number of unknowns in the problem, the functions ϕ_n and ψ_n (often referred as *shape functions*) are assumed to be zero at all locations where the boundary conditions are prescribed, and the term $g\phi_{n+1}$ must satisfy the boundary conditions. Different choices for the functions ϕ_n and ψ_m exist. However, throughout all this study the Galerkin method [Galerkin, 1915] is employed; therefore $\phi_n = \psi_n$. The advantage of this method is that it usually leads to symmetric matrices. To simplify the notation, I adopt the standard symbol of the shape functions $N_n = \phi_n = \psi_n$. The approximate Galerkin solution of the weak form is obtained substituting eqs. (A.12) and (A.13) into eq. (A.11):

$$\int_{\Omega} \sum_{n=1}^d N_{m,x} b_m \sum_{n=1}^d N_{n,x} a_n d\Omega = \int_{\Omega} \sum_{m=1}^d N_n b_m f d\Omega + \sum_{m=1}^d N_n b_m(0)h - \int_{\Omega} \sum_{m=1}^d b_m N_{n,x} g N_{n+1} \quad (\text{A.14})$$

Using the bilinearity of the integrals in A.11, the problem is rewritten as:

$$0 = \sum_{m=1}^d b_m G_m \quad (\text{A.15})$$

with

$$G_m = \sum_{n=1}^d \int_{\Omega} N_{m,x} N_{n,x} a_n d\Omega - \int_{\Omega} N_n f d\Omega - N_n b_m(0)h + \int_{\Omega} N_{n,x} g N_{n+1} d\Omega \quad (\text{A.16})$$

Eq. (A.15) must hold for all the weighting functions $w \in \mathcal{W}$. Given the arbitrariness of the coefficients b_n , it is necessary that each G_m , with $m = 1, 2, \dots, d$, is equal to zero. Thus the weak form becomes:

$$\sum_{n=1}^d \int_{\Omega} N_{m,x} N_{n,x} a_n d\Omega = \int_{\Omega} N_n f d\Omega + N_n b_m(0)h - \int_{\Omega} N_{n,x} g N_{n+1} d\Omega \quad (\text{A.17})$$

Finally, the original strong form has been reduced to a set of d ordinary differential equations (A.17) that can be expressed as:

$$\sum_{n=1}^d K_{mn} a_n = f_m \quad (\text{A.18})$$

$$(\text{A.19})$$

where

$$K_{mn} = \int_{\Omega} N_{m,x} N_{n,x} d\Omega \quad (\text{A.20})$$

$$f_m = \int_{\Omega} N_n f d\Omega + N_n b_m(0)h - \int_{\Omega} N_{n,x} g N_{n+1} d\Omega \quad (\text{A.21})$$

or using a compact matrix notation:

$$\mathbf{K}\mathbf{a} = \mathbf{f} \quad (\text{A.22})$$

The physical interpretation of the matrix \mathbf{K} and the vector \mathbf{f} depends on the nature of the problem described by the set of PDEs. For instance, in elasticity or viscous flow problems, \mathbf{K} is the so-called stiffness matrix and \mathbf{f} is the force vector; for thermal advection and diffusion, \mathbf{K} is the conductivity matrix and \mathbf{f} is the heat flow vector; and \mathbf{K} is the permeability matrix in porous flow problems.

The parameters a_n are obtained solving the linear system of equations in (A.22). Once a_n is obtained, u can be calculated at any point $x \in \Omega$:

$$u(x) = \sum_{n=1}^d N(x)_n a_n \quad (\text{A.23})$$

A convenient way to approximate N consists in constructing a *mesh* formed by set of discrete points that subdivide the domain Ω . These points are called *nodes* and each one of the nodes is connected to the neighbouring nodes, defining the so-called *elements*. These elements may have different shapes, and the appropriateness of a specific shape is ultimately determined by the nature of the problem. Each shape function is defined so that $N_n = 1$ at the node n and $N_m = 0$ at the nodes $n \neq m$. Common elements shapes are one dimensional bars for 1D problems; triangles and quadrilaterals for 2D problems; and tetrahedron and hexahedrons for 3D problems. The accuracy of the solution obtained with the FEM depends on the number and spatial distribution of the nodes of the mesh and the definition of the shape functions.

The number of nodes in an element defines the polynomial order of the shape functions. For example, a 2D triangular element with 3 nodes has linear shape functions that are constructed from piecewise linear approximations. If one needed N to be quadratic, the 2D triangular element requires of 6 nodes. The choice of the polynomial order of N is not trivial and depends on the physical problem described by the PDE. To obtain an accurate FEM solution, the order of N should be equal or higher than the order of the partial derivatives of the PDE.

