

1 Ground motion optimised orbit feedback design for the
2 future linear collider

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7 **Abstract**

The future linear collider has strong stability requirements on the position of the beam along the accelerator and at the interaction point (IP). The beam position will be sensitive to dynamic imperfections in particular ground motion. A number of mitigation techniques have been proposed to be deployed in parallel: active and passive quadrupole stabilisation and positioning as well as orbit and IP feedback. This paper presents a novel design of the orbit controller in the main linac and beam delivery system. One global feedback controller is proposed based on an SVD-controller (Singular Value Decomposition) that decouples the large multi-input multi-output system into many independent single-input single-output systems. A semi-automatic procedure is proposed for the controller design of the independent systems by exploiting numerical models of ground motion and measurement noise to minimise a target parameter, e.g. luminosity loss. The novel design for the orbit controller is studied for the case of the Compact Linear Collider (CLIC) in integrated simulations, which include all proposed mitigation methods. The impact of the ground motion on the luminosity performance is examined in detail. It is shown that with the proposed orbit controller the tight luminosity budget for ground motion effects is fulfilled and accordingly, an essential feasibility issue of CLIC has been addressed. The orbit controller design is robust and allows for a relaxed BPM resolution, while still maintaining a strong ground motion suppression performance compared to traditional methods. We believe that the described method could easily be

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applied to other accelerators and light sources.

8 *Keywords:* orbit feedback system, ground motion, SVD decoupling

9 **1. Introduction**

10 The future linear collider [1, 2] requires beam sizes at the interaction point
11 in the nanometre range to achieve its nominal luminosity. These requirements
12 make the luminosity performance sensitive to ground motion. Ground motion
13 misaligns the accelerator components over time, which excites beam oscillations.
14 These beam oscillations degrade the average luminosity via generated beam-
15 beam offset and emittance increase due to filamentation [3]. The ground motion
16 problem is considered to be a severe feasibility issue for the design of the future
17 linear collider.

18 Several mitigation methods have been put in place to cope with the ground
19 motion issue. While static component misalignments are cured with the help
20 of beam-based alignment [4] and tuning methods [5], the effect of dynamic
21 misalignments is reduced with mechanical stabilisation systems [6] and beam-
22 based feedback controllers. In this paper, we present a novel design strategy
23 for the orbit controller for dynamic alignment in the main linac and the beam
24 delivery system of the future linear collider.

25 Significant contributions to the topic of orbit feedback systems for linear
26 accelerators have already been achieved, e.g. for the Stanford Linear Collider
27 (SLC) at SLAC [7] and the Next Linear Collider (NLC) summarised in [8]. These
28 feedback systems are based on local correction sections that are exchanging
29 information among each other in order to avoid overcorrection. This local nature
30 was introduced to cope with model errors and measurement noise. However, in
31 [8] it is mentioned that better performance was achieved by moving from many
32 local feedback systems to fewer more global ones. As a result of this observation,
33 we present in this work a global feedback strategy similar to the ones use in
34 circular colliders [9, 10]. The design is especially optimised to cope with the
35 higher sensitivity of the global feedback to model errors and measurement noise.

36 After the introduction of the necessary models of the accelerator, ground
37 motion and ground motion effects on the beam oscillations in section 2, the
38 proposed orbit control algorithm is presented in section 3. It is based on the
39 decoupling of the inputs (correctors) and outputs (measurements of the beam
40 position monitors (BPMs)) of the accelerator system with the help of the *sin-*
41 *gular value decomposition* (SVD), which is a well known approach for orbit
42 controllers [9, 10]. The innovation of the design in this paper is a method to
43 design each individual controller for the decoupled accelerator systems. A semi-
44 automatic procedure is proposed to choose open controller parameters, such
45 that a given target function (average beam orbit, emittance or luminosity) is
46 minimised with respect to ground motion excitation and measurement noise.
47 The designer still has the freedom to incorporate important system knowledge
48 in the design procedure. Since each controller does not have to be hand-tuned,
49 the design time is significantly reduced.

50 Also other orbit controller designs aim to minimise a given target function
51 with respect to the mentioned excitation signals, but do this only in a qualita-
52 tive way. The proposed method, on the other hand, uses explicit models of the
53 ground motion, the measurement noise and the influence of these signals on the
54 target function. In this way, the presented design method can incorporate the
55 rich ground motion knowledge [11, 12] in a quantitative manner. This model-
56 based approach improves the efficiency of the ground motion suppression. Ad-
57 ditionally, the found controllers are robust with respect to measurement noise,
58 since this imperfection has been explicitly included in the design procedure.

59 In section 4 the proposed design procedure is applied to design an orbit con-
60 troller for the *Compact Linear Collider* (CLIC). As a target function, the lumi-
61 nosity will be minimised. The flexibility of the design method will be demon-
62 strated by using the available design freedom to cope with a problem arising
63 from the specific structure of CLIC.

64 To evaluate the effectiveness of the ground motion optimised orbit controller,
65 an integrated simulation framework has been set up (section 5). This simulation
66 framework combines and extends existing codes to perform full-scale simulations

67 that include beam tracking, realistic ground motion misalignments produced by
 68 a ground motion generator and the beam-beam interactions. Also all ground
 69 motion mitigation methods of CLIC have been implemented. Beside simulations
 70 of the luminosity performance over different time scales, also robustness studies
 71 (section 6) will be presented. A discovered sensitivity to beam energy variations
 72 will be addressed by filtering the dispersive orbit from the measurement data.

73 2. Modeling

74 The orbit controller design in this work is based on models of the beam orbit
 75 excitations in the accelerator and site-dependent ground motion models, which
 76 are briefly introduced. It will turn out in section 4 that besides the accelerator
 77 and ground motion models also a model of the ground motion induced beam
 78 offset in the BPMs is required for the controller design. Such beam oscillation
 79 models can be calculated from the accelerator and ground motion models as
 80 will also be shown in this section.

81 2.1. Accelerator model

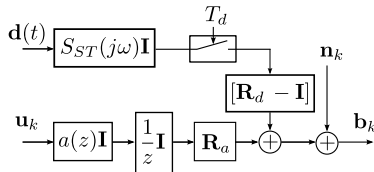


Figure 1: Block diagram of the model describing the beam oscillations along the beamline. The details are explained in the text.

82 In this section the accelerator system is described in a mathematical frame-
 83 work. As the accelerator is a discrete-time system, the so-called \mathcal{Z} -transform
 84 is used for its representation [13]. The \mathcal{Z} -transform transforms a discrete-time
 85 signal or system into its frequency representation and is therefore analogous to
 86 the Laplace-transform, used for continuous systems.

