Freeze-in production of FIMP dark matter

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In collaboration with John March-Russell, Lawrence Hall and Karsten Jedamzik arXiv: 0911:112 [hep-ph] JHEP 1003:080,2010

Freeze-in is relevant for particles that are feebly coupled (Via renormalisable couplings) – λ Feebly Interacting Massive Particles (FIMPs) X

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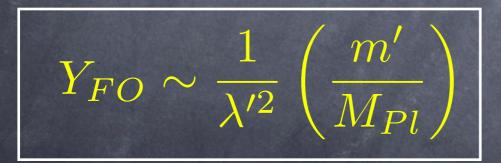
Although interactions are feeble they lead to some X production
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Increasing the interaction strength increases the yield opposite to Freeze-out...



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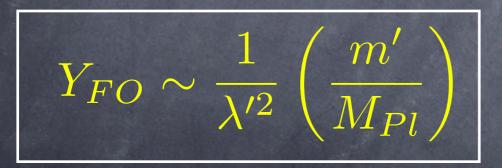
Using $\langle \sigma v \rangle \sim \lambda'^2/m'^2$



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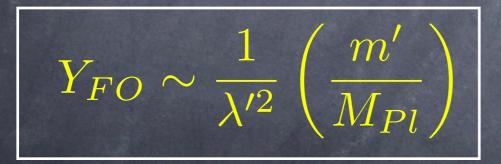


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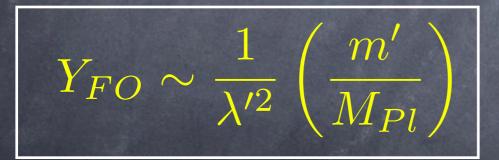
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Coupling strength λ



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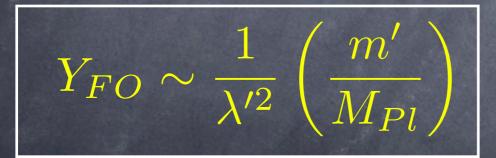


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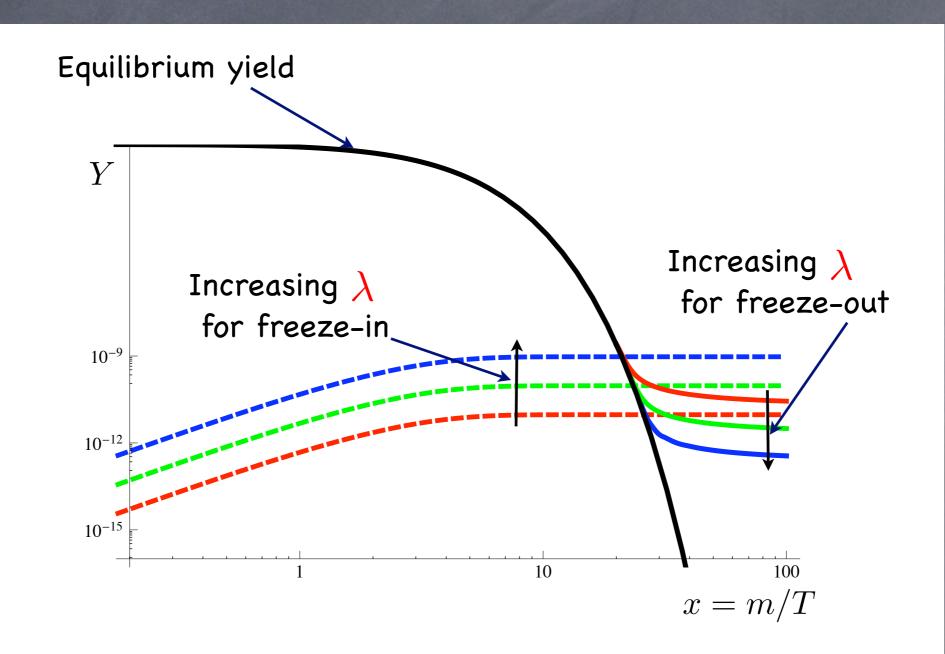
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 ψ_2