A.2 Weak formulation of the Stokes equations

A.2.1 Thermal diffusion

The Lagrangian time-dependent diffusion equation in a domain Ω is defined by the following boundary problem:

$$\rho C_p \frac{DT}{Dt} = \kappa \nabla^2 T + Q \quad (\text{A.24})$$

with the boundary conditions

$$T = g \quad \text{on } \Gamma_D \quad (\text{A.25})$$

$$-\mathbf{n}\kappa T = q \quad \text{on } \Gamma_N \quad (\text{A.26})$$

where n is the unit outward normal vector to the boundary Γ and q is the heat flux. As described in Section A.1, to obtain the weak forms we first need to pre-multiply eq. (A.24) by the weighting function and integrate over the Ω :

$$\int_{\Omega} w \rho C_p \frac{DT}{Dt} d\Omega = \int_{\Omega} w \kappa \nabla^2 T d\Omega + \int_{\Omega} w Q d\Omega \quad (\text{A.27})$$

The temperature field is then approximated as a linear combination of the shape functions and unknown parameters:

$$T \approx \sum_{n=1}^d N_n \tilde{T}_n \quad (\text{A.28})$$

Employing the Galerkin method, the strong form of eq. (A.27) yields:

$$\int_{\Omega} \mathbf{N} \rho C_p \frac{D(\mathbf{N}\tilde{T})}{Dt} d\Omega = \int_{\Omega} \mathbf{N}^T \kappa \nabla^2 (\mathbf{N}\tilde{T}) d\Omega + \int_{\Omega} \mathbf{N} Q d\Omega \quad (\text{A.29})$$

Eq. (A.29) is integrated by parts to reduce the order of the second derivatives:

$$\int_{\Omega} \mathbf{N} \rho C_p \frac{D(\mathbf{N}\tilde{T})}{Dt} d\Omega = - \int_{\Omega} (\nabla \mathbf{N}^T) \kappa (\nabla \mathbf{N}\tilde{T}) d\Omega + \int_{\Omega} \nabla (\mathbf{N}^T \kappa \nabla \mathbf{N}\tilde{T}) d\Omega + \int_{\Omega} \mathbf{N} Q d\Omega \quad (\text{A.30})$$

and using the divergence theorem we finally obtain the strong form of the equation of heat diffusion:

$$\int_{\Omega} \mathbf{N} \rho C_p \frac{D(\mathbf{N}\tilde{T})}{Dt} d\Omega + \int_{\Omega} (\nabla \mathbf{N}^T) \kappa (\nabla \mathbf{N}\tilde{T}) d\Omega = \int_{\Omega} \mathbf{N} Q d\Omega - \int_{\Gamma_N} \mathbf{N} q d\Gamma_N \quad (\text{A.31})$$

A.2.2 Conservation of momentum and conservation of mass

The resulting set of governing equations of the problem is solved numerically using the FEM to generate a system of linear equations. The governing equations (A.48) and (A.49) are transformed into their weak forms with help of the trial solutions and weighting functions. As explained previously in this chapter, the first step consists in integrating the Stokes equations over the domain Ω and pre-multiplying them by two sets of weighting functions: w for the velocity field; and q for the pressure. For the conservation of momentum, that is:

$$\int_{\Omega} w \tau_{ij,j} d\Omega - \int_{\Omega} q p_{,j} d\Omega = - \int_{\Omega} w \rho g_i d\Omega \quad (\text{A.32})$$

Integrating (A.32) by parts we obtain:

$$\int_{\Omega} (w \tau_{ij})_{,j} d\Omega - \int_{\Omega} w_{,j} \tau_{ij,j} d\Omega + \int_{\Omega} q_{,j} p d\Omega = - \int_{\Omega} w \rho g_i d\Omega \quad (\text{A.33})$$

The primary variables u and p are then approximated using the shape functions N_u for the velocity field and N_p for pressure:

$$u(x, y) \approx \sum_{n=1}^d N_u^n(x, y) \tilde{u}_n \quad (\text{A.34})$$

$$p(x, y) \approx \sum_{n=1}^d N_p^n(x, y) \tilde{p}_n \quad (\text{A.35})$$

where the subindex n is the nodal index and nn is the number of nodes in the element and \tilde{u}_n and \tilde{p}_n are the trial solutions of velocity and pressure, respectively. Employing the Galerkin procedure, the governing

equations (2.16) and (2.17) are transformed into their weak forms using the shape functions as trial functions. Substituting the Galerkin approximations of the primary fields (A.34) and (A.35) into the corresponding weak forms, and following the same procedure as in the thermal diffusion problem, we obtain:

$$\int_{\Omega} N_{u,j} \left(2\eta \left[\dot{\epsilon}_{ij} - \frac{1}{3} \dot{\epsilon}_{kk} \delta_{ij} \right] \right) d\Omega - \int_{\Omega} N_{p,j} p d\Omega = \quad (\text{A.36})$$

$$\int_{\Omega} N_u \rho g_i d\Omega - \int_{\Omega} N_{u,j} \chi \widehat{\tau}_{ij} d\Omega + \int_{\Gamma_N} N_u n_i t d\Gamma$$

$$\int_{\Omega} N_p u_{i,i} d\Omega = 0 \quad (\text{A.37})$$

or

$$\int_{\Omega} (\nabla \mathbf{N}_u)^T \left(2\eta \left[\dot{\epsilon} - \frac{1}{3} \text{tr}(\dot{\epsilon}) \mathbf{I} \right] \right) d\Omega - \int_{\Omega} \nabla \mathbf{N}_p \mathbf{p} d\Omega = \quad (\text{A.38})$$

$$\int_{\omega} \mathbf{N}_u \rho \mathbf{g} d\Omega - \int_{\Omega} \nabla \mathbf{N}_u \chi \widehat{\boldsymbol{\tau}} d\Omega + \int_{\Gamma_N} \mathbf{N}_u \mathbf{n} t d\Gamma$$

$$\int_{\Omega} \mathbf{N}_p \nabla \mathbf{u} d\Omega \quad (\text{A.39})$$

A.3 Mathematical description of a visco-elastic flow

The deformation of lower mantle material is dominated by viscous flow, whereas lithospheric and upper mantle materials have a non-negligible contribution of elastic stresses. Elastic deformation is incorporated by employing a Maxwell material model, where the deviatoric strain rate is the summation of the viscous and elastic strain rates:

$$\dot{\epsilon}'_{ij} = \dot{\epsilon}'_{ij}{}^{(viscous)} + \dot{\epsilon}'_{ij}{}^{(elastic)} = \frac{\tau_{ij}}{2\eta} + \frac{1}{2G} \frac{D\tau_{ij}}{Dt} \quad (\text{A.40})$$

where G is the shear modulus and $D\tau_{ij}/Dt$ is the objective deviatoric stress rate. The Zaremba-Jaumann derivative (e.g. Hashiguchi and Yamakawa [2012]) is used to compute the objective deviatoric stress rate in eq. (A.40):

$$\frac{D\tau_{ij}}{Dt} = \frac{\partial \tau_{ij}}{\partial t} + u_k \tau_{ij,k} - \omega_{ik} \tau_{kj} + \tau_{ik} \omega_{kj} \quad (\text{A.41})$$

where the spin tensor ω_{ij} is the skew-symmetric part of the velocity gradient tensor:

$$\omega_{ij} = \frac{1}{2} (u_{i,j} - u_{j,i}) \quad (\text{A.42})$$

Because a Lagrangian formulation is used, the advection term $u_k \tau_{ij,k}$ in the right-hand-side of eq. (A.41) vanishes. Following the implementation of large-strain elastic deformation described by Moresi et al. [2003] and Kaus [2010], eq. (A.41) is approximated in an implicit manner:

$$\frac{D\tau_{ij}}{Dt} \approx \frac{\tau_{ij}^{n+1} - \tau_{ij}^n}{\Delta t} + \tau_{ij}^{rot} \quad (\text{A.43})$$

where $\tau_{ij}^{rot} = -\omega_{ik}^n \tau_{kj}^n + \tau_{ik}^n \omega_{kj}^n$ are the terms of eq. (A.43) associated with rotation of the stress field, and the superscript n denotes the time step. Substitution of eq. (A.43) into eq. (A.40) followed by the rearrangement of its terms leads to the visco-elastic constitutive law:

$$\tau_{ij} = 2\eta_{eff} \left(\dot{\epsilon}_{ij} - \frac{1}{3} \dot{\epsilon}_{kk} \delta_{ij} \right) + \chi \widehat{\tau}_{ij} \quad (\text{A.44})$$

where

$$\eta_{eff} = \frac{1}{\frac{1}{\eta} + \frac{1}{G\Delta t}} \quad (\text{A.45})$$

$$\chi = \frac{1}{1 + \frac{G\Delta t}{\eta}} \quad (\text{A.46})$$

$$\widehat{\tau}_{ij} = \tau_{ij}^n + (\omega_{ik}^n \tau_{kj}^n - \tau_{ik}^n \omega_{kj}^n) \Delta t \quad (\text{A.47})$$