87 The accelerator model is depicted in Figure 1. The BPM readings \mathbf{b}_k (where
 88 k is the discrete-time index) are influenced by

- 89 • the BPM noise \mathbf{n}_k ;
- 90 • the positions of the quadrupoles via the quadrupole response matrix \mathbf{R}_d
- 91 and the positions of the BPMs themselves ($-\mathbf{I}$). The quadrupoles and
- 92 BPMs are displaced by the continuous ground motion $\mathbf{d}(t)$ folded with
- 93 the stabilisation transfer function $S_{ST}(j\omega)$. The element positions are
- 94 sampled with a repetition time T_d ;
- 95 • the actuator (corrector) settings \mathbf{u}_k via the actuator dynamics $a(z)$ and
- 96 actuator response matrix \mathbf{R}_a .

97 As actuators, corrector dipole magnets or mechanically movable quadrupoles

98 can be used for example. In the latter case \mathbf{R}_a is equal to \mathbf{R}_d . It is assumed

99 that all actuators have the same dynamics $a(z)$ and all elements have the same

100 stabilisation transfer function $S_{ST}(j\omega)$.

101 Characteristics of the accelerator system are its large size (multi-input multi-

102 output (MIMO)) and its relatively simple structure without internal back cou-

103 pling.

104 2.2. Ground motion models

105 Two types of ground motion models are used in this paper: the *ATL law*

106 and models based on the *two-dimensional power spectral density* (PSD). Both

107 models include correlations in time and space. For long time periods (longer

108 than several minutes) the relatively simple ATL law is applied (see [12] for

109 further information). However, some short time scale effects like technical noise

110 and the micro-seismic peak (a strong ground motion in the 0.15 Hz region caused

111 by ocean waves coupling to the shore) are not taken into account [11]. Therefore,

112 for short time scales, models based on the two-dimensional PSD $P(\omega, k)$ are used

113 to model the more complex high-frequency ground motion behaviour properly.

114 The two-dimensional PSD corresponds to a superposition of independent ground

115 motion waves with an angular frequency ω and a wave number k . Contrary to

116 the ATL law, the short time scale models are only valid for a limited time,

117 typically up to one minute.

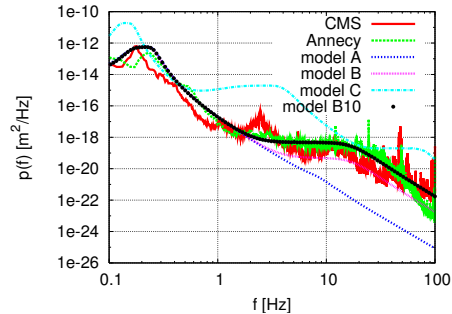


Figure 2: Ground motion power spectral density for several sites and models.

118 A generic form for ground motion models based on the two-dimensional
 119 PSD has been developed in [11]. The open parameters of this generic model
 120 can be adapted to the site-dependent ground motion properties. Three different
 121 scenarios taken from literature have been considered in this work (see Figure 2).
 122 Model A is based on measurements in the empty LEP tunnel, which is a very
 123 quiet site. Model B corresponds to measurements on the Fermilab site. Model
 124 B10 is model B with an amplified peak to match technical noise measured at
 125 LAPP (Anney) [14] and in the CMS hall [15]. The according model parameters
 126 are summarised for example in [16].

127 2.3. Beam oscillations due to ground motion

Beam oscillations due to ground motion are mainly caused by the misalign-
 ment of the quadrupole magnets. A model for the spectra of the BPM read-
 ings due to these beam oscillations can be derived with the help of the two-
 dimensional ground motion PSD. Taking into account the action of the quad-
 rupole stabilisation system $S_{ST}(j\omega)$, the spectra $B_i(\omega)$ of the BPM readings of
 the i^{th} BPM can be written as

$$B_i(\omega) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} P(\omega, k) |S_{ST}(j\omega)|^2 G_i(k)^2 dk, \quad (1)$$

The term $G_i(k)$ describes the average beam offset in the i^{th} BPM due to a
 ground motion wave with wave number k and an amplitude of 1. An expression

for $G_i(k)$ can be found by modifying the derivation of the spectrum of the beam-beam offset in [11] slightly, which leads to (see [16] for more details)

$$G_i(k)^2 = \sum_{m=0}^{N_i+1} \sum_{n=0}^{N_i+1} r_{i,m} r_{i,n} \cos(kL_{m,n}), \quad (2)$$

where N_i is the index of the last quadrupole influencing the i^{th} BPM and $L_{m,n}$ is the distance between the quadrupoles m and n . The parameter $r_{i,j}$ describes the change of the beam offset b_i in the i^{th} BPM due to a change of the misalignment x_j of the j^{th} quadrupole. To shorten the notation, the terms $r_{i,0}$ and r_{i,N_i+1} are used to describe the beam offset in the BPMs due to an initial beam offset and a misalignment of the BPM respectively. This leads to the definition

$$r_{i,j} = \begin{cases} 1 - \sum_{m=1}^{N_i} r_{m,j}, & \text{for } j = 0, \\ db_i/dx_j, & \text{for } j = 1, 2, \dots, N_i, \\ -1, & \text{for } j = N_i + 1. \end{cases}$$

Even though the expression for $G_i(k)$ in Eq. (2) seems to be difficult to evaluate, the spectra for all BPMs can be very efficiently calculated. By introducing the vector $\mathbf{G}(k)$, whose i^{th} element is $G_i(k)$, and expanding the cosine term with trigonometric identities, Eq. (2) can be rewritten as (see [16] for details)

$$\mathbf{G}(k)^2 = \left[\tilde{\mathbf{R}}_d \mathbf{c}(k) \right]^{\circ 2} + \left[\tilde{\mathbf{R}}_d \mathbf{s}(k) \right]^{\circ 2} \quad \text{with} \quad (3)$$

$$\tilde{\mathbf{R}}_d = \begin{bmatrix} \tilde{r} & \mathbf{R}_d & -\mathbf{I} \end{bmatrix}, \quad \tilde{r}_i = r_{i,0},$$

where $\circ 2$ symbolises the element-wise square of a vector also called Hadamard's square, \mathbf{R}_d is the orbit response matrix due to quadrupole magnet misalignments, \mathbf{I} is the identity matrix and

$$c_i(k) = \begin{cases} \cos(z_{i-1}k), & \text{for } i = 1, \dots, N_q + 1, \\ \cos(\tilde{z}_{i-N_q-1}k), & \text{for } i = N_q + 2, \dots, N_q + N_b + 1, \end{cases}$$

$$s_i(k) = \begin{cases} \sin(z_{i-1}k), & \text{for } i = 1, \dots, N_q + 1, \\ \sin(\tilde{z}_{i-N_q-1}k), & \text{for } i = N_q + 2, \dots, N_q + N_b + 1, \end{cases}$$

128 where $c_i(k)$ and $s_i(k)$ are the i^{th} element of $\mathbf{c}(k)$ and $\mathbf{s}(k)$, z_0 is the longitudinal
129 position of the entrance of the beamline, z_i and \tilde{z}_i are the longitudinal posi-
130 tions of the i^{th} quadrupole and BPM respectively and N_q and N_b are the total
131 numbers of quadrupoles and BPMs in the beam line. Note that too much noise
132 in the response matrix \mathbf{R}_d can lead to numerical problems at the evaluation of
133 expression Eq. (3).