X

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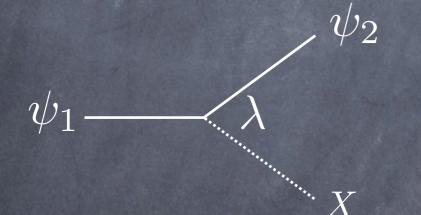
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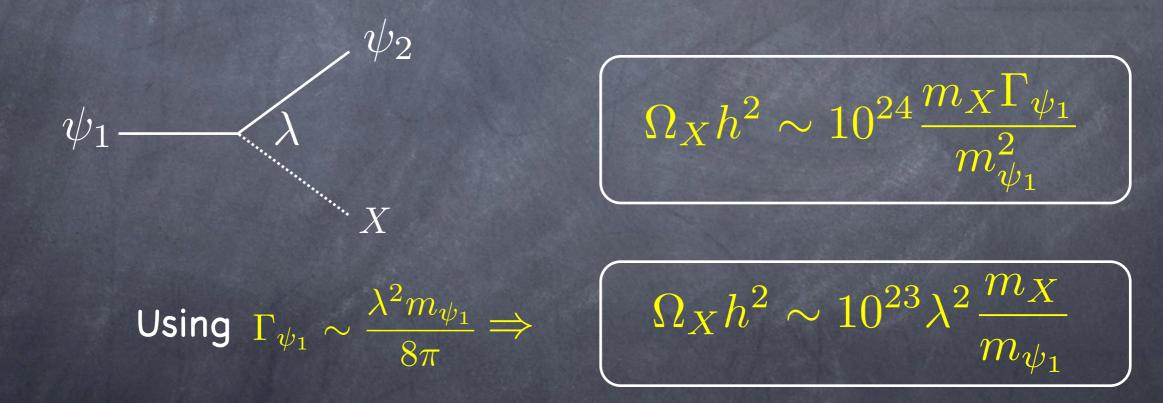
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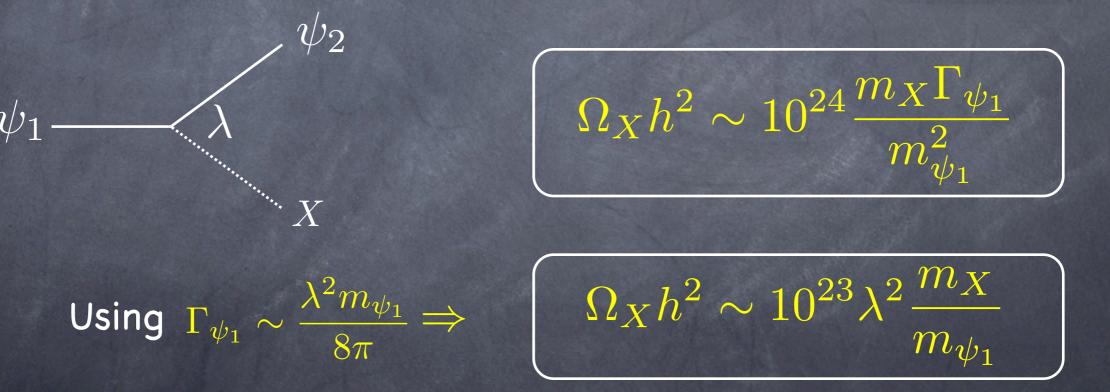