Here η_{eff} is the effective viscosity and Δt is the time step. A pure viscous rheology is recovered if $\Delta t \rightarrow \infty$. Note that the visco-elastic deformation obtained per time step depends on the time step size. However, the deformation after a certain simulation time has to be independent of the chosen time step. The Stokes flow problem for a visco-elastic fluid in Ω is determined by:

$$\tau_{i,j,j} - p_{,j} = -\rho g_i \quad (\text{A.48})$$

$$u_{i,i} = 0 \quad (\text{A.49})$$

and the boundary conditions

$$u_i = h_i \quad \text{on } \Gamma_D \quad (\text{A.50})$$

$$\sigma_{ij} n_j = t_i \quad \text{on } \Gamma_N \quad (\text{A.51})$$

Now we can substitute eq. (A.44) in eq. (A.48) and obtain the following expressions for the x axis:

$$2\eta_{eff} \left(\dot{\epsilon}_{xx} - \frac{1}{3} \dot{\epsilon}_{kk} \delta_{ij} \right)_{,x} + (2\eta_{eff} \dot{\epsilon}_{xz})_{,z} - p_{,x} = -\chi (\widehat{\tau}_{xx,x} + \widehat{\tau}_{xz,z}) \quad (\text{A.52})$$

and for the z axis:

$$2\eta_{eff} \left(\dot{\epsilon}_{zz} - \frac{1}{3} \dot{\epsilon}_{kk} \delta_{ij} \right)_{,z} + (2\eta_{eff} \dot{\epsilon}_{zx})_{,x} - p_{,z} = -\chi \left(\widehat{\tau}_{zx,x} + \widehat{\tau}_{zz,z} - \frac{1}{3} \dot{\epsilon}_{kk,z} \widehat{\tau}_{zz,z} \right) - \rho g_z \quad (\text{A.53})$$

and using the definition of the strain tensor we obtain the strong forms of the conservation of momentum:

$$\eta_{eff} \left(\frac{4}{3} u_{x,x} - \frac{2}{3} u_{z,z} \right)_{,x} + \eta_{eff} (u_{x,z} + u_{z,x})_{,z} - p_{,x} = -\chi (\widehat{\tau}_{xx,x} + \widehat{\tau}_{xz,z}) \quad (\text{A.54})$$

for the x axis, and:

$$\eta_{eff} \left(\frac{4}{3} u_{z,z} - \frac{2}{3} u_{x,x} \right)_{,z} + \eta_{eff} (u_{x,z} + u_{z,x})_{,z} - p_{,z} = -\chi (\widehat{\tau}_{zx,x} + \widehat{\tau}_{zz,z}) - \rho g_z \quad (\text{A.55})$$

for the z axis. Eqs.(A.54) and (A.55) are then transformed into weak form and solved using the FEM. Note that the visco-elastic Stokes problem is very similar to a purely viscous Stokes flow. The main difference lay in the subtraction of the rotated stress from the previous time step in the right-hand-side of the momentum equation. The "purely viscous" viscosity is also substituted by new effective viscosity that averages the resistance of the material to viscous and elastic strain. and using the divergence theorem:

$$\int_{\Omega} w_{,j} \tau_{ij} d\Omega - \int_{\Omega} q_{,j} p d\Omega = \int_{\Omega} w \rho g_i d\Omega + \int_{\Gamma_N} w \tau_{ij} n_j d\Gamma \quad (\text{A.56})$$

The weak form of the incompressible constraint is:

$$\int_{\Omega} q u_{i,i} d\Omega = 0 \quad (\text{A.57})$$

The variational form of the visco-elastic Stokes flow is then described as follows:

$$\int_{\Omega} w_{,j} \tau_{ij} d\Omega - \int_{\Omega} q_{,j} p d\Omega - \int_{\Omega} q u_{i,i} d\Omega = \int_{\Omega} w \rho g_i d\Omega + \int_{\Gamma_N} w t_i d\Gamma \quad (\text{A.58})$$

Using the constitutive relationship (A.44), eq. (A.58) yields:

$$\begin{aligned} \overbrace{\int_{\Omega} w_{,j} \left(2\eta \left[\dot{\varepsilon}_{ij} - \frac{1}{3} \dot{\varepsilon}_{kk} \delta_{ij} \right] \right)}^{\text{stress-strain relationship}} d\Omega - \overbrace{\int_{\Omega} q_{,j} p d\Omega}^{\text{pressure}} - \overbrace{\int_{\Omega} q u_{i,i} d\Omega}^{\text{incompressibility constrain}} = \\ \underbrace{\int_{\Omega} w \rho g_i d\Omega}_{\text{buoyancy forces}} - \underbrace{\int_{\Omega} w_{,j} \chi \widehat{\tau}_{ij} d\Omega}_{\text{rotated stresses}} + \underbrace{\int_{\Gamma_N} w t_i d\Gamma}_{\text{Neumann BC}} \end{aligned} \quad (\text{A.59})$$