134 3. Beam-based Orbit Controller

135 An orbit controller uses the BPM measurements \mathbf{b}_k to calculate corrector
136 settings for the next time step \mathbf{u}_{k+1} . These actuator settings are supposed to
137 steer the beam back onto its nominal trajectory.

138 A semi-automatic design procedure for an high performing SVD-based orbit
139 controller will be presented. The design procedure consists of the following three
140 steps.

- 141 • Decoupling of inputs and outputs
- 142 • General time dependent filter design for all decoupled channels
- 143 • Gain optimisation for each decoupled channel

144 3.1. Decoupling

For the orbit feedback system an SVD controller is chosen, which is a special
form of a decoupling controller [17]. A decoupling procedure converts a MIMO
system into a new system, in which every input acts only on one output. For
each of the decoupled system channels an independent single-input single-output
(SISO) controller can be designed. This splitting of one large control problem
into many smaller ones simplifies the design procedure significantly. For an
SVD controller, the decoupling is achieved by using the SVD of the response
matrix $\mathbf{R}_d = \mathbf{U}\mathbf{\Sigma}\mathbf{V}^T$, where \mathbf{U} and \mathbf{V} are orthonormal matrices and $\mathbf{\Sigma}$ is a
diagonal matrix with the singular values σ_i as elements. An important property

of an orthonormal matrix \mathbf{A} is that $\mathbf{A}^T \mathbf{A} = \mathbf{I}$. If the system in Figure 1 is pre-multiplied with \mathbf{V} and post-multiplied with \mathbf{U}^T , the new decoupled system

$$\begin{aligned}\hat{\mathbf{b}}_k &= \mathbf{U}^T \mathbf{U} \mathbf{\Sigma} \mathbf{V}^T \mathbf{V} \hat{\mathbf{u}}_{k-1} = \mathbf{\Sigma} \hat{\mathbf{u}}_{k-1}, \quad \text{with} \\ \hat{\mathbf{b}}_k &= \mathbf{U}^T \mathbf{b}_k \quad \text{and} \quad \hat{\mathbf{u}}_k = \mathbf{V}^T \mathbf{u}_k\end{aligned}\quad (4)$$

145 is formed. The inputs and outputs of the new system ($\hat{\mathbf{u}}_k$ and $\hat{\mathbf{b}}_k$) do not
 146 correspond to individual correctors and BPMs anymore, but to input and output
 147 vector directions, given by the columns of \mathbf{U} and \mathbf{V} . Consequently, also the
 148 ground motion and the BPM noise have to be transformed to $\hat{\mathbf{d}}_k = \mathbf{V}^T \mathbf{d}_k$ and
 149 $\hat{\mathbf{n}}_k = \mathbf{U}^T \mathbf{n}_k$. Since the system has no internal back coupling, the decoupling is
 valid for all frequencies, which is usually not achievable.

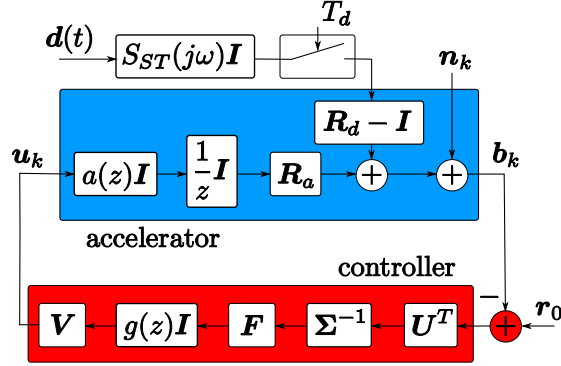


Figure 3: Block diagram of the orbit feedback system, where the values f_i and $1/\sigma_i$ have been collected in the diagonal matrices \mathbf{F} and $\mathbf{\Sigma}^{-1}$

150

151 3.2. Time dependent Filter

152 For each of the decoupled channels an individual controller of the form
 153 $g(z)f_i/\sigma_i$ is designed, where i is the channel index. The division by the sin-
 154 gular value σ_i corresponds to a normalisation of the loop gain in Eq. (4). The
 155 time dependent filter $g(z)$ is parametrised the same way for all controllers to
 156 reduce the problem complexity. Additionally, one gain factor f_i is left open per
 157 channel to account for the different ground motion excitations and BPM noise

158 for each channel. The complete system is visualised in Figure 3, where \mathbf{r}_0 is the
 159 target beam orbit which is different from zero in the BDS.

The designer is free to choose the time dependent filter $g(z)$ according to his needs. The filter changes the frequency response of the open loop to achieve required results (loop-shaping method [17]). Here, we propose a time dependent filter that is composed of 4 parts:

$$g(z) = I(z)L(z)P(z)E(z).$$

Integrator $I(z)$: The central element of $g(z)$ is the integrator

$$I(z) = \frac{z}{z-1}.$$

160 This element ensures good suppression of low frequency ground motion, but
 161 amplifies BPM noise strongly.

Low pass filter $L(z)$: To improve the noise behaviour of the controller a low pass filter

$$L(z) = \frac{z(1 - e^{-\frac{T_d}{T_1}})}{z - e^{-\frac{T_d}{T_1}}}$$

162 is added, with T_d the sampling rate and T_1 a time constant that determines
 163 the cutoff frequency, i.e. how fast the controller reacts to the measured BPM
 164 values. It is a basic first-order low pass filter multiplied by z to avoid the large
 165 phase change for high frequencies [16].

166 **Peak element $P(z)$:** This element is added to allow the possibility to
 167 strengthen and weaken the controller performance in certain frequency ranges
 168 by adding higher gain in these ranges. This enables the user to incorporate
 169 system-specific knowledge to the design. An example will be given in section 4.