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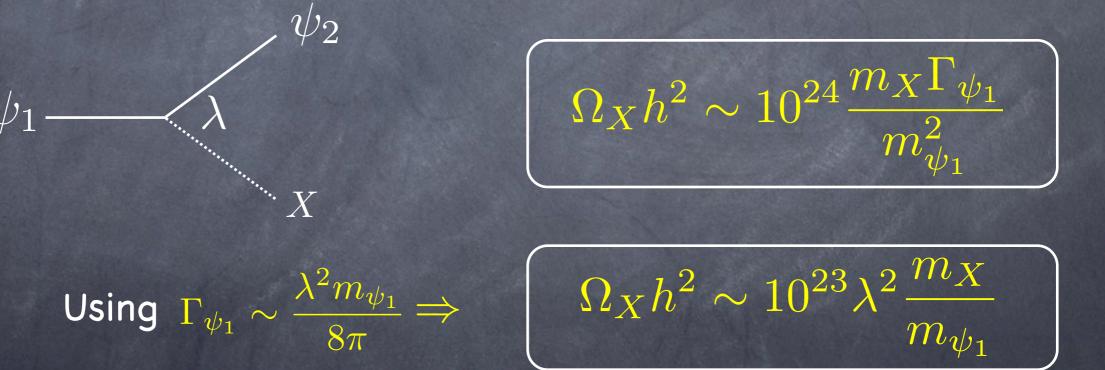
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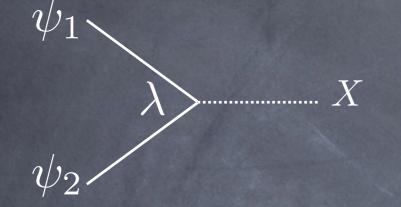
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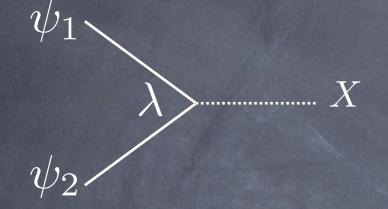
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Using $\Gamma_X \sim \frac{\lambda^2 m_X}{8\pi}$

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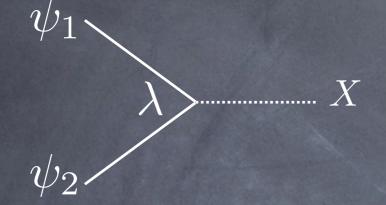
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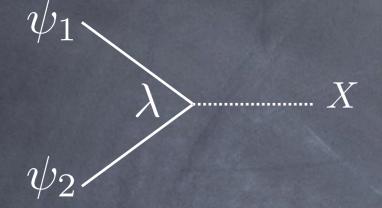
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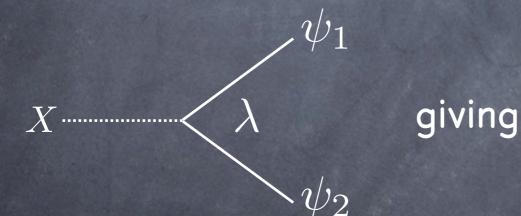
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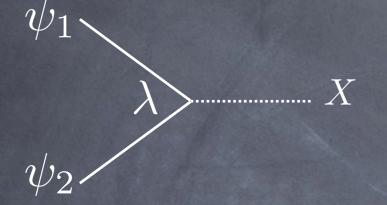
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$$\Omega_{\psi_1} h^2 = \frac{m_{\psi_1} \Omega_X h^2}{m_X} \sim 10^{23} \lambda^2 \frac{m_{\psi_1}}{m_X}$$

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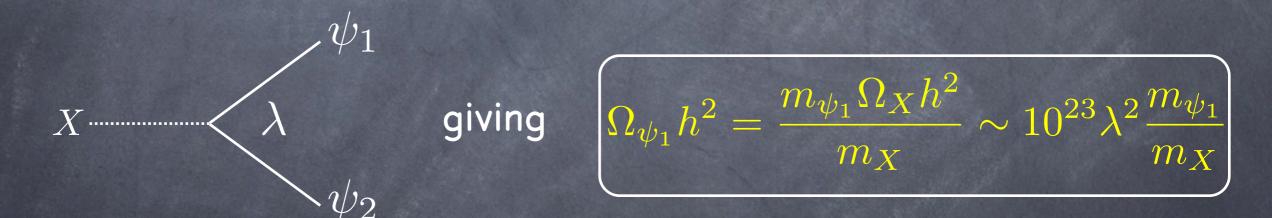
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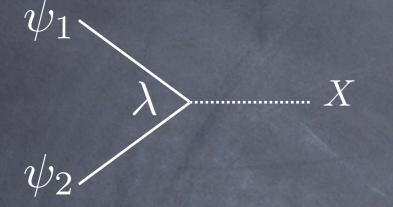
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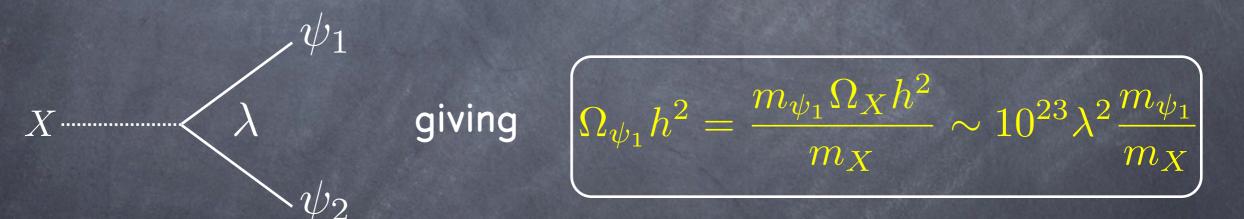
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Moduli and Modulinos associated with SUSY breaking