To increase the gain only in a limited frequency range, $P(z)$ is chosen as

$$P(z) = \frac{(1 - n_1)(1 - n_2)}{(1 - z_1)(1 - z_2)} \cdot \frac{(z - z_1)(z - z_2)}{(z - n_1)(z - n_2)},$$

170 where the poles $n_{1,2}$ and the zeros $z_{1,2}$ are conjugate complex pairs of the form
 171 $\exp((-a \pm jb)T_d)$, with $b \in \mathbb{R}$ and $a \in \mathbb{R}^+ \setminus \{0\}$ to ensure that $P(z)$ is stable and
 172 minimum-phase. The poles $n_{1,2}$ create a low-pass of second order, where the

173 location of the poles is chosen to create a significant overshoot of the frequency
 174 response in a limited frequency range before the cutoff frequency. The zeros
 175 $z_{1,2}$ do not alter the frequency response of $P(z)$ for low frequencies, but create
 176 amplification at high frequencies, which cancels the effect of the low pass of the
 177 denominator. By positioning the poles and zeros such that the high frequency
 178 effect of the denominator and the nominator cancels completely, only the over-
 179 shoot of the two second order elements remains, which creates the desired peak
 180 in the frequency response in a limited frequency range.

Phase lifting element $E(z)$: The combination of $I(z)$, $L(z)$ and $P(z)$ can lead to insufficient stability properties. An important measure for the stability of a control circuit is the phase margin, defined as the difference of the phase of the open loop frequency response at the cross-over frequency (magnitude of one) and -180° (Nyquist criterion). The element

$$E(z) = \frac{1 - n_3}{1 - z_3} \cdot \frac{z - z_3}{z - n_3}$$

181 is added to increase the phase margin. Therefore, the zero z_3 is positioned to lift
 182 the phase shortly before the cross-over frequency. The pole n_3 is positioned to
 183 cancel the action of the zero z_3 for frequencies above cross-over frequency. Both
 184 z_3 and n_3 are located on the real axis of the z -plane inside the unit circle, which
 185 can be achieved by choosing them as $\exp(-aT_d)$ with $a \in \mathbb{R}^+ \setminus \{0\}$. The phase
 186 increase and the resulting inevitable magnitude amplification can be adjusted
 187 by moving the positions of the pole and the zero .

188 3.3. Gain optimisation

For each controller loop there remains one gain parameter f_i , which can be chosen to minimise the power of each output signal with respect to the system excitation. The power of the output signal $\hat{b}_{i,k}$ can be calculated from the power spectra $\hat{B}_i(\omega)$, which are given by

$$\hat{B}_i(\omega) = \hat{S}_i(z = e^{j\omega T_d}) \hat{D}_i(\omega) - \hat{T}_i(z = e^{j\omega T_d}) \hat{N}_i(\omega),$$

where we use that the \mathcal{Z} -transform of a system evaluated at $z = \exp(j\omega T_d)$ corresponds to the transfer function of the system. The ground motion suppression

and noise transfer functions $\hat{S}_i(z)$ and $-\hat{T}_i(z)$ of the system are given by

$$\hat{S}_i(z) = \frac{1}{1 + \hat{G}_i(z)\hat{C}_i(z)}, \quad \hat{T}_i(z) = \frac{\hat{G}_i(z)\hat{C}_i(z)}{1 + \hat{G}_i(z)\hat{C}_i(z)},$$

189 where $\hat{G}_i(z)$ and $\hat{C}_i(z)$ are the \mathcal{Z} -transforms of the decoupled system channel
190 and its associated controller given by $\hat{G}_i(z) = a(z)\sigma_i/z$ and $\hat{C}_i(z) = g(z)f_i/\sigma_i$.

To calculate the channel ground motion spectra $\hat{D}_i(\omega)$ it is not possible to simply project the basic spectrum $D(\omega, k)$ on the channel input vectors \mathbf{v}_i . The problem with this approach is that the projection results in correlation between the individual channel spectra $\hat{D}_i(\omega)$. This correlation is not taken in consideration by the used model that assumes that all $\hat{D}_i(\omega)$ are independent of each other. This leads to large errors when the channel BPM spectra $\hat{B}_i(\omega)$ are calculated from the $\hat{D}_i(\omega)$. For that reason, virtually independent channel ground motion spectra $\hat{D}_i(\omega)$ are calculated by

$$\hat{D}_i(\omega) = s_i^{-1}\hat{B}_i(\omega), \quad (5)$$

which creates in the model the correct channel BPM spectra $\hat{B}_i(\omega)$, where s_i is the channel singular value of \mathbf{R}_a . The calculation of the channel BPM spectra $\hat{B}_i(\omega)$ is very similar to the calculation of the BPM spectra $B_i(\omega)$ as performed in section 2.3, only that $B_i(\omega)$ has to be additionally projected on the output directions \mathbf{u}_i . Hence $\hat{B}_i(\omega)$ can be calculated with Eqs. (1) and (3), only that (3) has to be slightly modified to

$$\mathbf{G}(k)^2 = [\mathbf{U}^T \tilde{\mathbf{R}}_d \mathbf{c}(k)]^{\circ 2} + [\mathbf{U}^T \tilde{\mathbf{R}}_d \mathbf{s}(k)]^{\circ 2}.$$

191 The decoupled noise spectra $\hat{N}_i(\omega)$ can be modelled as white noise (flat spec-
192 trum). Neglecting correlation, the according variances are given by the diagonal
193 elements of the expression $\mathbb{E}\{\hat{\mathbf{n}}_k \hat{\mathbf{n}}_k^T\} = \mathbf{U}^T \mathbb{E}\{\mathbf{n}_k \mathbf{n}_k^T\} \mathbf{U}$, where $\mathbb{E}\{\cdot\}$ is the ex-
194 pectation value of a random variable.

To find the optimal value for f_i , the signal $\hat{b}_{i,k}(f_i)$ is minimised with respect to the power norm. This is equivalent to minimising the L_1 -norm of the power spectrum of $\hat{b}_{i,k}$

$$\min_{f_i} \|\hat{b}_{i,k}(f_i)\|_{pow} = \min_{f_i} \|\hat{B}_i(\omega, f_i)\|_1 = \min_{f_i} \int_{\omega=-\infty}^{+\infty} \hat{B}_i(\omega, f_i) d\omega \quad \forall i \quad (6)$$

195 Eq. (6) can be solved numerically by evaluating the integral for different values
196 of f_i over a sufficient frequency range.

197 These parameters f_i can be optimised for different target functions than
198 the beam orbit by modeling the influence of the ground motion onto the target
199 function. An example of a luminosity optimised design will be shown in the
200 adaptation to the CLIC case.

201 4. Adaptation to the CLIC case

202 In this section, after a general introduction to CLIC and its ground motion
203 mitigation techniques, an orbit controller based on the method that was pro-
204 posed in the previous section is designed for the main linac (ML) and beam
205 delivery system (BDS).

206 4.1. Introduction to CLIC

207 CLIC [2] is an electron-positron collider, which is together with the Inter-
208 national Linear Collider (ILC) the most promising proposal for a future high
209 energy linear accelerator. Its main parameters are summarised in Table 1. CLIC
210 implements a novel two-beam acceleration scheme where the main beam, which
211 is used for the collisions, is accelerated by a second so called drive beam. The
212 very small bunch interval Δ_b of the main beam does not allow for a global intra-
213 train feedback in the ML and BDS, and thus the orbit controller will act from
214 train to train with a repetition rate f_R .

215 The CLIC ML accelerates the beams to the final energy and is about 20 km
216 long. It has a FODO lattice with a gradually increasing cell length to accustom
217 to the increasing energy. The CLIC BDS [18] is mainly responsible for collima-
218 tion and for the focusing of the beams to the required very small beam sizes σ_x^*
219 and σ_y^* at the interaction point (IP).

220 4.2. CLIC ground motion mitigation techniques

221 To counter the impact of the ground motion several mitigation techniques
222 are deployed in CLIC. Currently there are four mitigation techniques foreseen,

Centre of mass energy	E	3 TeV
Total/peak (1%) luminosity	$\mathcal{L}/\mathcal{L}_{1\%}$	$5.9/2.0 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
Hor./vert. beam size at IP	$\sigma_{x/y}^*$	40/1 nm
Hor./vert. norm. emittance at IP	$\epsilon_{x/y}^*$	660/20 nm rad
Nr. of particles per bunch	N	3.72×10^9
Repetition rate	f_R	50 Hz
Nr. bunches per beam train	N_b	312
Bunch interval	Δ_b	0.5 ns
RF gradient	G_{RF}	100 MV/m

Table 1: Most important design parameters of CLIC.

223 namely the quadrupole stabilisation, preisolator, orbit controller and IP feed-
224 back. They will be shortly summarised.

225 The IP feedback [19, 20] corrects the beam-beam offset at the IP on a train
226 to train basis by measuring the deflection angles of the colliding beams and
227 adjusting the beam position with a dipole kicker positioned between QD0 and
228 the IP. An additional intra-train IP feedback is foreseen to work within the
229 bunch train. The intra-train IP feedback is considered as a reserve option and
230 is not taken into account in this paper, contrary to the regular IP feedback.

231 Note that since the repetition rate of CLIC is 50 Hz, beam-based feedbacks
232 (orbit controller and IP feedback) are mostly effective for frequencies below a
233 few Hz. For higher frequencies other systems have to be deployed. To reduce the
234 motion of the quadrupoles for high frequencies (≥ 1 Hz), each quadrupole will be
235 positioned on an active stabilisation system [6]. Its theoretical transfer function
236 is shown in Figure 4. The peak at 0.2 Hz of the quadrupole stabilisation is close
237 to the micro-seismic peak which is unfavorable. Therefore, a targeted future
238 design is shown in the figure as well. The final doublet, the last quadrupoles QD0
239 and QF1, which are especially sensitive to luminosity loss due to ground motion,
240 will be put on a large mass block supported by air springs, the preisolator [21].
241 The preisolator acts as a passive ground motion isolation system. The combined

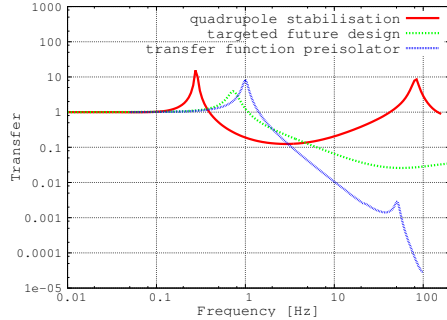


Figure 4: Amplitude of the theoretical transfer functions of the quadrupole stabilisation.

242 transfer function is shown in Figure 4.

243 4.3. Adaptation of the orbit controller

244 In this section, the orbit controller that was described in section 3 will be
 245 tuned to the CLIC case. The current baseline for the CLIC actuators is quad-
 246 rupole movers and dipole kickers as an alternative option. For the quadrupole
 247 movers a perfect actuator response $a(z) = 1$ is assumed. This is in accordance
 248 with the specifications of the actuator design [22]. The orbit controller in the
 249 ML and BDS has 2122 BPMs and 2104 correctors to its avail.

250 4.3.1. Time dependent filter

251 The parameters chosen for the time dependent filter are as follows. For the
 252 low-pass $L(z)$, T_1 was chosen as 0.1 s. In this case the $L(z)$ demagnifies signals
 253 above its cutoff frequency of about 1.4 Hz.

254 As mentioned in the previous section, the final doublet quadrupoles and the
 255 other quadrupoles are stabilised by different methods (preisolator and quad-
 256 rupole stabilisation). The according transfer functions differ strongly around
 257 0.3 Hz, see Figure 4, which causes a beam offset in the final doublet quadrupoles
 258 resulting in beam size growth at the IP due to dispersion and coupling. The el-
 259 ement $P(z)$ has been chosen to strengthen the controller in this frequency range
 260 by using $z_{1,2} = \exp((-1.43 \pm 0.4\pi j)T_d)$ and $n_{1,2} = \exp((-0.3 \pm 0.6\pi j)T_d)$.

261 This combination of $I(z)$, $L(z)$ and $P(z)$ leads to an insufficient phase
 262 margin. To improve the stability properties $E(z)$ with $z_3 = \exp(-17T_d)$ and
 $n_3 = \exp(-38T_d)$ is added to increase the phase margin to 36.3° .

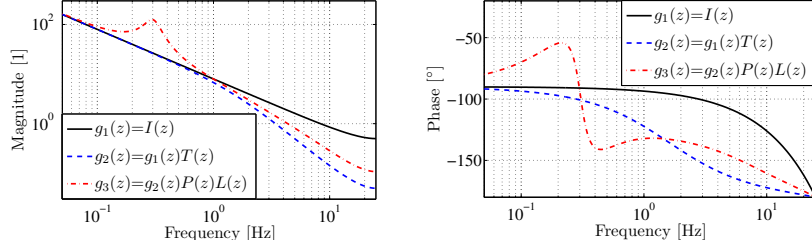


Figure 5: Magnitude (left) and phase (right) of the frequency responses of the elements of the time dependent filter $g(z)$.