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• Dirac neutrino masses with SUSY – RH sneutrino FIMPs $\mathcal{L}_{\rm Dirac} = \lambda_{\nu} L H_u N \qquad \lambda_{\nu} \sim 10^{-12} \qquad \frac{\text{See Moroi et al}}{\text{for related}}$

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• FIMPs from kinetic mixing: hidden sector particles coupling to the MSSM via mixing of $U(1)_{Y}$ and hidden U(1) tiny mixing tiny coupling

• Others...

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 Enhanced indirect and direct detection: Relic abundance and DM annihilation cross section no longer related.
 Freeze-in dominantly produces DM abundance annihilation cross section must be large – freeze-out abundance is small

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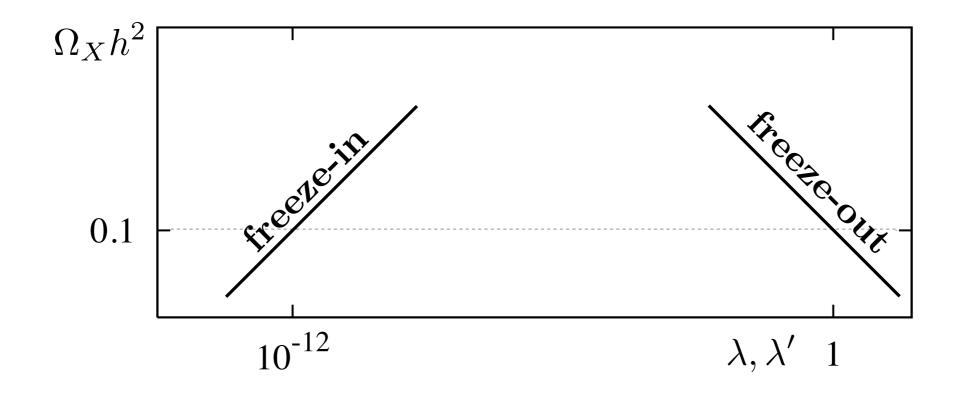
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Much more to come...

Freeze-in vs Freeze-out

• For a TeV scale mass particle we have the following picture.



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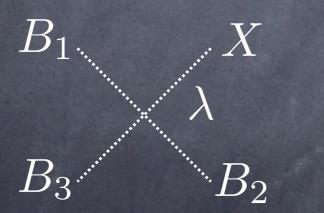
$$B_1$$
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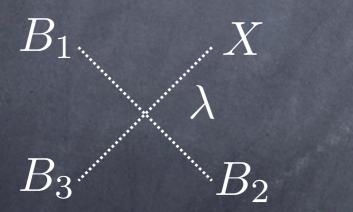


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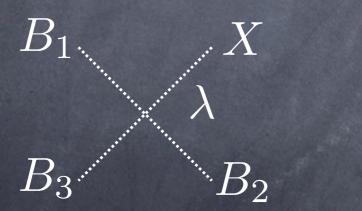
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For correct DM abundance $\Rightarrow \lambda \sim 10^{-11}$

NOTE: Abundance in this case is independent of the FIMP mass