263

264 The frequency responses of combinations of the elements for the time de-
 265 pendent filter $g(z)$ are shown in Figure 5. It can be seen that $L(z)$ reduces
 266 the sensitivity of the controller for high frequencies above the cutoff frequency,
 267 that $P(z)$ strengthens the controller in the region around 0.3 Hz and that $E(z)$
 268 increases the phase margin, while slightly increasing the sensitivity to high fre-
 269 quencies.

270 4.3.2. Gain optimisation

271 By solving Eq. (6) the gain factor of each controller loop is optimised to
 272 reduce the ground motion offsets in the BPMs. However, the ultimate objective
 273 of CLIC is to reduce the luminosity loss. The equation is slightly modified to
 274 account for this objective. The luminosity loss due to ground motion is caused
 275 by two effects, beam-beam offset at the IP and beam size growth. For each
 276 controller loop i the peak luminosity loss ΔL_i , which can be decomposed in
 277 the peak luminosity loss due to beam-beam offset $\Delta L_{o,i}$ and due to beam size
 278 growth $\Delta L_{c,i}$, is estimated. This is accomplished by simulations that misalign
 279 the quadrupoles with the i^{th} column of \mathbf{V} and calculate ΔL_i and $\Delta L_{c,i}$ by
 280 centering the beams. For small values $\Delta L_{o,i}$ can be calculated with $\Delta L_{o,i} =$
 281 $\Delta L_i - \Delta L_{c,i}$. The normalised peak luminosity losses for the vertical direction

are shown in Figure 6.

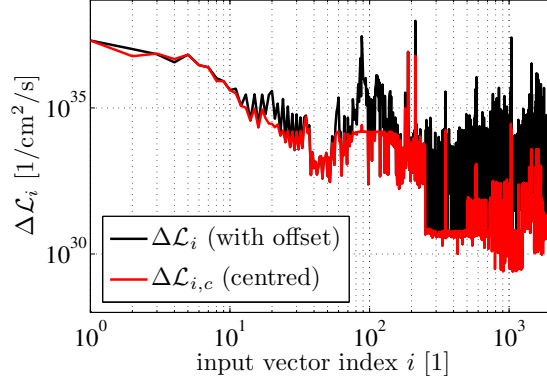


Figure 6: Normalised peak luminosity loss due to beam-beam offset at the IP and beam size growth (centred) for each decoupled controller loop for the vertical direction.

282

The modified version of Eq. (6) becomes:

$$\min_{f_i} \int_{\omega=-\infty}^{+\infty} \hat{B}_i(\omega, f_i) \left(\frac{\Delta L_{c,i}}{\Delta L_i} + \frac{\Delta L_{o,i}}{\Delta L_i} \left| \frac{1}{z} S_{IP}(e^{j\omega T_d}) \right|^2 \right) d\omega, \quad (7)$$

283 where $1/z S_{IP}(\exp(j\omega T_d))$ is the frequency response of the IP feedback, which
 284 has to be taken into account. The gain factors obtained with this method for
 285 ground motion model B10 are shown in Figure 7 for both the horizontal and
 286 vertical direction. It can be seen that all 2104 gain factors are below 1 and have
 287 been artificially given a minimum value of 10^{-6} to avoid any open control loops.

288

289 5. Simulations

290 5.1. Simulation framework

291 All simulations are performed tracking the beams through both MLs and
 292 BDSs of CLIC with PLACET [23] and beam-beam interaction code GUINEA-
 293 PIG [24], similar as in [25] for the ILC. A BPM resolution of 100 nm is assumed
 294 for the ML BPMs and 50 nm for the BDS BPMs. A ground motion generator
 295 that simulates the short term and long term ground motion models of section

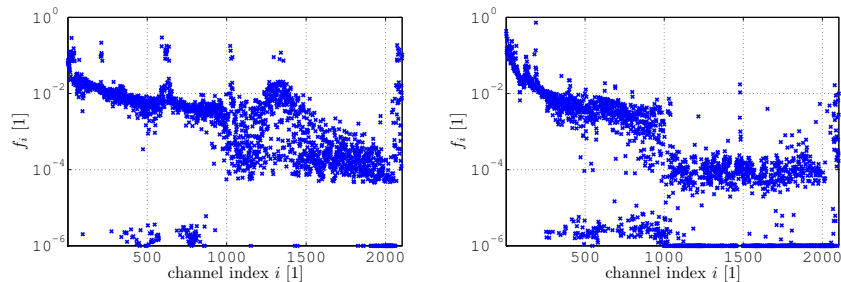


Figure 7: Gain factors f_i of the decoupled control loops for the horizontal (left) and vertical (right) direction.

296 2.2, all ground motion mitigation methods of section 4.2 and the controller of
 297 section 4.3 have been integrated in the simulations. Both the current version
 298 and the future design of the quadrupole stabilisation are studied. Besides the
 299 automatically tuned controller gains of Figure 7, a former hand-tuned version
 300 of the controller gains is tested for comparison.

301 5.2. Pre-assumptions

302 The foreseen emittance budget due to the static imperfections of the RTML,
 303 ML and BDS combined is a growth from 5 nm rad normalised geometric emit-
 304 tance at the exit of the damping rings to 20 nm rad at the IP. Instead of integrat-
 305 ing the static imperfections directly in the simulations, a simplified approach is
 306 taken by injecting a beam with an emittance of 20 nm rad at the beginning of
 307 the ML (instead of 10 nm rad) to approximate the static imperfections budget.
 308 This approach makes the assumption that the effects of the static and dynamic
 309 imperfections can be factorised and that the misaligned lattice is not adversely
 310 effecting the orbit feedback operation. Further studies will be carried out to
 311 understand the combined effect. The foreseen budget for peak luminosity loss
 312 due to dynamic imperfections in the ML and BDS is about 20%.

313 5.3. Luminosity evolution for short time scales

314 The short term ground motion generator based on the two-dimensional PSD
 315 $P(\omega, k)$ misaligns the beamline every pulse. All of the four mitigation techniques

316 are applied, i.e. preisolator, quadrupole stabilisation, IP feedback and orbit
 317 controller. Figure 8 shows that the luminosity is well preserved over a time
 318 period of 60s, which is about the maximum time for which the used ground
 319 motion generator is valid.

320 The jitter on the luminosity is caused by the remaining high frequencies of
 the ground motion and the BPM noise.

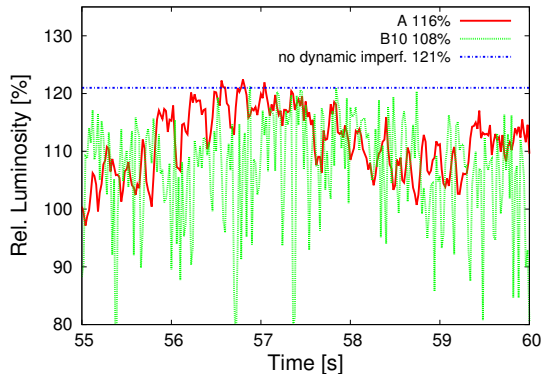


Figure 8: Example of peak luminosity for the current design over a longer time scale (60s) for several ground motion models.

321
 322 In Table 2 the relative peak luminosity performance using the auto-tuned
 323 controller for several configurations of the applied stabilisation system and the
 324 ground motion model are shown. Note that for each result a different controller
 325 has been used, since the f_i were optimised with respect to the stabilisation system
 326 and the ground motion model according to Eq. (7). It can be concluded
 327 that depending on the ground motion different stabilisation measures are re-
 328 quired. Note that for ground motion model A mitigation methods can even
 329 lower the luminosity performance. This is due to offsets between the preisolator
 330 and the rest of the beamline, which is caused by a discrepancy between the two
 331 transfer functions. Also note that an enhanced quadrupole stabilisation can
 332 improve the luminosity performance significantly.

333 The auto-tuned controller (named C_5) is compared in the following with
 334 other global feedback algorithms. The controller C_1 is considered to be the

Stabilisation	no GM	A	B	B10
Current	121 (0)	117 (4)	116 (5)	109 (12)
Future	121 (0)	121 (0)	121 (0.5)	120 (1)

Table 2: Overview of the peak luminosity performance (and luminosity loss) in % with respect to the nominal peak luminosity $\mathcal{L}_{1\%}$ for different ground motion (GM) models and quadrupole stabilisation system averaged over 20 seeds.

335 standard SVD feedback that employs an integrator instead of $g(z)$ and a basic
336 weighting of the modes by using $\hat{f}_i = 1, \forall i \leq 20$ and $\hat{f}_i = 10^{-4}, \forall i > 20$. The
337 next two controllers improve the basic controller C_1 by using either the optimised
338 gains f_i of C_5 (C_2) or the optimised time dependent filter $g(z)$ instead of the
339 integrator (C_3). Finally, C_4 consists of the optimised time dependent filter $g(z)$
340 and gains \tilde{f}_i that have been optimised by hand by applying the knowledge of
341 Figure 6.

GM model	C_1	C_2	C_3	C_4	C_5
B	24 (97)	103 (18)	24 (97)	111 (10)	116 (5)
B10	23 (98)	97 (24)	23 (98)	107 (14)	109 (12)

Table 3: Peak luminosity performance (and luminosity loss) in % with respect to the nominal peak luminosity $\mathcal{L}_{1\%}$ of different controllers (C_1 to C_5) for different ground motion models B and B10. The current stabilisation system was used and the results are averaged over 10 seeds.

342 As can be seen in Table 3, the application of non-optimised gains f_i results
343 in a very large luminosity loss (controllers C_1 and C_3). Also the use of optimised
344 gains without optimising the time dependent filter (controller C_2) causes still too
345 strong performance degradation. Only the controllers C_4 and C_5 can preserve
346 the luminosity to an acceptable level. The reason is that for the design of both
347 controllers detailed system model information has been used. It should also be
348 pointed out that C_5 still performs better than C_4 even though the gains f_i for
349 C_5 have been found automated within a few minutes of calculation, while for

350 C_4 several weeks of hand tuning has been necessary.

351 5.4. Luminosity evolution for large time scales

352 For longer time scales the ground motion generator based on the two-dimensional
353 PSDs is no longer valid. In addition, the simulations would become computa-
354 tionally intensive. Therefore, to study ground motion for longer time scales
355 the ATL law has been applied with the constant $A = 0.5 \cdot 10^{-6} \mu\text{m}^2/(\text{s m})$,
356 which is based on measurements from the LEP-tunnel. After applying the orbit
357 controller the resulting relative peak luminosity for CLIC is shown in Figure 9
358 as a function of time. It can be seen that after about half an hour the peak
359 luminosity is decreased by 10% and further optimisation is required. However,
360 it has been shown that tuning the accelerator with the BDS sextupole knobs
361 [5] can recover the luminosity fully. Since these BDS sextupole knobs are not
362 deployed in the orbit controller, this shows that the response matrix of the ac-
363 celerator does not change significantly for this time scale and therefore the orbit
364 controller performs well. Further studies will have to be done to estimate the
365 time period for which the response matrix is still accurate.

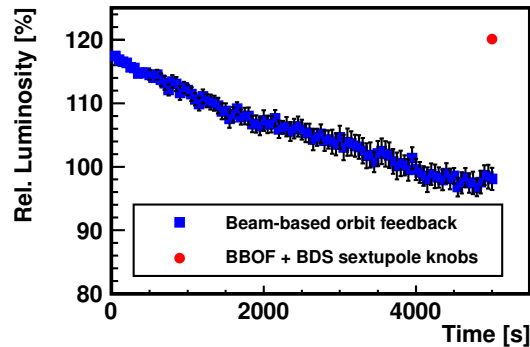


Figure 9: Peak luminosity evolution for long time scales of ground motion. After about 30 minutes, the luminosity is decreased by 10%. To correct for this loss, further optimization procedures, e.g. tuning of the BDS sextupole knobs, are required.

366 **6. Robustness Studies**

367 The robustness of the CLIC orbit controller has been studied in detail. Many
368 types of dynamic and static imperfections and their effect on the controller per-
369 formance have been studied and the controller has been found to be sufficiently
370 robust for all tested imperfections. If not stated differently, the studies were
371 carried out with the auto-tuned controller optimised for ground motion B10
372 and the current stabilisation function. Two types of imperfections (BPM reso-
373 lution and beam energy jitter) are especially important for the CLIC controller
374 and are therefore discussed separately. The outcomes of the studies of the other
375 tested imperfections are summarised in the last section and we would also like
376 to refer also to the more detailed publications [26] and [16].

377 *6.1. BPM resolution*

378 BPM noise degrades the effectiveness of the orbit controller, since a BPM
379 measurement error will propagate into the orbit correction. To evaluate the
380 required BPM resolution in the BDS, simulations have been performed where all
381 other dynamic effects as e.g. ground motion are not applied. In Figure 10 (left)
382 the relative peak luminosity loss is shown as a function of the BPM resolution.
383 It can be seen that a BPM resolution of 50 nm is required in the BDS to limit
384 the luminosity loss to 2%, while the BPM resolution in the ML can be more
385 relaxed. Previously, a BPM resolution of 10 nm was considered to be necessary.
386 However, due to consideration of static imperfections in the final focus system,
387 tighter BPM resolutions might still be required. Also shown in Figure 10 (right)
388 is the peak luminosity loss for different versions of the orbit controller, once with
389 the full controller and two times with a simple integrator ($g(z) = I(z)$), with
390 full gains ($\mathbf{F} = \mathbf{I}$) and optimised gains ($\mathbf{F} = \text{diag}(f_i)$) respectively. It can be
391 seen that the gain optimisation as well as the design of the low pass of the time
392 dependent filter reduces the back-coupling of the BPM noise.

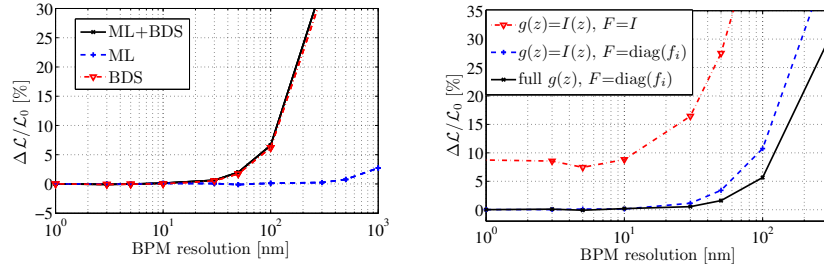


Figure 10: Relative peak luminosity loss as a function of the BPM resolution for the ML and BDS, separated and combined (left) and for different versions of the controller (right). Note that the luminosity loss is only due to BPM noise and that no other dynamic effects, e.g. ground motion, has been applied.

393 6.2. Beam energy jitter

The basic feedback algorithm encounters problems in the presence of beam energy jitter. This beam energy jitter is caused by small deviations of the initial beam energy and acceleration gradients from their nominal values. In the dispersive collimation section of the BDS, such energy variations result in beam offsets up to the millimetre range. The orbit controller reacts strongly on these large offsets. As a result the beam is mis-steered and the according luminosity loss is not tolerable. To counteract this effect, we deploy the fact that the beam offsets due to energy variations follow a specific pattern \mathbf{b}_D . By filtering this dispersion pattern from the BPM measurements with

$$\tilde{\mathbf{b}}_k = \mathbf{b}_k - \frac{\mathbf{b}_k^T \mathbf{b}_D}{\mathbf{b}_D^T \mathbf{b}_D} \mathbf{b}_D$$

394 the luminosity can be recovered almost fully. The use of this dispersion filtering
 395 is only necessary in the horizontal plane, since the coupling to the vertical plane
 396 can be neglected. The remaining luminosity loss due to the energy jitter coupling
 397 with the orbit controller has been observed to be 0.1%.

398 6.3. Other imperfections

399 Apart from the already mentioned imperfections, several other effects have
 400 been investigated. The scaling errors of BPMs and correctors are restricted

401 by the orbit controller action. For a relative luminosity loss of 0.5% one can
402 allow for a corrector scaling error up to 30%, while the BPM scaling error
403 tolerance for the same luminosity loss is as small as 1%. The tolerances for
404 static and jitter-like quadrupole strength errors are known to be very tight due
405 to the lattice design. The action of the orbit controller does not worsen these
406 tolerances in a notable way. Also the tolerances for the incoming beam jitter
407 at the entrance of the ML are hardly altered by the orbit controller operation.
408 Breakdown studies of BPMs, correctors and the stabilisation systems revealed
409 sensitivity to malfunctions of certain stabilisation systems and BPMs in the
410 BDS. An especially robust design for these systems is advisable and possibly also
411 redundancy has to be foreseen. For the positioning capability of the stabilisation
412 system, a tolerance of 0.25 nm has been identified. This tight tolerance has
413 been confirmed to be achievable by the CLIC stabilisation group. In case this
414 tolerance will turn out to be problematic for other accelerator designs, dipole
415 kickers can be used instead of the quadrupole movers. As mentioned before,
416 a fast and perfect actuator response $a(z) = 1$ has been assumed. For the
417 case of CLIC, this assumption is valid as shown by the stabilisation group [22].
418 However, if slower actuators are used no performance degradation is expected,
419 since the corresponding low pass behaviour of $a(z)$ can substitute parts of the
420 artificially introduced low pass $L(z)$.

421 **7. Conclusions**

422 In this paper, we have presented a ground motion optimised orbit controller
423 design method. This design method exploits a model of the beam oscillations
424 spectra in the BPMs, which have been derived from an accelerator model, ex-
425 isting noise and ground motion models. The orbit controller design method
426 consists of the three steps: decoupling of the inputs and outputs, time depen-
427 dent filter design and gain optimisation. This design method was applied to
428 create an orbit controller for CLIC. The orbit controller design method has
429 several advantages:

- 430 • Since the design is based on SVD decoupling, the overall controller system
431 is split up in SISO systems, which simplifies the design.
- 432 • The time dependent filter enables the user to incorporate expert knowl-
433 edge.
- 434 • The tedious task of optimising each decoupled loop (several thousand in
435 the case of the future linear accelerator) by hand is overtaken by an au-
436 tomated procedure. This eases the task of the designer.
- 437 • The controller design makes it possible to incorporate models of the ground
438 motion and the measurement noise. This closes a gap between the ground
439 motion research of the accelerator community and the orbit controller
440 design practice in a quantitative way.
- 441 • The controller performs better than a hand tuned controller in the case of
442 CLIC.

443 To evaluate the luminosity preservation ability of this controller, an inte-
444 grated simulation framework was set up. Full-scale simulations revealed that
445 the orbit controller, in combination with the other ground motion mitigation
446 methods, is capable of keeping the ground motion induced luminosity loss within
447 the allowed specifications. This is an essential contribution to resolve the feasi-
448 bility issue of ground motion for CLIC. Furthermore, robustness studies showed
449 that the controller is robust against imperfections, especially to measurement
450 noise. For CLIC the required BPM resolution could be loosened from 10 nm to
451 50 nm in the BDS. A sensitivity to variations of the beam energy was observed,
452 which could be resolved by filtering the resulting large dispersive orbits from
453 the BPM measurements.

454 Even though the presented orbit controller design method was developed for
455 the high demands of the future linear collider with respect to orbit stability, the
456 generic design procedure can be easily adapted to other linear machines. To
457 use the procedure for ring accelerators, the ground motion models would have
458 to be extended to the circular geometry, which could be an interesting subject

459 for future work. Alternatively, for existing accelerators the necessary power
460 spectral densities of the BPM readings can be obtained from measurements.

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